

Dark Photon Search at PHENIX

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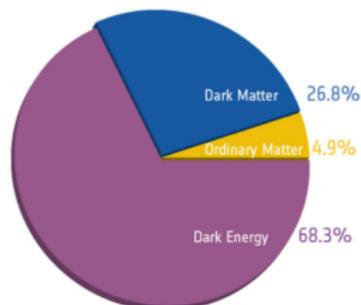


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University

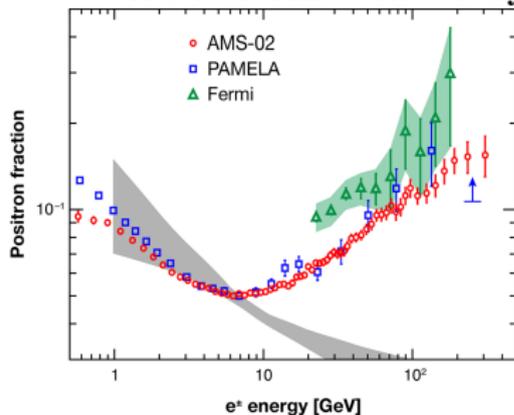


Our universe is composed of 27% dark matter and 68% dark energy.

- Considerable interest in exploring the nature of dark matter
 - Long list of candidates; e.g. WIMPS, Axions, extra dimensions and many more.
- A new possibility - dark sector.
 - 1 Extension of the Standard model.
 - 2 Additional new gauge field –**Dark photon**



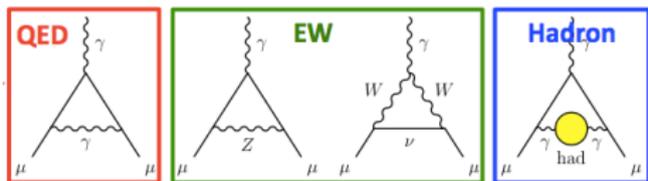
Positron excess in cosmic rays



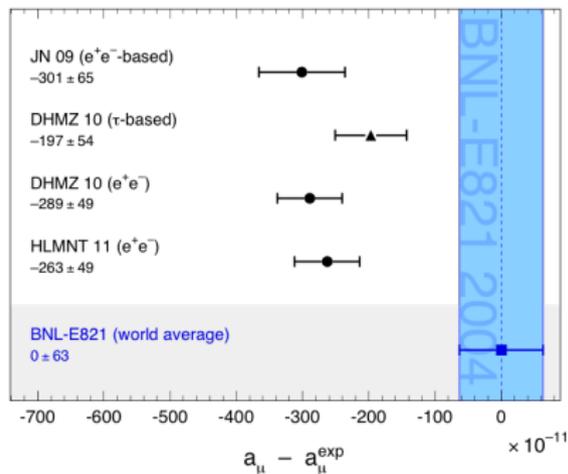
Experimental results suggesting possible BSM phenomenon

- Excess in positron fraction compared to SM prediction; seen by PAMELA and AMS-02.
- μ anomalous magnetic moment
 - Very precise measurement as well as theoretical calculations
 - **Strong motivation for our measurement.**

μ – magnetic moment anomaly



(M. Davier et al. 2011)



Standard Model predicts very precisely the value of "a" defined as:

- $a_i = \frac{(g-2)_i}{2}$, $i = e, \mu$.
- $a^{SM} = a^{QED} + a^{EW} + a^{Had}$.
- For the electron: $a_e^{SM} \approx a^{QED}$
 → serves as a QED test
 $a_e^{EXP} - a_e^{SM} = -(1.06 \pm 0.82) \times 10^{-12}$
- Muon: a_μ^{EW}, a_μ^{Had} are not negligible
 → a SM test
- $a_\mu^{SM} = (116591802 \pm 49) \times 10^{-11}$
- Precise measurement by BNL-E821 experiment
 $a_\mu^{EXP} - a_\mu^{SM} = (287 \pm 80) \times 10^{-11}$
- About $2.4\sigma - 3.6\sigma$ discrepancy.

Possibility to explain a_μ^{EXP} by dark photon

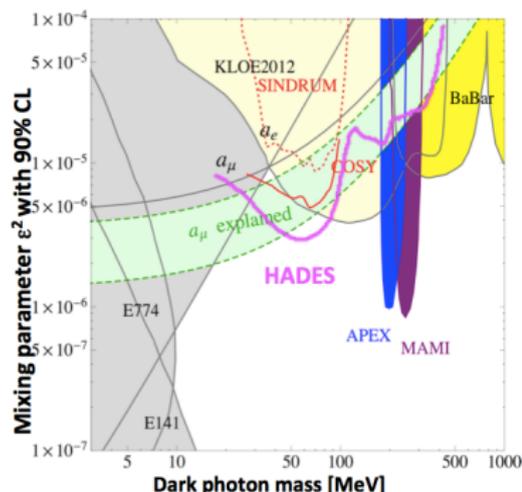
The Dark photon

- Simplest hidden sector model introduces one extra $U(1)$ gauge symmetry and a corresponding gauge boson: the dark photon.
 - leaves the SM particles unchanged.
 - Associated gauge boson can communicate with the SM through a small mixing on the kinetic term of the QED Lagrangian.

$$\mathcal{L}_{mix} = -\frac{\varepsilon}{2} F_{\mu\nu}^{QED} F_{dark}^{\mu\nu}$$



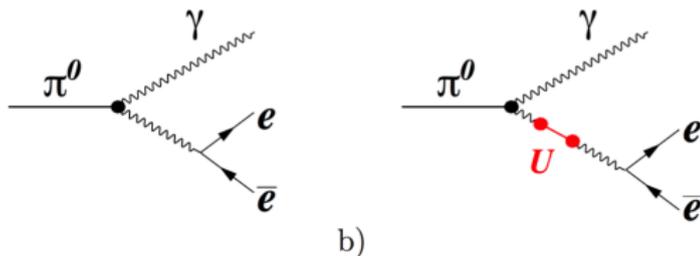
- The gauge boson U (also A' or Z'_d) is referred to as a dark photon since it can mix with the real photon in all processes.
- Strength of the mixing parameter ε as suggested by phenomenological arguments must be of the order of $10^{-4} - 10^{-2}$ and the boson mass M_U below 2 GeV.



Mapping in the parameter space shown as a function of the dark photon mass (about 1 year ago). Theory curves from Hye-Sung Lee & Bill Marciano.

Batell, Pospelov and Ritz, PRD80 (2009) 095024

Ref: PLB 726, 187 (2013)

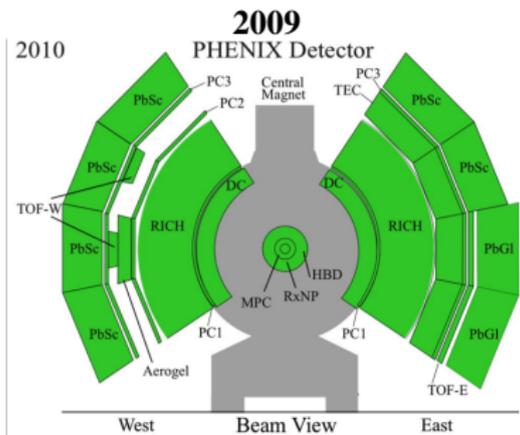
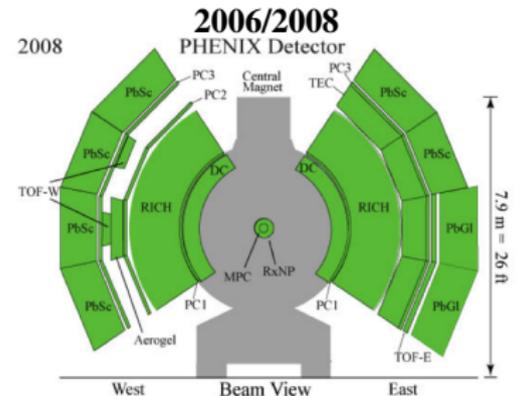


Measurement of $\pi^0/\eta \rightarrow \gamma U \rightarrow \gamma e^+ e^-$ in π^0/η Dalitz decays

- **Assumption:** Dark photon exclusively decays into an $e^+ e^-$ pair.
 - its natural width is very narrow !
 - Expected peak width = detector mass resolution.
 - Same approach was used in COSY-WASA & HADES

Important requirements for this measurement

- 1 Large statistics of $e^+ e^-$ from π^0/η Dalitz decays.
- 2 Good mass resolution of $e^+ e^-$ spectrum.



PHENIX Central arms Acceptance:

$$-0.35 < \eta < 0.35, 2 \times 90^\circ \text{ in } \varphi$$

- Data sets used in the analysis are from the years 2006, 2008 and 2009.
 - Hadron Blind Detector installed in 2009 (radiation length increase by 2.4%).
- Vertex determination: **BBC**
- Tracking: **DC/PC1**
 - $\delta p/p = 1\% \oplus 1.1\% \times p$ [GeV/c].

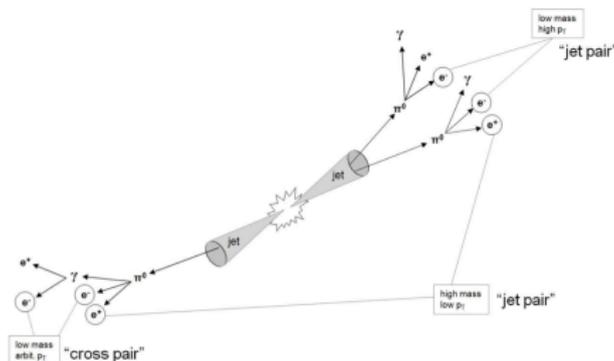
Electron identification based on:

- **RICH** (Ring Imaging Čerenkov detector) (e/π rejection >1000)
- **EMCal** (Electromagnetic Calorimeter) (E-p matching, e/π rejection ~ 10)

Understanding the background

The unlike sign mass spectrum consists of two types of backgrounds:

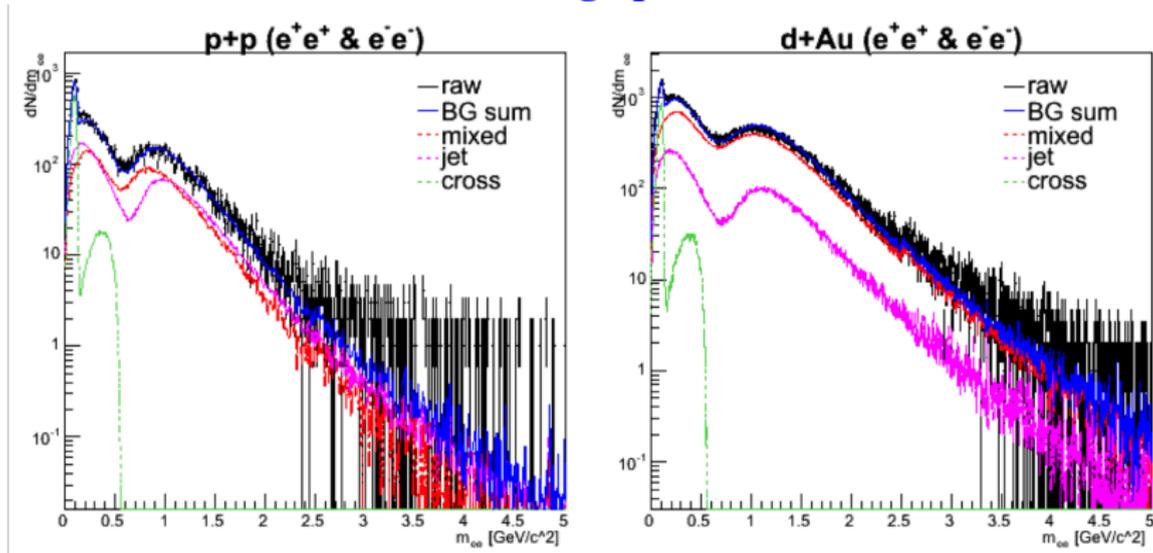
- **Uncorrelated Combinatorial Background:** arises from all the combinations where the origin of two electrons is totally uncorrelated.
- **Correlated Background:**
 - Cross Pairs: If there are two e^+e^- pairs in the final state of a meson, e.g. π^0 double Dalitz decay ($\pi^0 \rightarrow e_1^+ e_1^- e_2^+ e_2^-$), or a Dalitz decay ($\pi^0 \rightarrow \gamma e_1^+ e_1^-$) where γ converts to $e_2^+ e_2^-$.
 - Jet Pairs: Hadrons either from the same jet or in back-to-back jets, that decay into electron pairs.



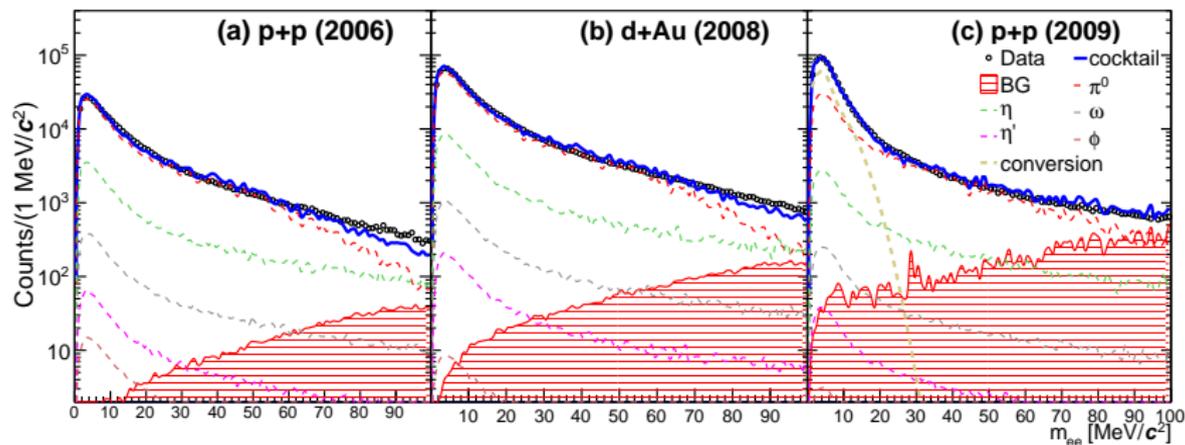
Two methods are used to estimate the background:

- Like sign event technique that takes into account both combinatorial and correlated background (2009 analysis).
- Subtract the combinatorial using Mixed Event Technique and then subtract the correlated background estimated from simulations (2006 and 2008 analysis).

Like-sign pairs



- Like-sign pairs are used to evaluate the contributions from different components.
- Sum of all the components describe the distribution well.

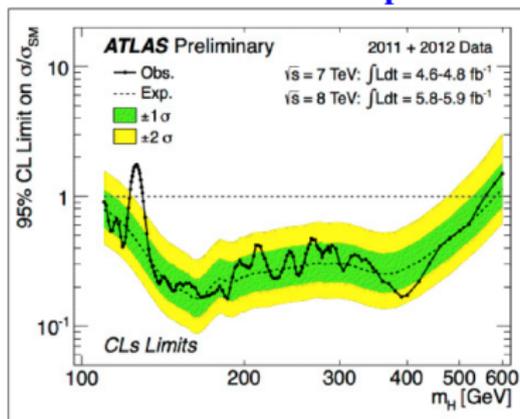
e^+e^- mass spectra for the different analyzed data sets

- Measured e^+e^- spectra is well described by an expected “cocktail” of hadronic decays + the background.
- No significant dark photon peak visible.
- Use a statistical analysis based approach to extract the dark photon signal, *if any*.

CLs approach

- Widely accepted method to set confidence levels for hypothetical particles.
 - Famous “Brazil band plot “ for Higgs search at LHC.
- Relative likelihoods of how well the data is described by
 - Only background (Dalitz continuum).
 - Signal (Dark photon) + Background.
- Requires an **expected shape** of the dark photon peak and background.

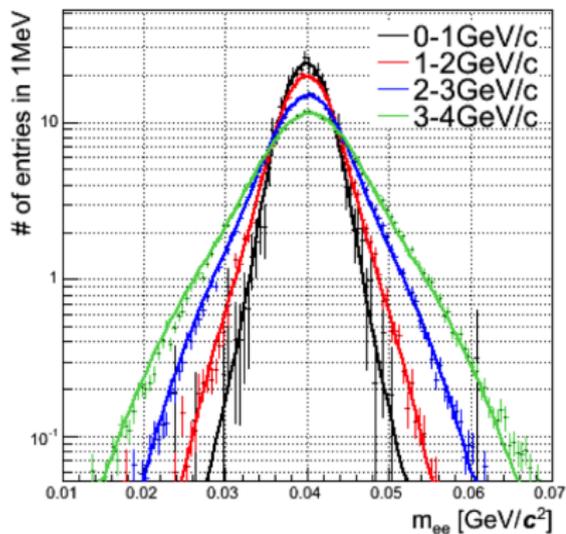
ATLAS Brazil band plot



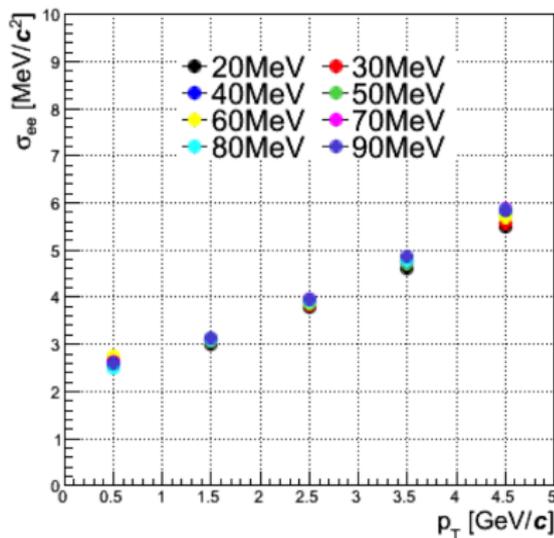
Extracting dark photon mass resolution in PHENIX

- The dark photon couples weakly to QED photon \Rightarrow narrow natural width.
 - \Rightarrow expected line shape of the dark photon is set by the mass resolution σ of the detector.
- Mass resolution of the PHENIX was calculated by the GEANT based simulation tuned to match the real data.
 - The expected dark photon peak width is ~ 3 MeV (for inclusive p_T).

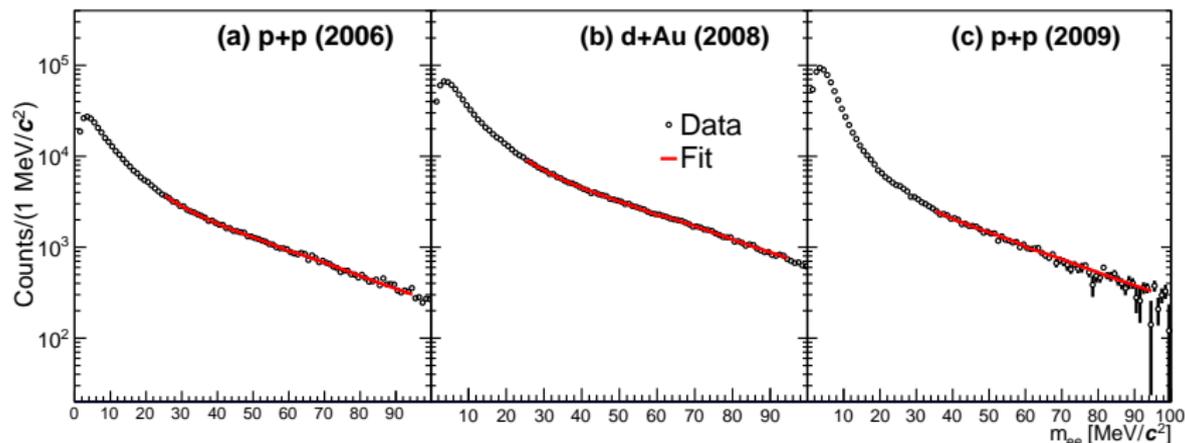
Dark photon mass = 40MeV



Mass resolution as a function of p_T



Background shape for the peak search

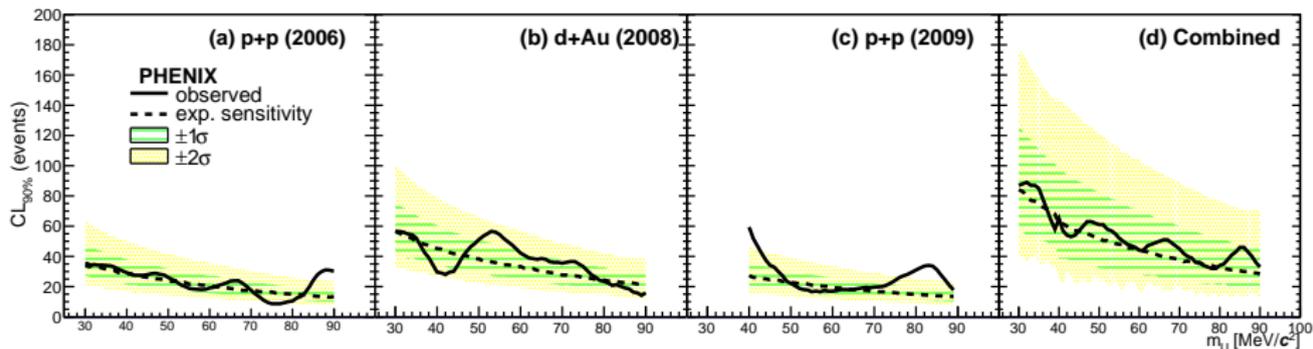


- Common fit function for the three data sets with different scale factors.

$$f_{fit}(m_{ee}) = \frac{1}{m_{ee}} \times \left[\left(1 - \frac{m_{ee}^2}{m_{\pi^0}^2} \right)^3 + r_{\eta/\pi^0} \cdot \left(1 - \frac{m_{ee}^2}{m_{\eta}^2} \right)^3 \right] \times f_{chebychev}^i(m_{ee})$$

- **Chebychev polynomial** allows for any slight deviations from **Kroll-Wada** shape due to detector effects.
- Two separate fit ranges to avoid having a local bad χ^2 .
 - Smoothly connects at the break-point.

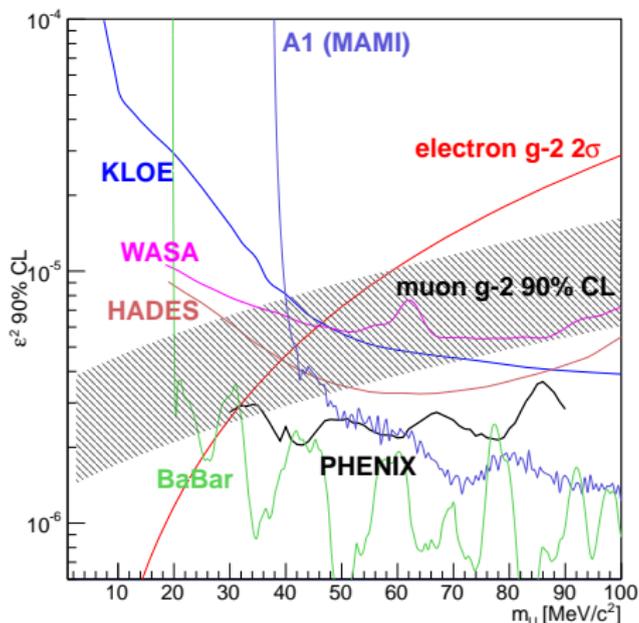
Possible dark photon events



- Possible events with dark photon candidates with 90% CL using CLs approach.
- Shown are experimental sensitivity, and its $\pm 1,2 \sigma$ uncertainties.
- Observed limits within the 2σ fluctuation of our sensitivity.
- \Rightarrow No dark photon signal is observed.

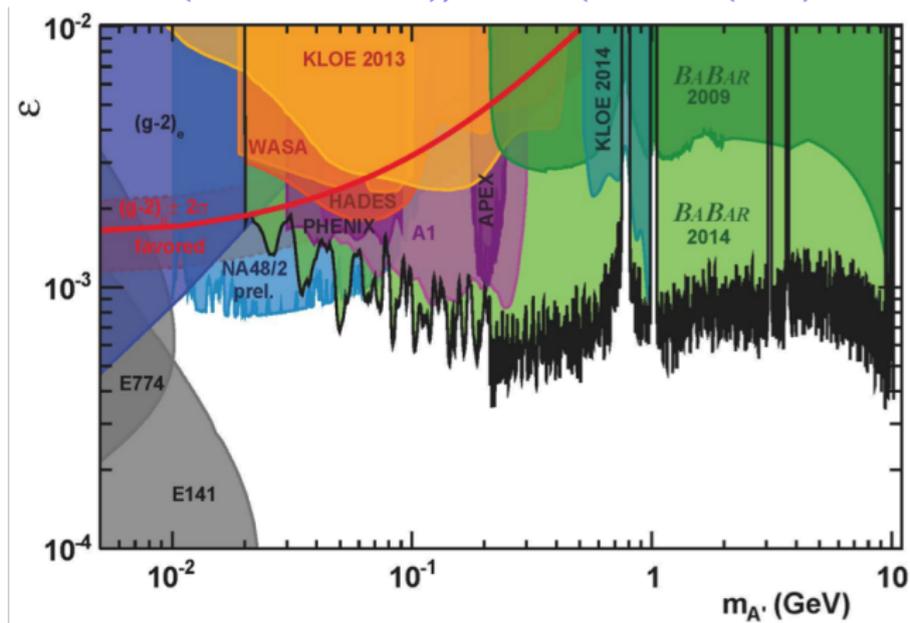
Results- dark photon mixing

$$\epsilon^2 = \frac{2\alpha_{EM}}{3\pi} \frac{\sigma}{m_U} \sqrt{2\pi} R(m_U); R(m_U) = \frac{N_U(m_U)}{N_{Dalitz}(m_U)}$$



- PHENIX results cover the mass range $30 < m_U < 90 \text{ MeV}/c^2$
 - Set a stricter limit than those of WASA, HADES or KLOE in the mass region between $30 < m_U < 50 \text{ MeV}/c^2$.
 - Complement the A1(MAMI) results by covering their less sensitive area.
 - Excludes the values of the coupling favored by the $(g-2)_\mu$ region for $m_U > 32 \text{ MeV}/c^2$.

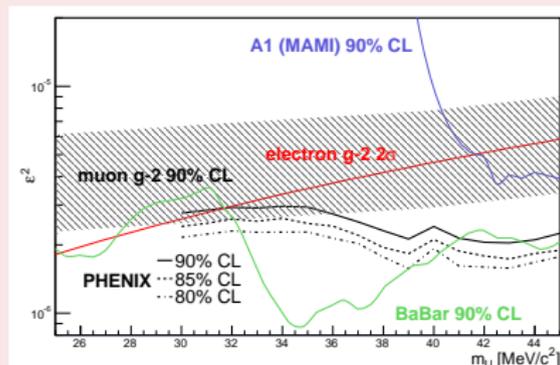
NA48 Results(arXiv:1412.8053); BABAR(PRL 113 (2014) 201801)



- Together with the BABAR and NA48/2 results, full “g-2 band” is excluded at the 90% CL.

Summary

- PHENIX carried out a search for the dark photon in π^0/η Dalitz decays.
- Results set limits for the coupling of a dark photon to the QED photon over the mass range $30 < m_U < 90 \text{ MeV}/c^2$.
- Combining with the other world data, dark photon is ruled out with 90% CL as the explanation for observed $(g - 2)_\mu$ anomaly.
- Accepted to be published as Phys.Rev.C. Rapid ([arXiv:1409.08501](https://arxiv.org/abs/1409.08501)).



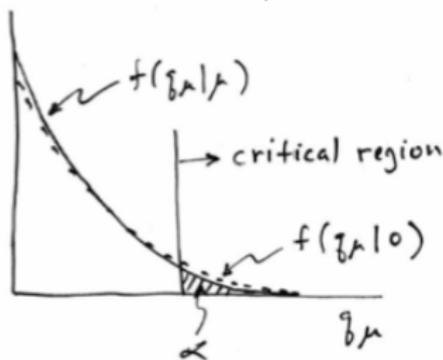
Outlook

- Analysis of the high statistics dataset taken in 2014 using Vertex detector will provide more stringent limits.

Back-ups

Why CLs method is better?

- ❖ Test of the hypothesis:
 - q_μ : likelihood ratio with a signal level, μ
 - $f(q_\mu | \mu)$: likelihood distribution for μ
 - ✓ $f(q_\mu | 0)$ means the distribution for only background.
 - P-value for observed q_μ :
$$p_\mu = \int_{q_{\mu, \text{obs}}}^{\infty} f(q_\mu | \mu) dq_\mu$$
 - Normally 90%CL limit for μ is the highest value for $p_\mu > 0.10$.
- ❖ If a given μ is very small:
 - $f(q_\mu | \mu, 0)$ are almost consistent.
 - Probability to reject μ for $f(q_\mu | 0)$ is slightly larger than that for $f(q_\mu | \mu)$ = 90%CL limit for normal case.
 - This leads to exclude hypotheses to which one has no sensitivity.



“Spurious exclusion”

CLs method

- ❖ The CLs method is very popular in HEP to calculate a upper limit with avoiding the spurious exclusion.

1. Calculate the likelihood ratio of Poisson probabilities with two different hypotheses, Q:

a. signal+background

$$Q = \frac{P_{\text{poiss}}(\text{data}|\text{signal} + \text{background})}{P_{\text{poiss}}(\text{data}|\text{background})},$$

b. only background

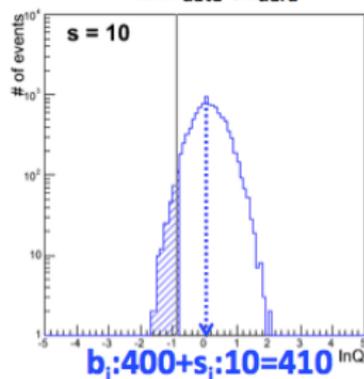
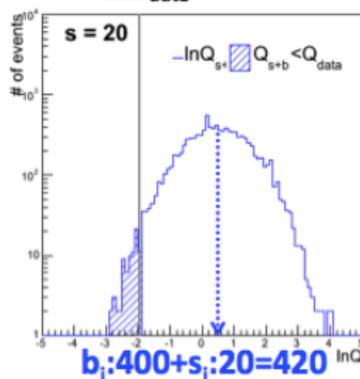
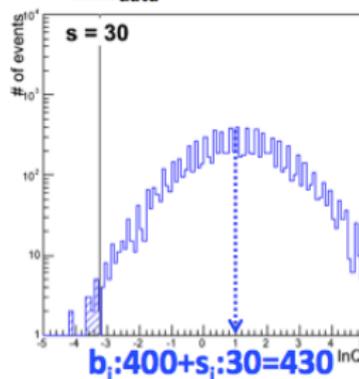
2. Log-likelihood ratio, lnQ:

$$\ln Q = \sum_{i=1}^{n_{bins}} \{s_i - n_i \ln(1 + s_i/b_i)\},$$

- s_i = estimated dark photon signal in bin i
 - ✓ Scan a signal amplitude of expected dark photon peak
- b_i = Dalitz continuum background in bin i
 - ✓ From the fit function of Dalitz continuum distribution
- n_i = the number of entries in bin i
 - ✓ lnQ_{data} : n_i=n_{data}
 - ✓ lnQ_{s+b} : n_i=Poisson(s_i+b_i)
 - ✓ lnQ_b : n_i=Poisson(b_i)

CLs method (cont.)

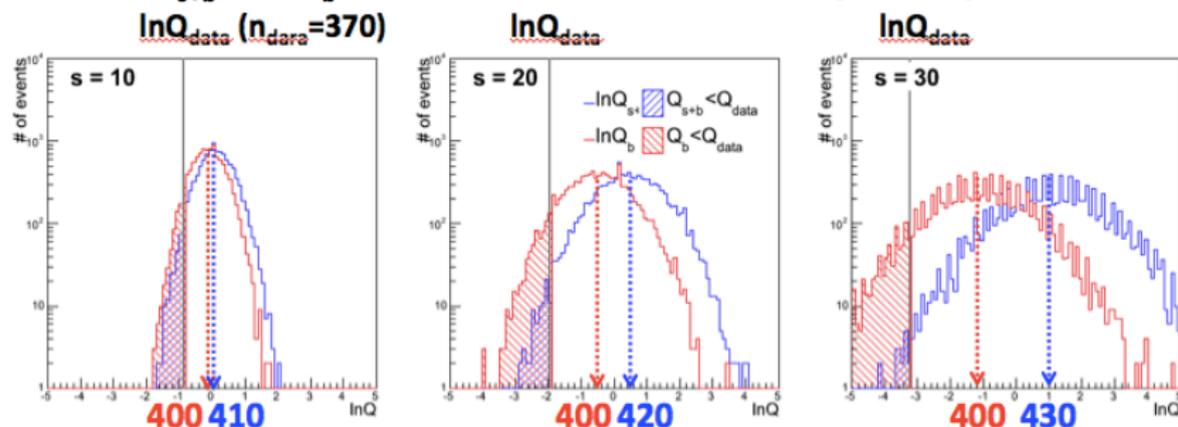
3. $\ln Q_{s+b}$ & $\ln Q_b$ are calculated with 10000 sample, then calculate:
- $CL_{s+b} = P(Q_{data} > Q_{s+b})$
 - $CL_b = 1 - P(Q_b > Q_{data}) (= P(Q_{data} > Q_b))$
- Example: $n_{data}=370$, $b_i=400$

 $\ln Q_{s+b}$ distributions with different signal amplitudes $\ln Q_{data}$ ($n_{data}=370$) $\ln Q_{data}$  $\ln Q_{data}$ 

→ $P(Q_{data} > Q_{s+b})$ is interpreted as the probability that n_{data} can be explained by a statistical fluctuation with a mean value of $s_i + b_i$.

CLs method (cont.)

3. $\ln Q_{s+b}$ & $\ln Q_b$ are calculated with 10000 sample, then calculate:
- $CL_{s+b} = P(Q_{data} > Q_{s+b})$
 - $CL_b = 1 - P(Q_b > Q_{data}) (= P(Q_{data} > Q_b))$
- Example: $n_{data} = 370$, $b_i = 400$

 $\ln Q_{s+b}$ & $\ln Q_b$ distributions with different signal amplitudes

→ The hatched areas indicate the fractions that the data can be explained by the $s+b$ and b hypotheses out of 10^4 samples.

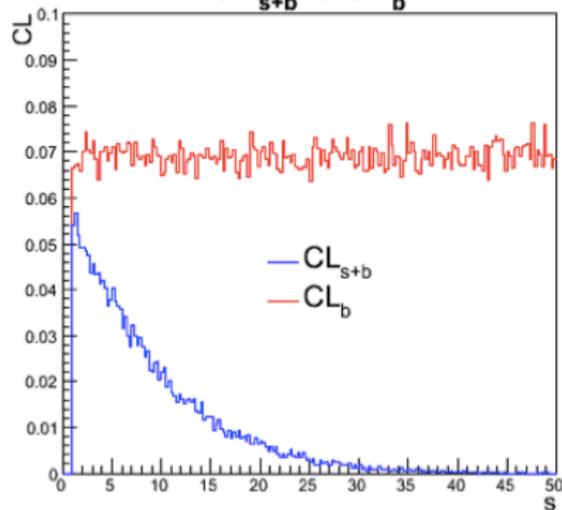
CLs method (cont.)

4. CL_s is calculated as:

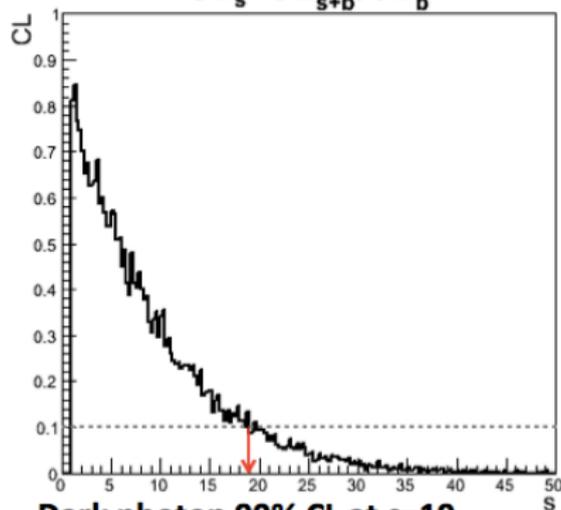
– $CL_s = CL_{s+b}/CL_b$

✓ CL_s is explained as CL_{s+b} re-normalized by CL_b .

CL_{s+b} & CL_b



$CL_s = CL_{s+b}/CL_b$



Dark photon 90% CL at s=19

References of CLs method

- ❖ More detailed descriptions for the CLs method are:
 1. A.L. Read, J. Phys. G 28, 2693, 2002
 2. V. Buescher, et al., D0 Note 4629, 2004
 3. G. Cowan, Lecture note for Day 3 of Statistical Data Analysis for High Energy Physics 2011
http://www.pp.rhul.ac.uk/~cowan/stat_freiburg.html