

E_T and other Event-by-Event Distributions

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- (Charged) Multiplicity n or dn/dy or $dn/d\eta$
- Transverse Energy E_T
- Average p_T per event, in contrast to
 - $\langle p_T \rangle$ averaged over all events

Multiplicity in collision of nuclei

- The study of relativistic collisions involving nuclei has a long tradition, dating from studies in the 1930's of cosmic ray interactions in emulsions and in cloud chambers.
- One of the burning issues in the early 1950's was whether more than one meson could be produced in a single nucleon-nucleon collision ("multiple production") or whether the multiple meson production observed in nucleon-nucleus interactions was the result of several successive n-n collisions, with each collision producing only a single meson ("plural production").
- The issue was decided when multiple meson production was first observed in 1954 in collisions between neutrons of energy up to 2.2 GeV produced at the Brookhaven Cosmotron and protons in a hydrogen filled cloud chamber. Phys. Rev. **95**, 1026 (1954)

Single particle Inclusive Reactions

- A single particle “inclusive” reaction involves the measurement of just one particle coming out of a reaction,

$$a + b \rightarrow c + \text{anything} \quad .$$

The terminology comes from the fact that all final states with the particle c are summed over, or **included**.

- A “semi-inclusive” reaction refers to the measurement of all events of a given topology or class, e.g.

$$a + b \rightarrow n_1 \text{ particles of class 1} + \text{anything} \quad .$$

Kinematics

A particle of mass m has energy E , longitudinal momentum p_L , transverse momentum p_T , rapidity y :

$$y = \ln \left(\frac{E + p_L}{m_T} \right) \quad (1)$$

$$\cosh y = E/m_T \quad \sinh y = p_L/m_T \quad (2)$$

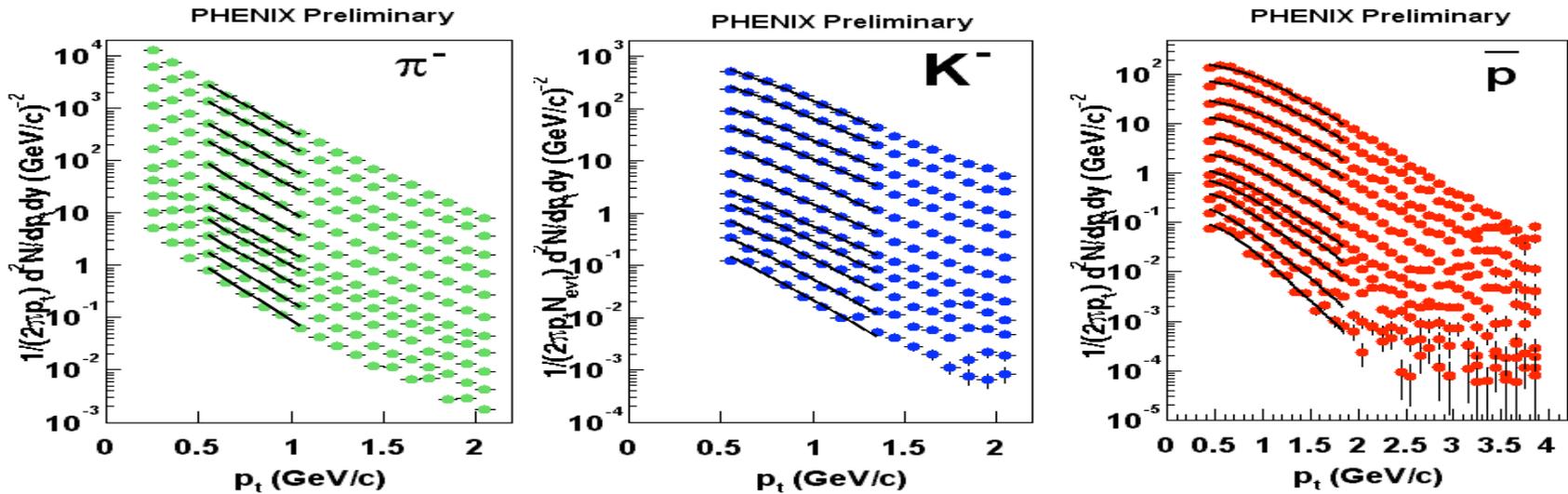
where

$$m_T = \sqrt{m^2 + p_T^2} \quad \text{and} \quad E = \sqrt{p_L^2 + m_T^2} \quad (3)$$

In the limit when ($m \ll E$) the rapidity reduces to the pseudorapidity (η)

$$\eta = -\ln \tan \theta/2 \quad (4)$$

Semi-Inclusive p_T spectra Au+Au $s_{NN}=200$ GeV



- Measurements are presented in terms of the (Lorentz) Invariant single particle inclusive differential cross section (or Yield per event in the class if semi-inclusive):

$$\frac{E d^3\sigma}{dp^3} = \frac{d^3\sigma}{p_T dp_T dy d\phi} = \frac{1}{2\pi} \mathbf{f}(p_T, y) \quad . \quad (5)$$

- A uniform azimuthal distribution is usually assumed, so that $\int d\phi = 2\pi$:

$$\frac{d^2\sigma}{p_T dp_T dy} = \mathbf{f}(p_T, y) \quad . \quad (6)$$

- $\langle p_T \rangle$ averaged over all events is obtained the usual way:

$$\langle p_T(y) \rangle = \frac{\int p_T dp_T p_T \mathbf{f}(p_T, y)}{\int dp_T p_T \mathbf{f}(p_T, y)} \quad (7)$$

Integrals of the cross section= $\langle n \rangle$

- Integrals of the single particle inclusive cross section (Eq. 5) are not equal to σ_I the interaction cross section, but rather equal to the mean multiplicity times the interaction cross section : $\langle n \rangle \times \sigma_I$.

- Integration over p_T gives the average multiplicity density in rapidity, dn/dy :

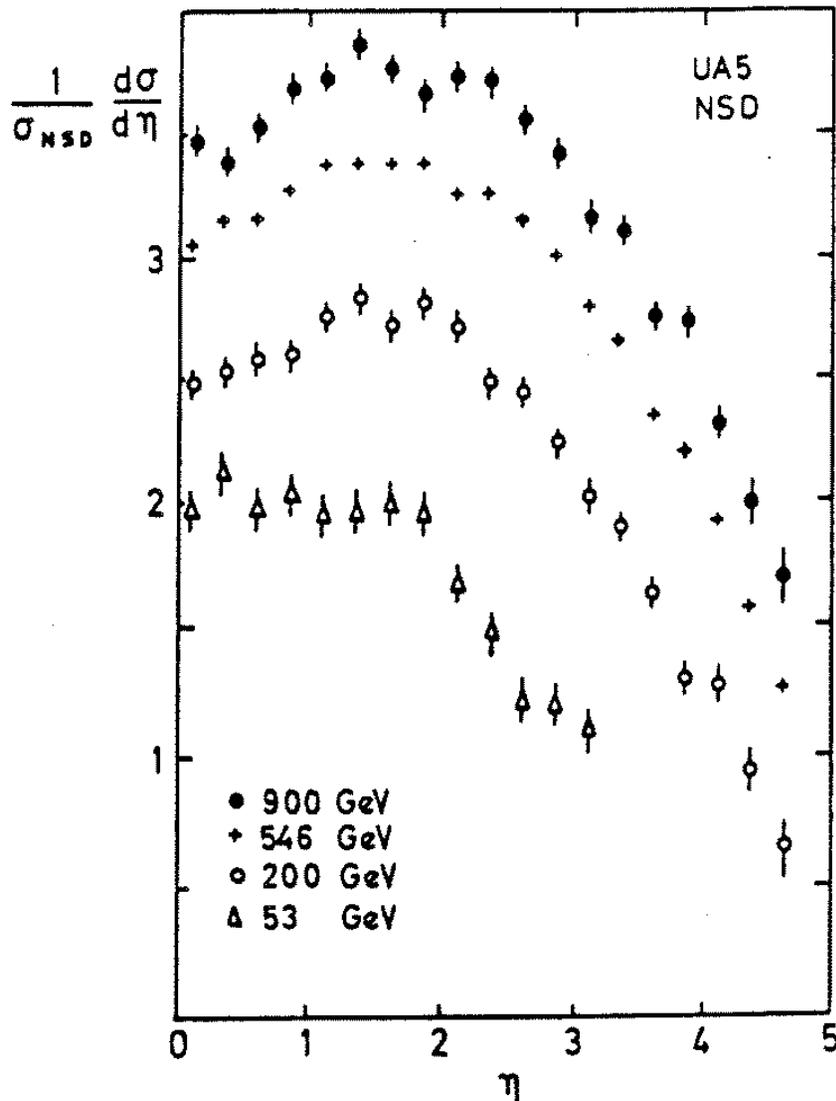
$$\frac{dn}{dy} = \frac{1}{\sigma_I} \frac{d\sigma}{dy} = \frac{1}{\sigma_I} \int dp_T p_T \mathbf{f}(p_T, y) = \rho(y) \quad . \quad (8)$$

- Integration over the kinematically possible rapidity gives the overall mean multiplicity per interaction, $\langle n \rangle$:

$$\langle n \rangle = \frac{1}{\sigma_I} \int dy dp_T p_T \mathbf{f}(p_T, y) = \int dy \frac{dn}{dy} = \int dy \rho(y) \quad (9)$$

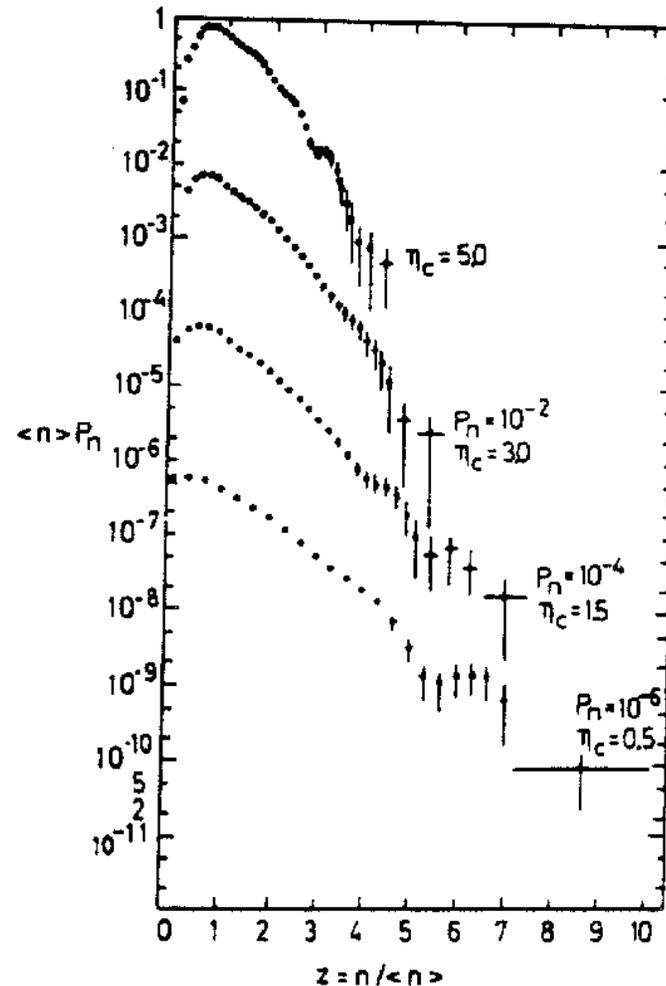
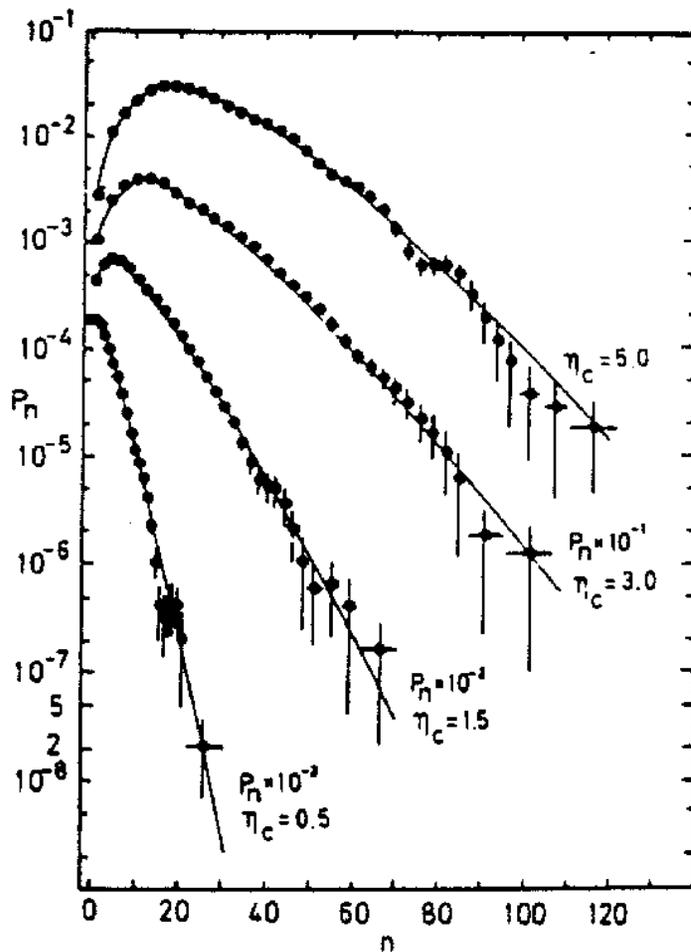
- Terminology is dn/dy for identified particles (known m), $dn/d\eta$ for non-identified particles (m unknown, assumed massless).

UA5 (1985) $dn/d\eta$ in $p\bar{p}$ -p



- Normalized to Non Single-Diffractive Cross Section
- This is an average quantity, averaged over all events.
- The event-by-event multiplicity distribution in regions of η is much more interesting: Instead of averaging over all events, plot a frequency distribution of the multiplicity n in an interval of pseudorapidity $|\eta| < \eta_c$
- Study of fluctuations in small intervals near mid-rapidity might allow observation of the "real" fluctuations freed of constraints like energy, momentum and charge conservation which need not be locally conserved.

UA5 Multiplicity Distribution $|\eta| < \eta_c$ shows huge variation event-by-event



UA5 PLB **160**, 193,199 (1985); **167**, 476 (1986)

Distributions are Negative Binomial, NOT POISSON: implies correlations

E_T distributions

- E_T is an event-by-event variable defined as:

$$E_T = \sum_i E_i \sin \theta_i \quad \text{and} \quad dE_T(\eta)/d\eta = \sin \theta(\eta) dE(\eta)/d\eta \quad , \quad (10)$$

- The sum is taken over all particles emitted on an event into a fixed but large solid angle, (which is different in every experiment).

- Measured in Hadronic and Electromagnetic Calorimeters and even as sum of charged particles $\sum_i |p_{T_i}|$

- Introduced by High Energy Physicists as an “improved” method to detect and study the jets from “hard-scattering” compared to high p_T single particle spectra by which hard scattering was discovered in p-p collisions and used as a hard-probe in Au+Au collisions at RHIC. **Didn't work as expected: E_T distributions are dominated by soft particles near $\langle p_T \rangle$.**

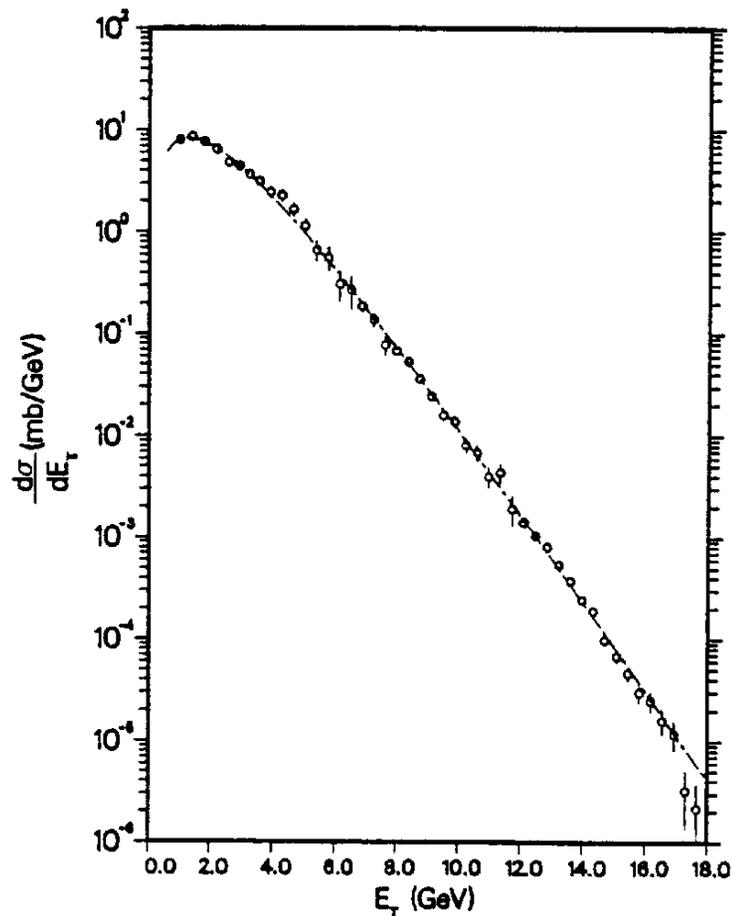
- The importance of E_T distributions in relativistic heavy ion (RHI) collisions is that they are largely dominated by the *nuclear geometry* of the reaction and so provide a measure of the overall character or *centrality* of individual RHI interactions. **Bjorken Energy density**

- Normalization Problem: Measurement is not accurate but is very precise, good for A dependences in a single setup; not as good (larger systematic errors) for comparison of different experiments.

♥ correct from E_T in measured aperture to E_T in $\Delta\Phi = 2\pi$ $\Delta\eta = 1$ with or without HAD/EM calorimeter response correction.

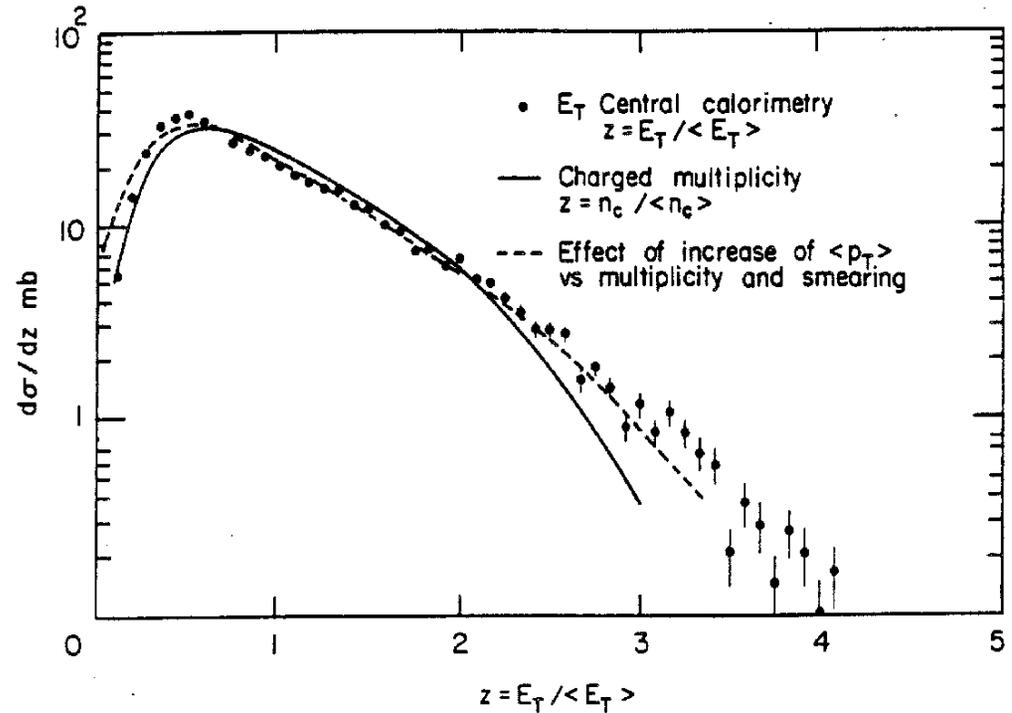
♥ Measure $\langle E_T \rangle_{p-p}$ in same aperture and correct scale to $E_T / \langle E_T \rangle_{p-p}$, i.e. E_T in units of $\langle E_T \rangle$ per participant pair.

NA5 (CERN) (1980) First E_T dist. pp



NA5 300 GeV PLB 112, 173 (1980)
 $2 \square$, $-0.88 < y < 0.67$ **NO JETS!**
 $s = 23.7$ GeV
 Fit (by me) is \square dist $p = 2.39 \pm 0.06$

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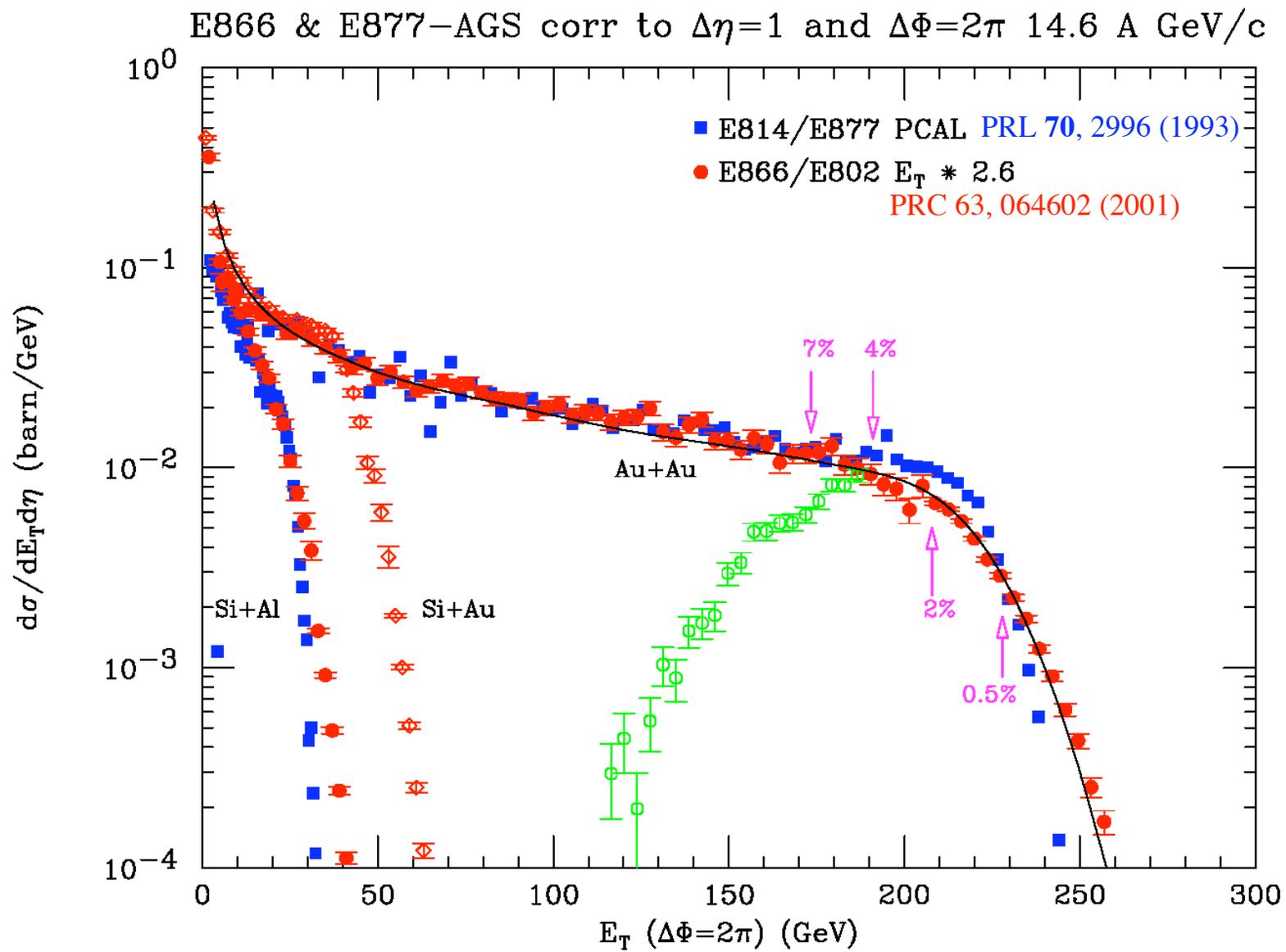


UA1 (1982) (C.Rubbia) $s = 540$ GeV.
 No Jets because E_T is like multiplicity
 (n), composed of many soft particles near
 $\langle p_T \rangle$! CERN-EP-82/122.

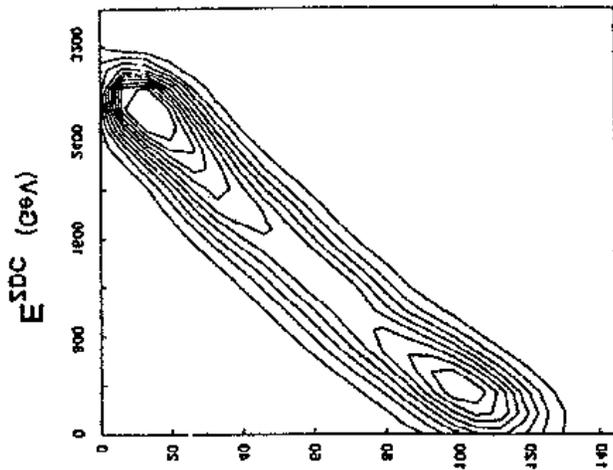
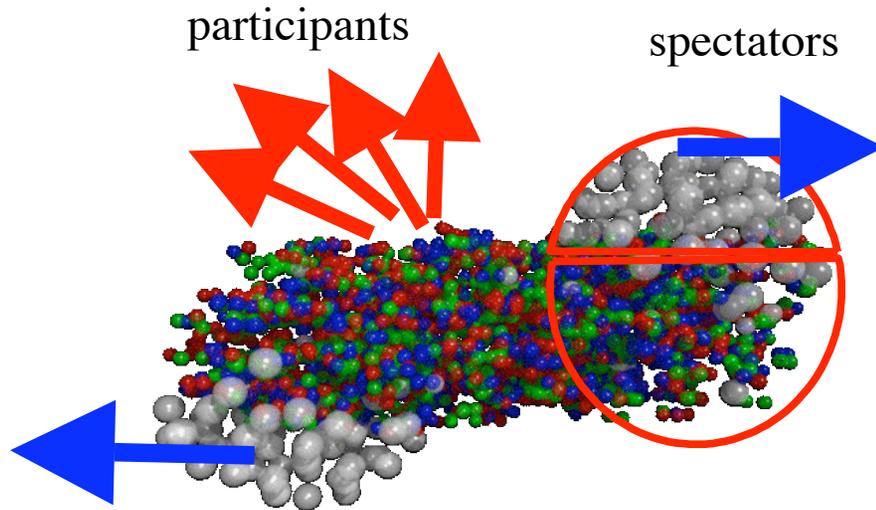
OOPS UA2 discovers jets 5 orders of
 magnitude down E_T distribution!

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Typical E_T distributions in RHI collisions

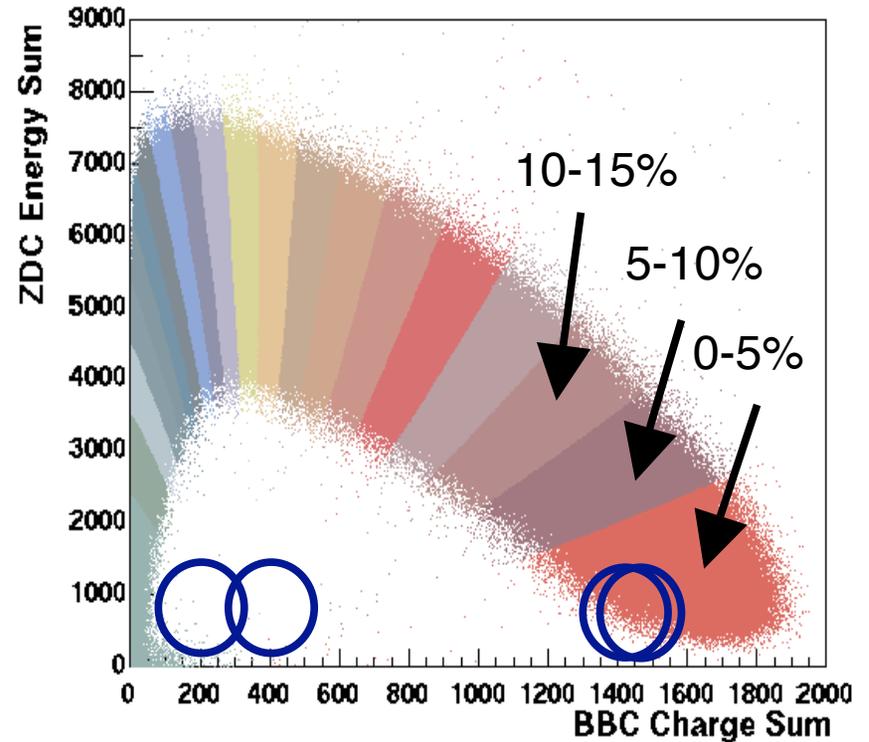


Collision Centrality Measurement ZeroDegreeCalorimeter



WA80 O+Au CERN

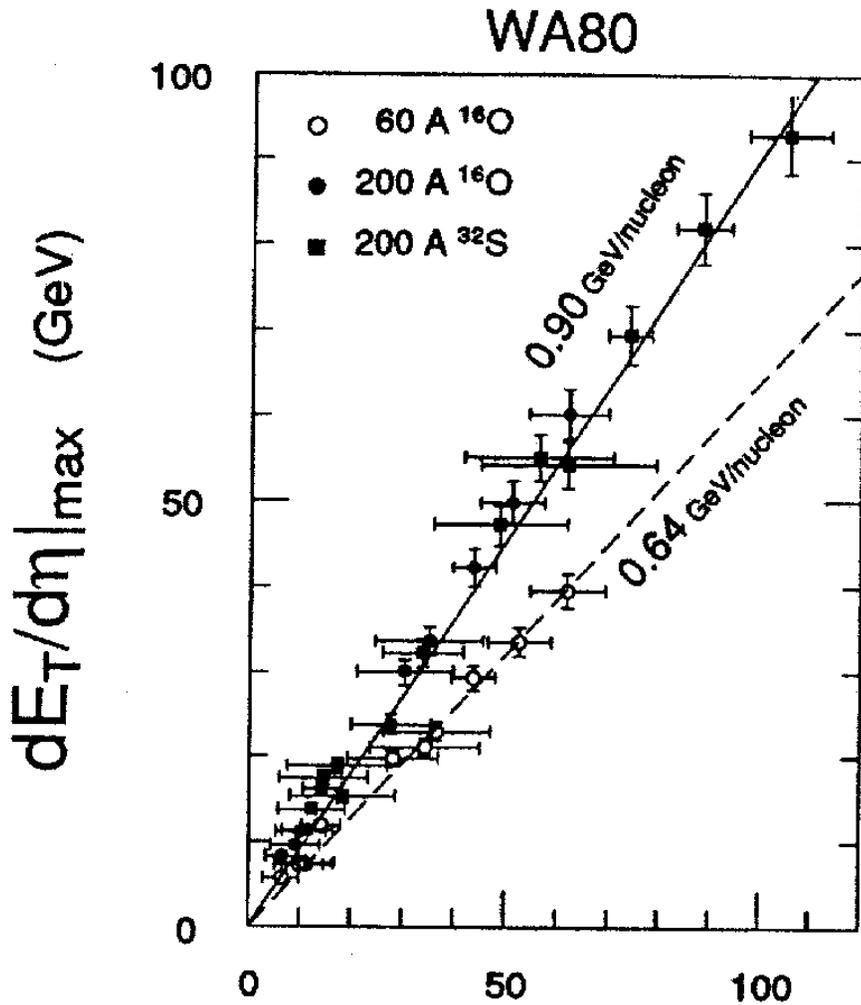
PHENIX at RHIC Au+Au



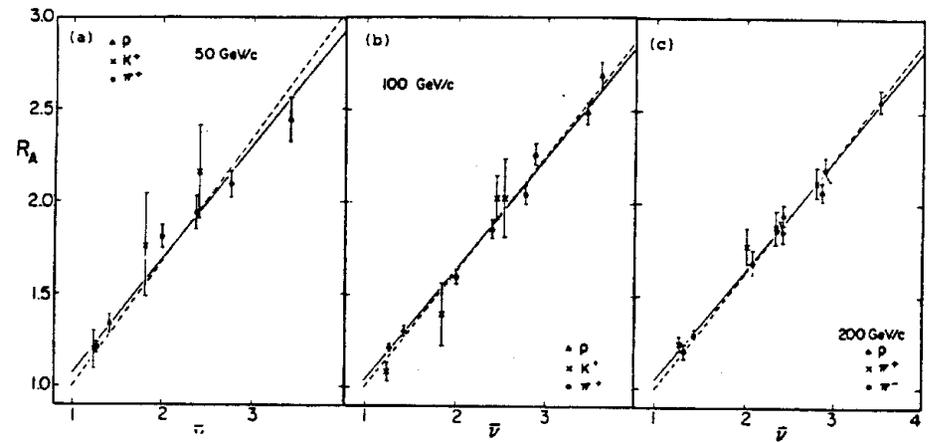
Extreme-Independent or Wounded Nucleon Models

- Number of Spectators (i.e. non-participants) N_s can be measured directly in Zero Degree Calorimeters (more complicated in Colliders)
- Enables unambiguous measurement of (projectile) participants = $A_p - N_s$
- For symmetric A+A collision $N_{\text{part}} = 2 N_{\text{projpart}}$
- Uncertainty principle and time dilation prevent cascading of produced particles in relativistic collisions $\lambda_{\text{h/m}_\pi c} > 10\text{fm}$ even at AGS energies: particle production takes place outside the Nucleus in a p+A reaction.
- Thus, Extreme-Independent models separate the nuclear geometry from the dynamics of particle production. The Nuclear Geometry is represented as the relative probability per B+A interaction w_n for a given number of total participants (WNM), projectile participants (WPNM), wounded projectile quarks (AQM), or other fundamental element of particle production.
- The dynamics of the elementary underlying process is taken from the data: e.g. the measured E_T distribution for a p-p collision represents, 2 participants, 1 n-n collision, 1 wounded projectile nucleon, a predictable convolution of quark-nucleon collisions.

WA80 proof of Wounded Nucleon Model at 60, 200 A GeV using ZDC



Original Discovery by W. Busza, et al
at FNAL $\langle n \rangle_{pA}$ vs $\langle \bar{\nu} \rangle$ (N_{coll})
PRD 22, 13 (1980)



$$R_A = \langle n \rangle_{pA} / \langle n \rangle_{pp} = (1 + \langle \bar{\nu} \rangle) / 2$$

$\langle N_{part} \rangle_{pA}$

$\langle N_{part} \rangle_{pp}$

PRC 44, 2736 (1991)

$$\overline{W} = \langle N_{part} \rangle$$

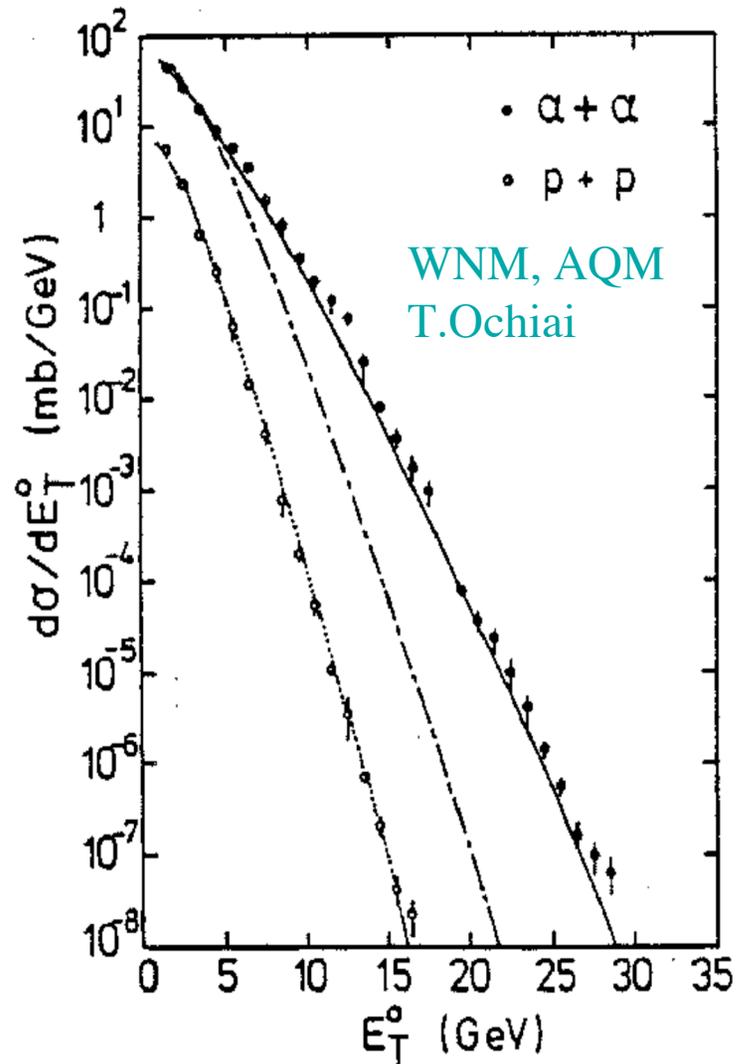
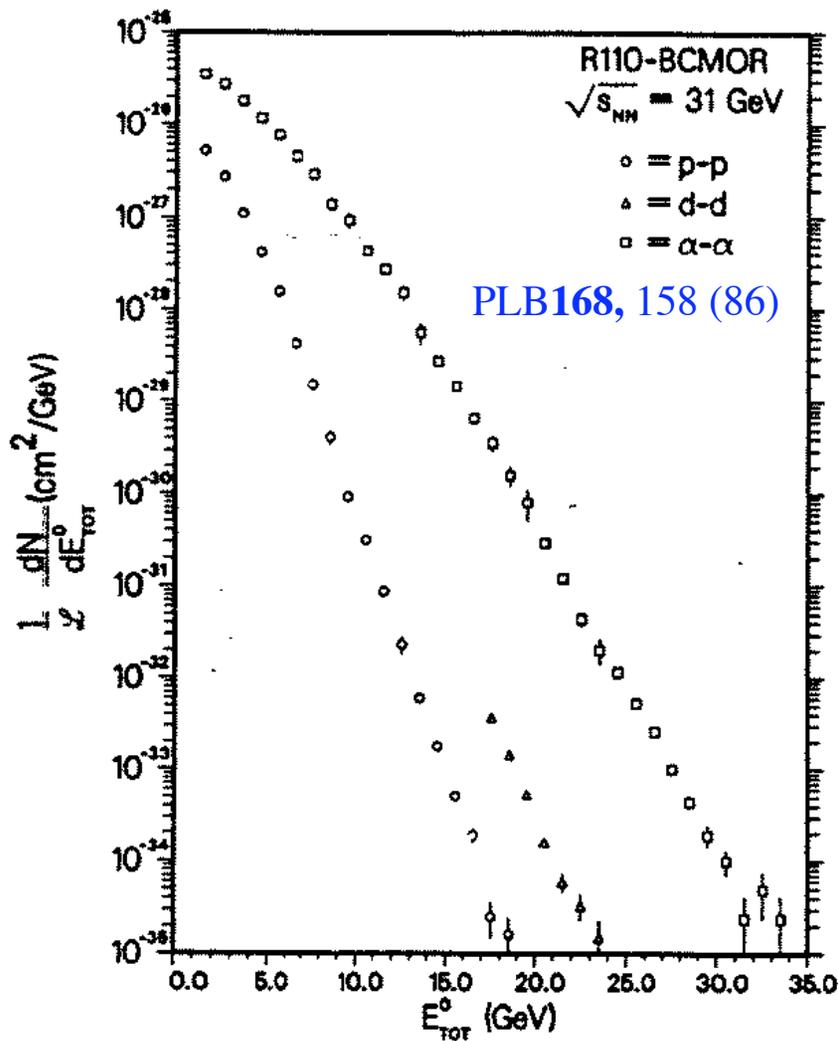
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Summary of Wounded Nucleon Models

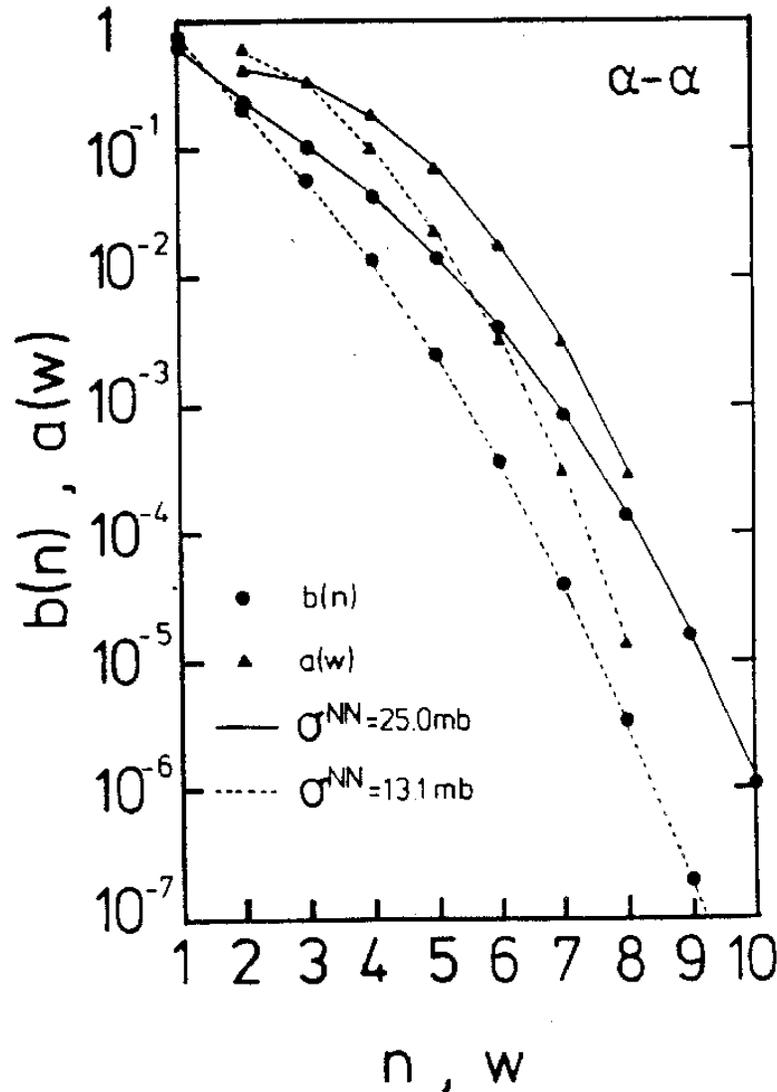
- The classical Wounded Nucleon (N_{part}) Model (WNM) of Bialas, Bleszynski and Czyz (NPB **111**, 461 (1976)) works only at CERN fixed target energies, $\sqrt{s_{\text{NN}}}\sim 20$ GeV.
- WNM overpredicts at AGS energies $\sqrt{s_{\text{NN}}}\sim 5$ GeV (WPNM works at mid-rapidity)
- WNM underpredicts for $\sqrt{s_{\text{NN}}}\geq 31$ GeV

ISR-BCMOR-pp,dd,αα $\sqrt{s_{NN}}=31\text{ GeV}$ WNM FAILS!



Note WNM edge is parallel to p-p data!

AQM works with σ_I (and p_0) or σ_{obs} for N_{part}



Observed $\sigma_{obs}=13\text{mb}$ is $\sim 0.5 \sigma_I=25\text{mb}$. Use either σ_{obs} for the nuclear geometry calculation or σ_I (properly taking account of the probability $p_0 \sim 0.5$ of detecting zero in the detector for a collision).

Calculation by T.Ochiai, ZPC **35**, 209 (87), PLB **206**, 535 (88).

This is the answer to Kopeliovich's complaint

Quick Course in Statistics

- Two of the most popular statistics are the sum and the average:

$$S_n \equiv \sum_{i=1}^n x_i \quad \bar{x}_{(n)} \equiv \frac{1}{n} \sum_{i=1}^n x_i = \frac{1}{n} S_n \quad (11)$$

where the x_i are n samples from a the same population or probability density function, $f(x)$.

- The probability distribution of a random variable S_n , which is itself the sum of n **independent** random variables with a common distribution $f(x)$:

$$S_n = x_1 + x_2 + \cdots + x_n \quad (12)$$

is given by $f_n(x)$, the n -fold convolution of the distribution $f(x)$:

$$f_n(x) = \int_0^x dy f(y) f_{n-1}(x-y) \quad . \quad (13)$$

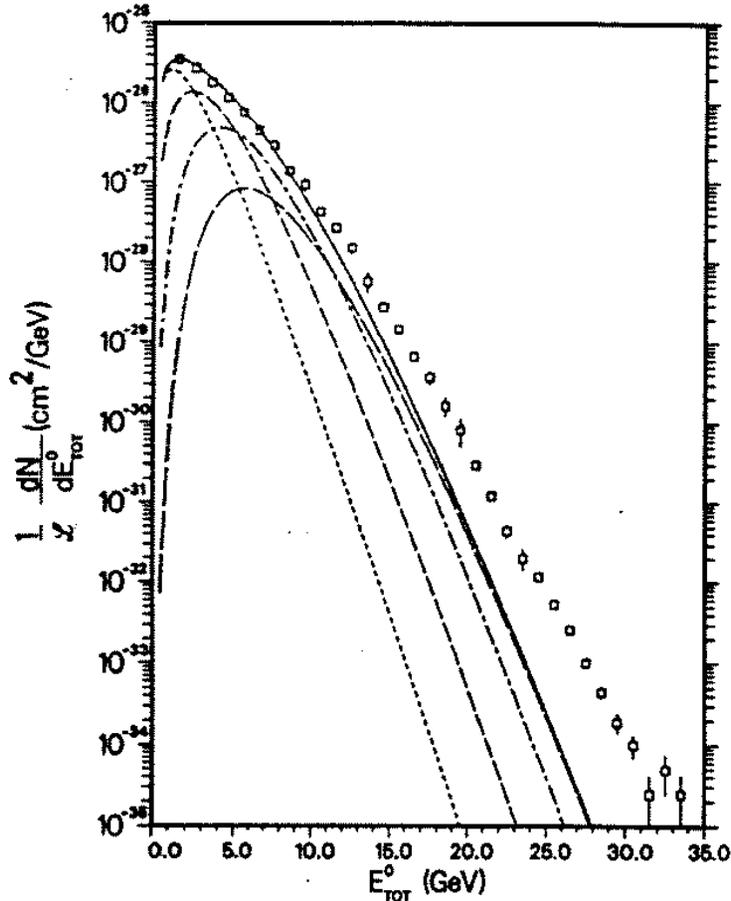
The mean, $\mu_n = \langle S_n \rangle$, and standard deviation, σ_n , of the n -fold convolution obey the familiar rule

$$\mu_n = n\mu \quad \sigma_n = \sigma\sqrt{n} \quad , \quad (14)$$

where μ and σ are the mean and standard deviation of the distribution $f(x)$.

- A complementary case is that of a scale transformation $x \rightarrow nx$. The behavior of the mean and the standard deviation for a scale transformation is $\mu \rightarrow n\mu$, $\sigma \rightarrow n\sigma$, which is quite different than the behavior of the standard deviation under convolution (Eq. 14).

Details of WNM Calculation



The WPNM calculation for a B+A reaction

$$\left(\frac{d\sigma}{dE_T}\right)_{\text{WPNM}} = \sigma_{BA} \sum_{n=1}^B w_n P_n(E_T) \quad (15)$$

- σ_{BA} is the measured B+A cross section in the detector aperture,
- w_n is the relative probability for n projectile nucleons in the B+A reaction.
- $P_n(E_T)$ is the calculated E_T distribution on the detector aperture for n **independently interacting** projectile nucleons.
- If $f_1(E_T)$ is the measured E_T spectrum on the detector aperture for one projectile nucleon, and p_0 is the probability for the elementary collision to produce no signal on the detector aperture, then, the correctly normalized E_T distribution for one projectile nucleon collision is:

$$P_1(E_T) = (1 - p_0)f_1(E_T) + p_0\delta(E_T) \quad , \quad (16)$$

where $\delta(E_T)$ is the Dirac delta function and $\int f_1(E_T) dE_T = 1$.

- $P_n(E_T)$ (including the p_0 effect) is obtained by **convoluting** $P_1(E_T)$ with itself $n - 1$ times

$$P_n(E_T) = \sum_{i=0}^n \frac{n!}{(n-i)! i!} p_0^{n-i} (1 - p_0)^i f_i(E_T) \quad (17)$$

where $f_0(E_T) \equiv \delta(E_T)$ and $f_i(E_T)$ is the i -th convolution of $f_1(E_T)$:

$$f_i(x) = \int_0^x dy f(y) f_{i-1}(x - y) \quad . \quad (18)$$

A Favorite Function

The Gamma distribution is an example of a probability density function (pdf) which has particularly simple properties under convolutions and scale transformations. The Gamma distribution is a function of a continuous variable x and has parameters p and b

$$f(x) = f_{\Gamma}(x, p, b) = \frac{b}{\Gamma(p)} (bx)^{p-1} e^{-bx} \quad (19)$$

where

$$p > 0, \quad b > 0, \quad 0 \leq x \leq \infty$$

$\Gamma(p) = (p-1)!$ if p is an integer, and $f(x)$ is normalized, $\int_0^{\infty} f(x) dx = 1$. The mean and standard deviation of the distribution are

$$\mu \equiv \langle x \rangle = \frac{p}{b} \quad \sigma \equiv \sqrt{\langle x^2 \rangle - \langle x \rangle^2} = \frac{\sqrt{p}}{b} \quad \frac{\sigma^2}{\mu^2} = \frac{1}{p} \quad (20)$$

The n -fold convolution of the Gamma distribution (Eq. 19) is simply given by the function

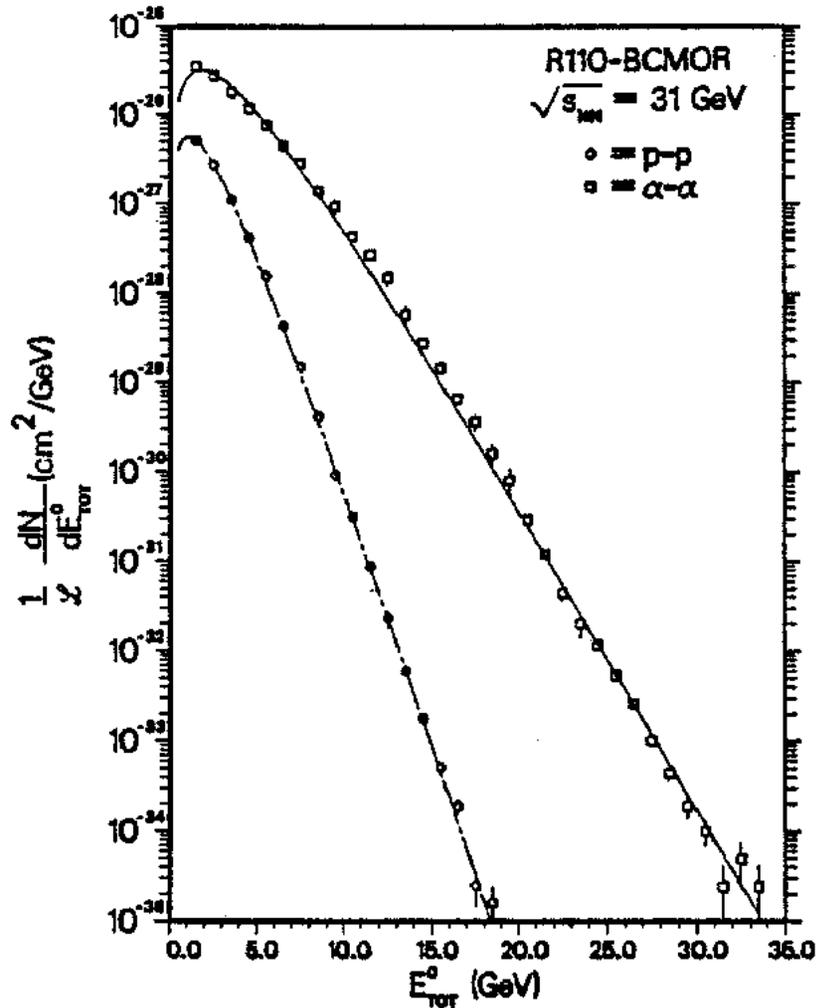
$$f_n(x) = \frac{b}{\Gamma(np)} (bx)^{np-1} e^{-bx} = f_{\Gamma}(x, np, b) \quad (21)$$

i.e. $p \rightarrow np$ and b remains unchanged. Note that the mean and standard deviation of Eq. 21

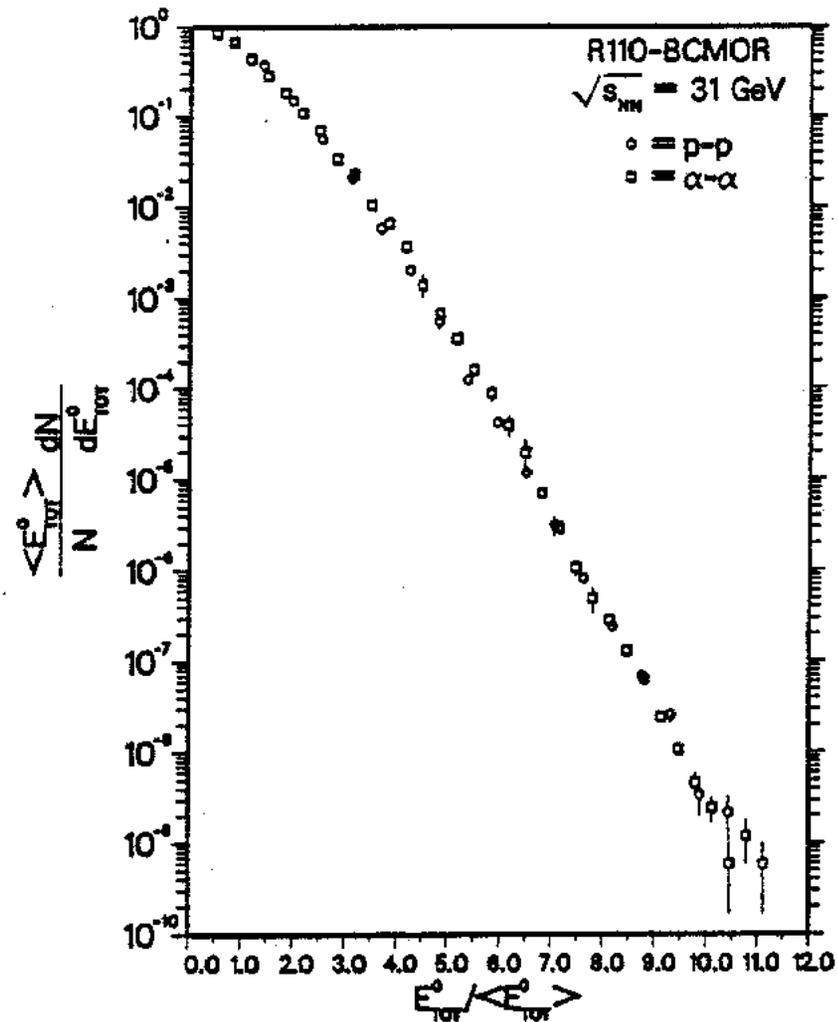
$$\mu_n = \frac{np}{b} \quad \sigma_n = \frac{\sqrt{np}}{b} \quad \frac{\sigma_n}{\mu_n} = \frac{1}{\sqrt{np}} \quad (22)$$

when compared to Eq. 20 explicitly obey Eq. 14. **To summarize, the n -th convolution of the Gamma distribution $f_{\Gamma}(x, p, b)$ is $f_{\Gamma}(x, np, b)$ — b remains unchanged.**

But-Gamma Dist. fits uncover Scaling in the mean over 10 decades??

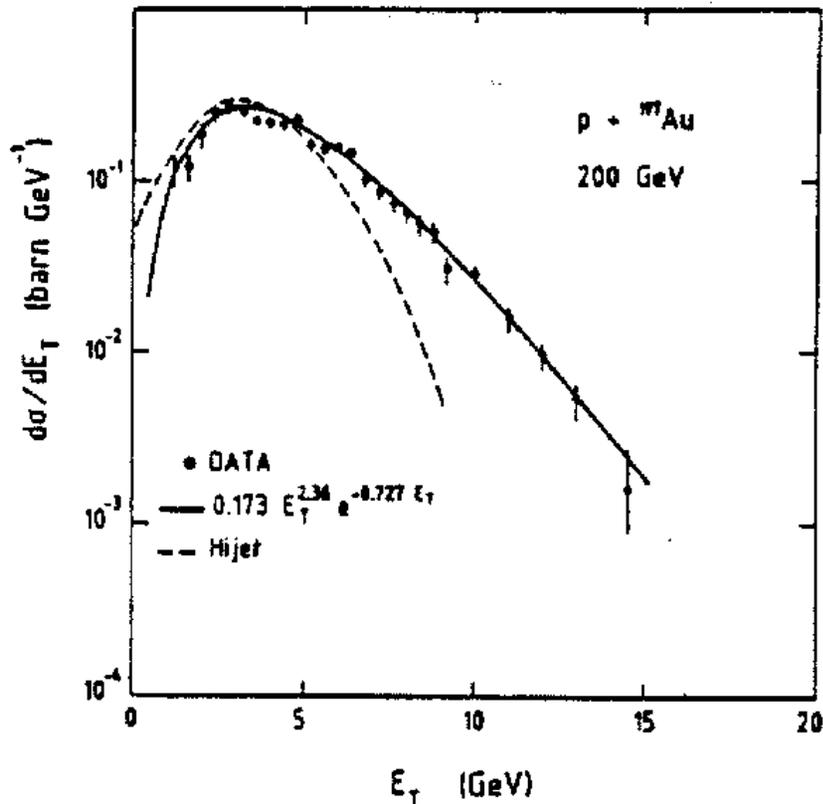


p-p $p=2.50 \pm 0.06$ \square - \square $p=2.48 \pm 0.05$

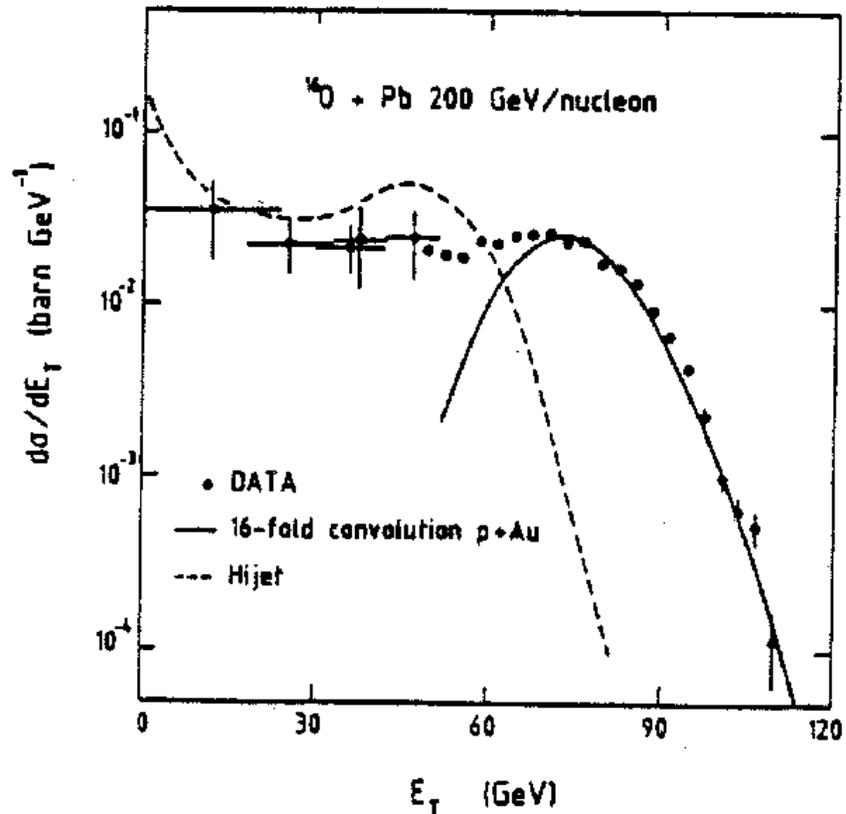


Is it Physics or a Fluke?

First RHI data NA35 (NA5 Calorimeter)
 CERN $^{16}\text{O}+\text{Pb}$ $\sqrt{s_{\text{NN}}}=19.4$ GeV midrapidity
 PLB 184, 271 (1987)



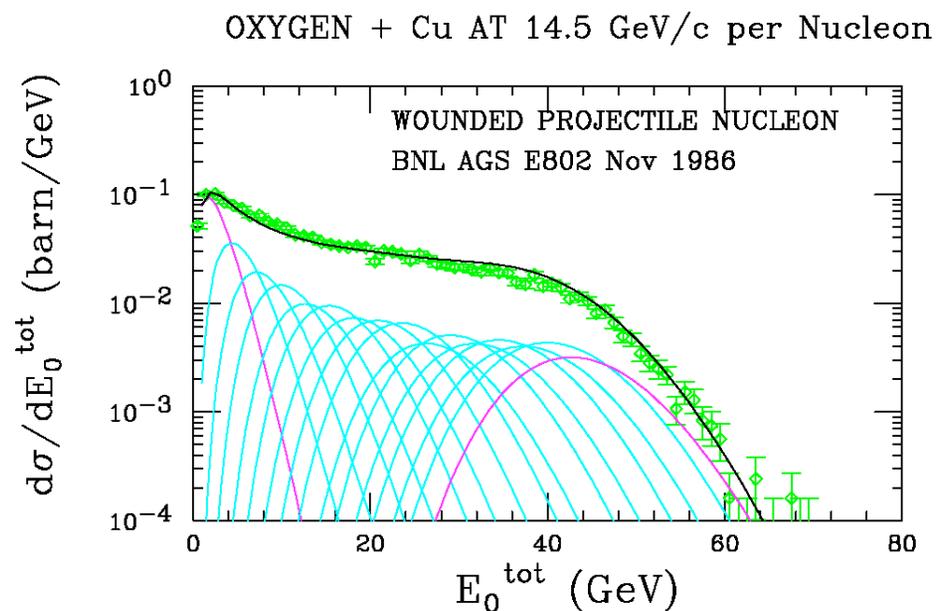
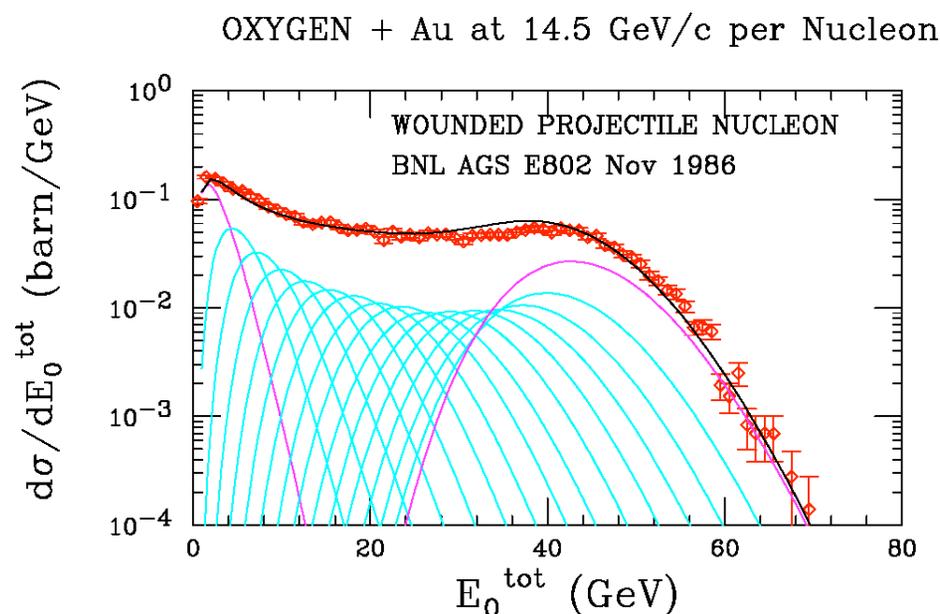
$p+\text{Au}$ is a \square dist w. $p=3.36$



Upper Edge of O+Pb is 16
 convolutions of p+Au. WPNM!!

E802-O+Au, O+Cu midrapidity at $\sqrt{s_{NN}}=5.4\text{GeV}$ WPNM works in detail

PLB **197**, 285 (1987)
ZPC **38**, 35 (1988)



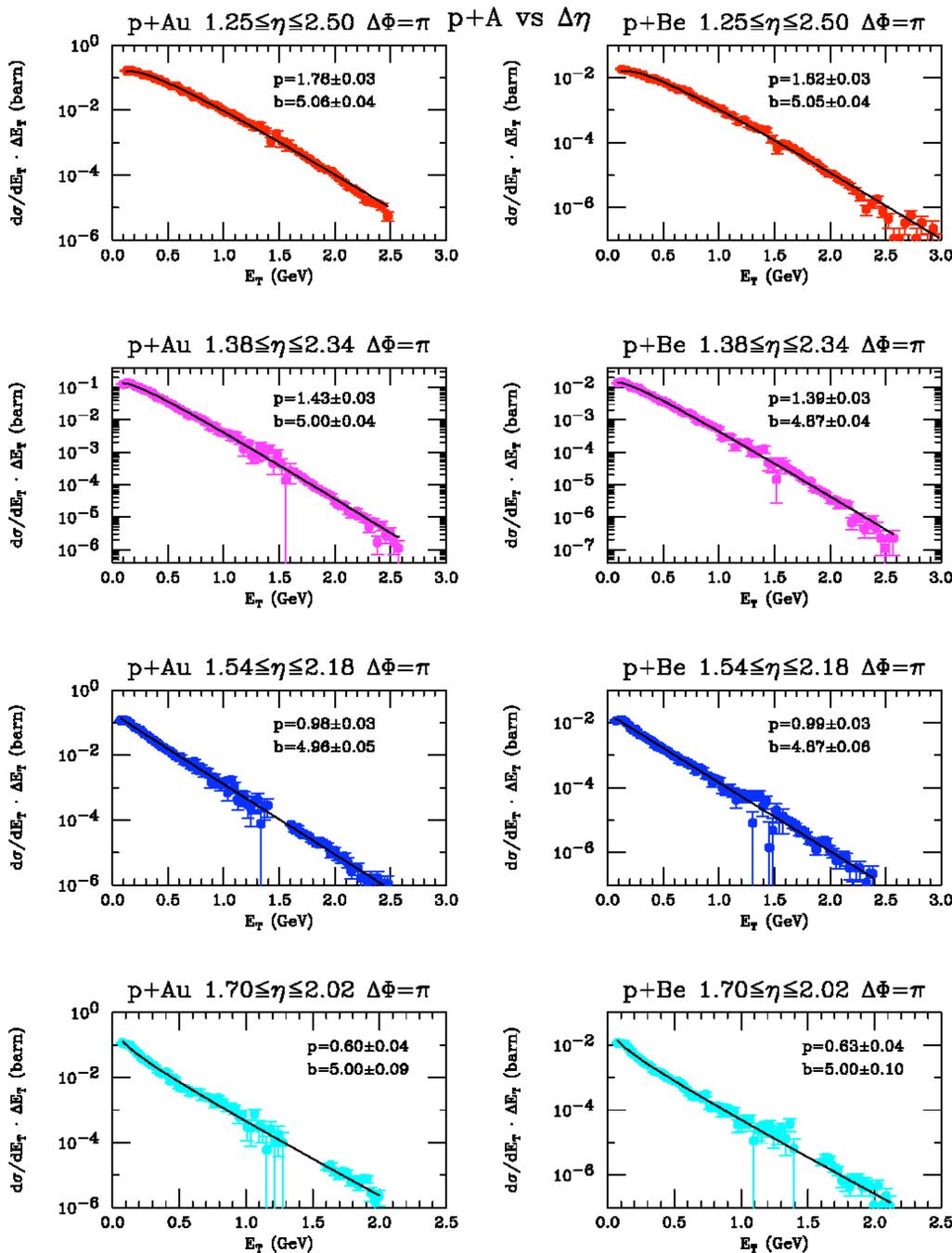
- Maximum energy in O+Cu ~ same as O+Au--Upper edge of O+Au identical to O+Cu $d\sigma/dE * 6$
- Indicates large stopping at AGS ^{16}O projectiles stopped in Cu so that energy emission (mid-rapidity) ceases
- Full O+Cu and O+Au spectra described in detail by WPNM based on measured p+Au

E802-AGS

pBe & pAu have same
shape at midrapidity
over a wide range of $\Delta\eta$

PRC **63**, 064602 (2001)

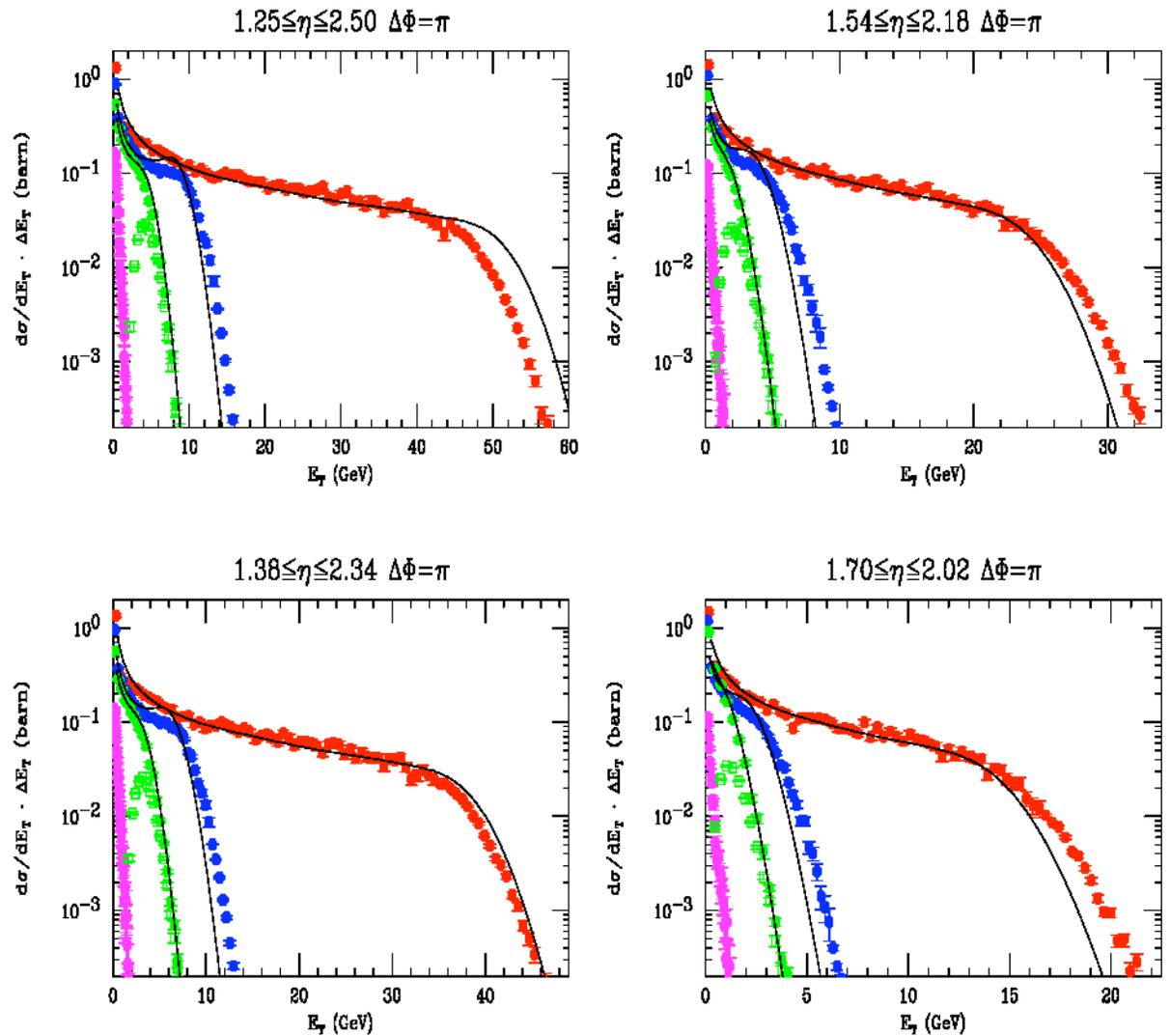
- confirms previous measurement
PRC **45**, 2933 (1992)
that pion distribution from second
collision shifts by > 0.8 units in y ,
out of aperture. Explains WPNM.



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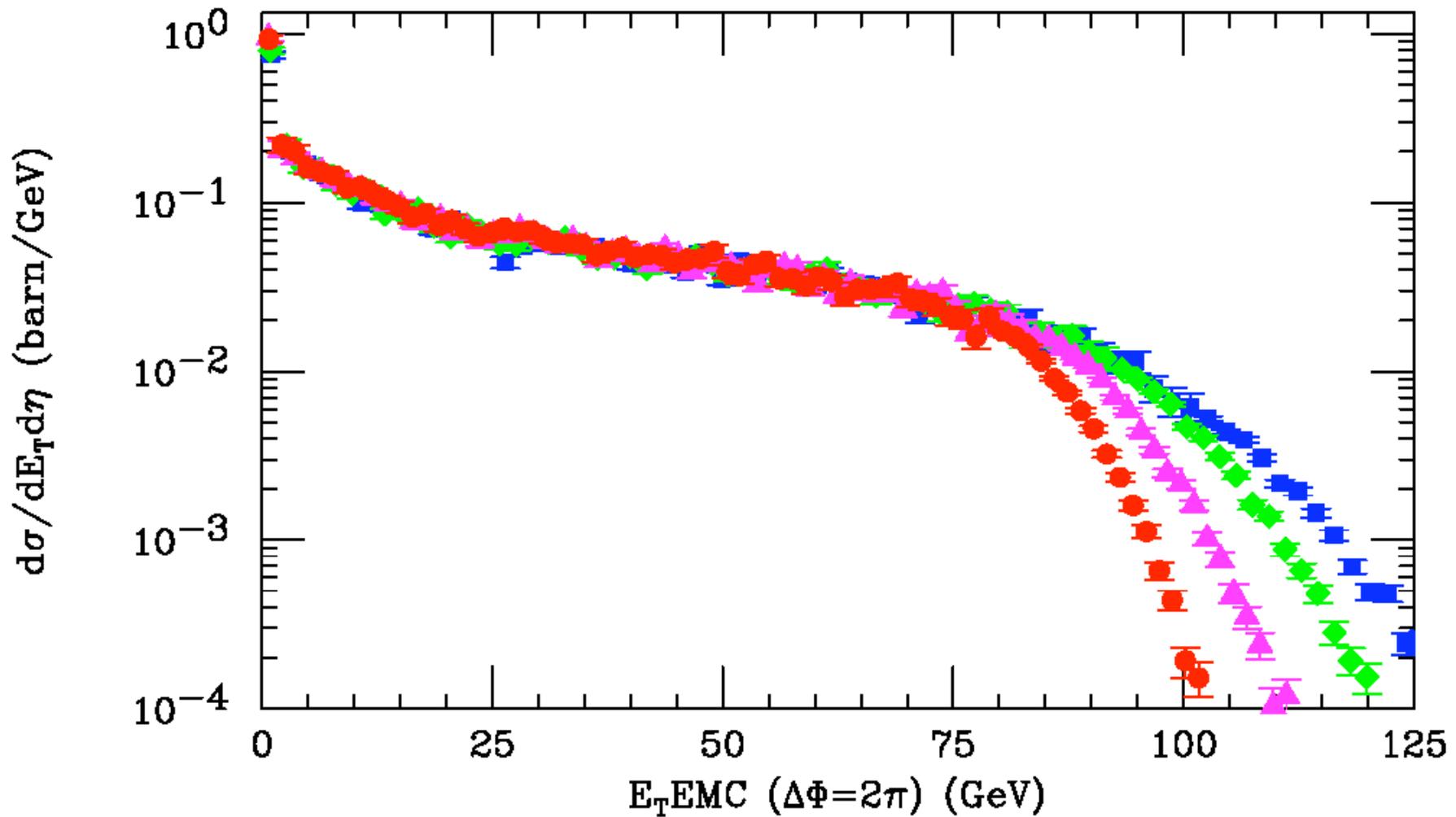
E802-AGS $E_T \sim$ follow WPNM vs \square



Spectra in different apertures hard to compare--Normalize 2 different ways

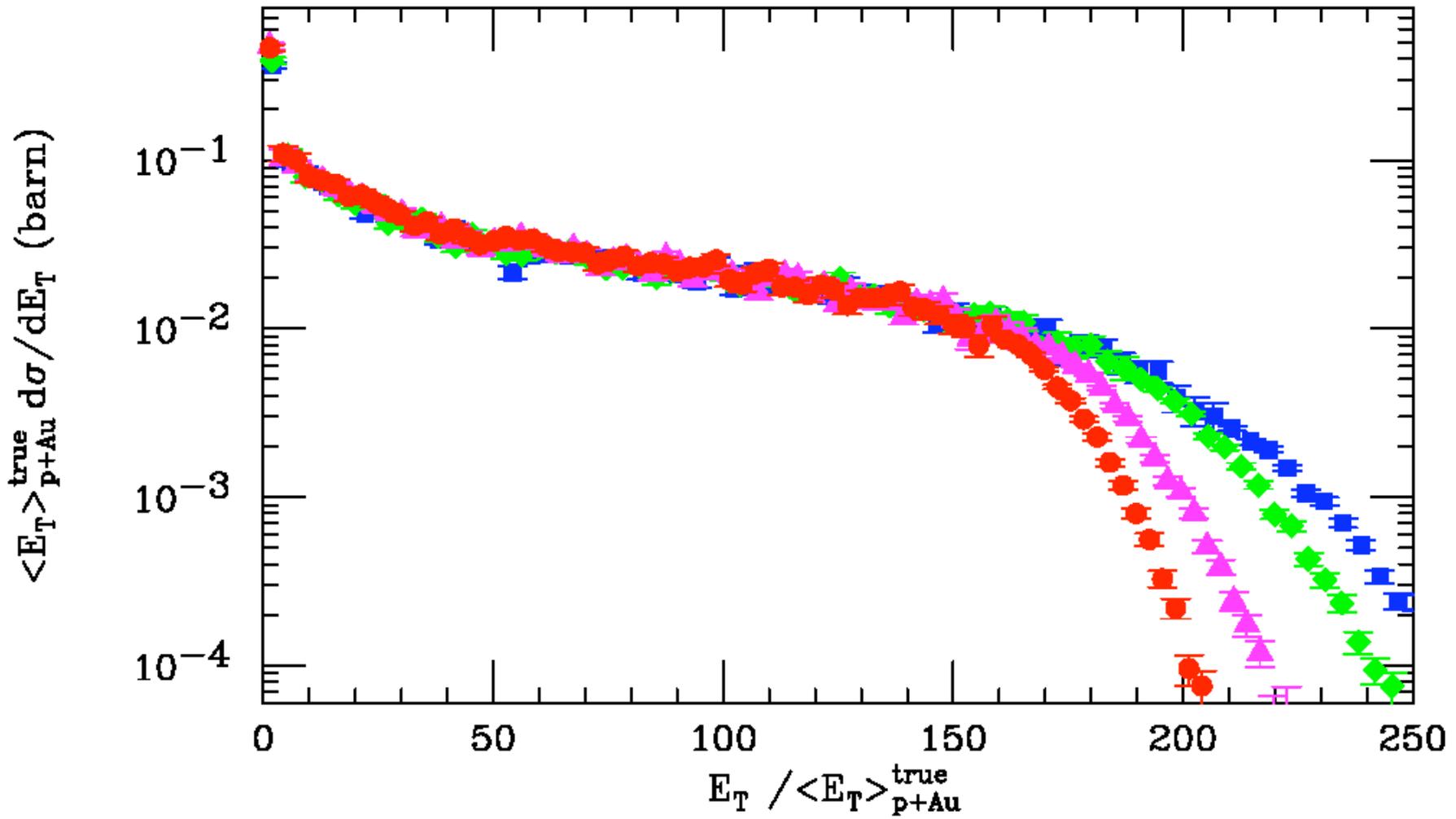
E802-AGS corrected to full azimuth

E802-AGS corr to $\Delta\eta=1$ and $\Delta\Phi=2\pi$

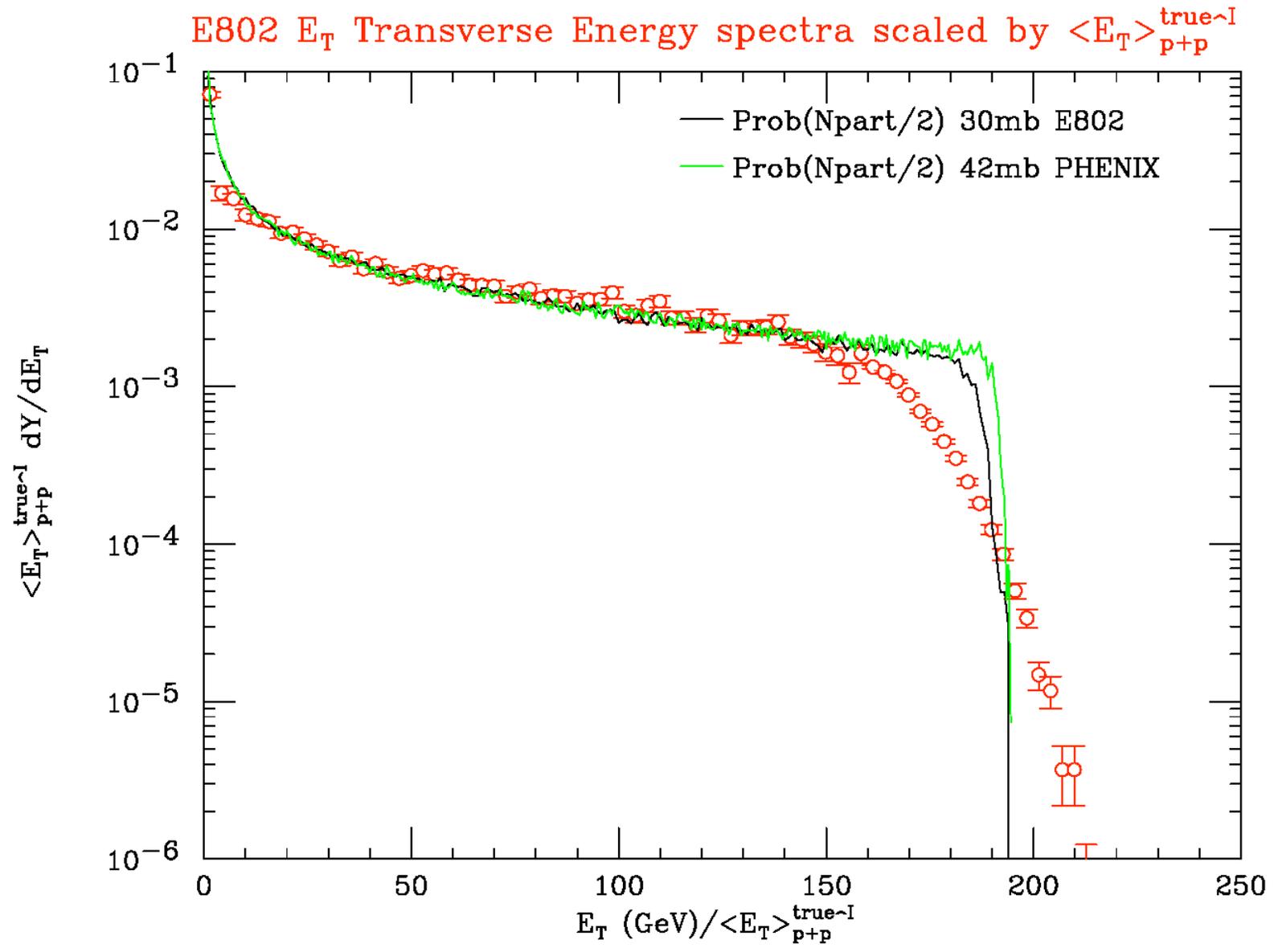


Normalized plots in units of $\langle E_T \rangle^{\text{true}} = (1-p_0) \langle E_T \rangle^{\text{obs}}$

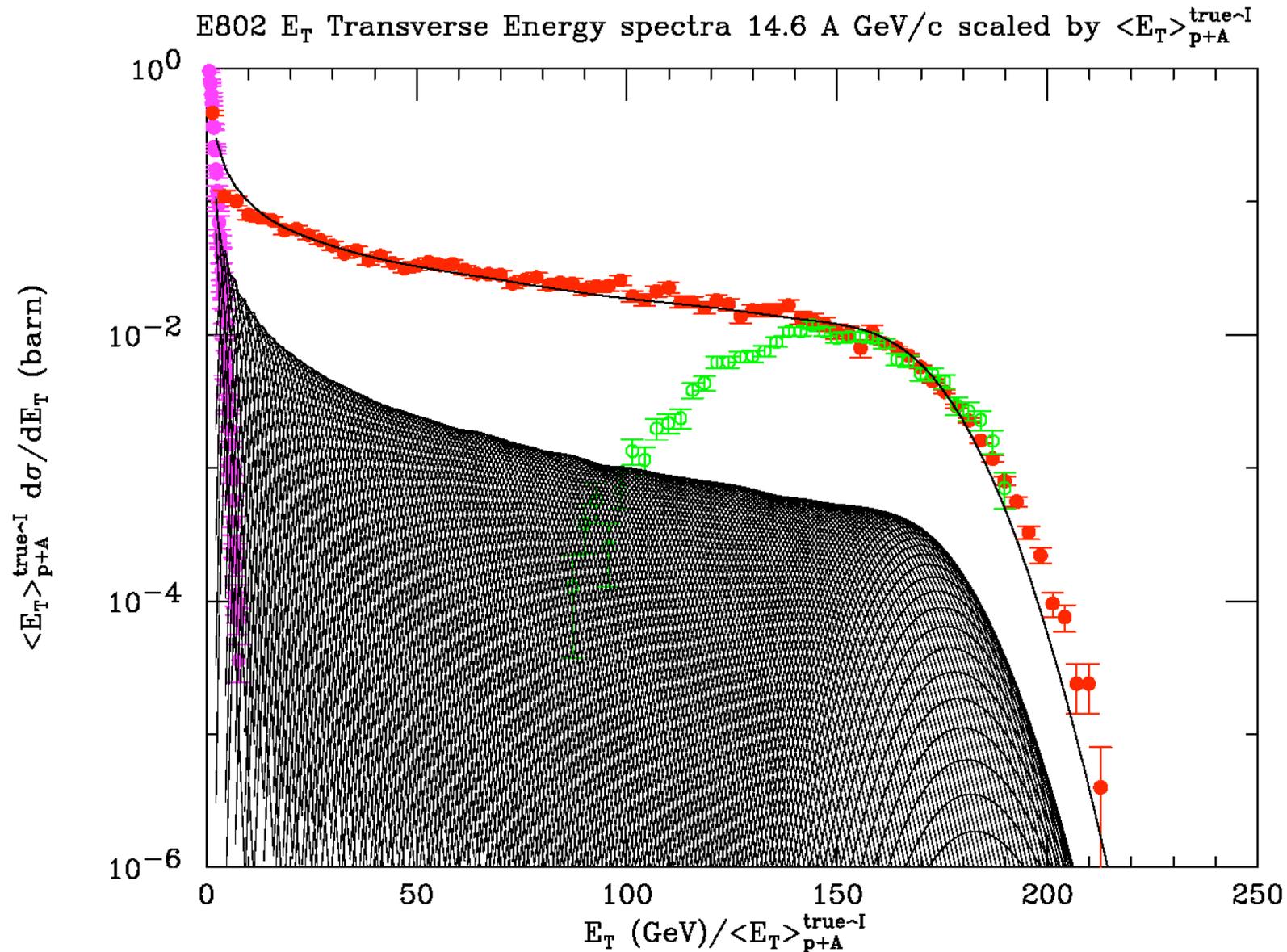
$1.22 \leq \eta \leq 2.50 \quad \Delta\Phi = \pi \quad \text{E802-AGS}$



Illustrate the Nuclear Geometry

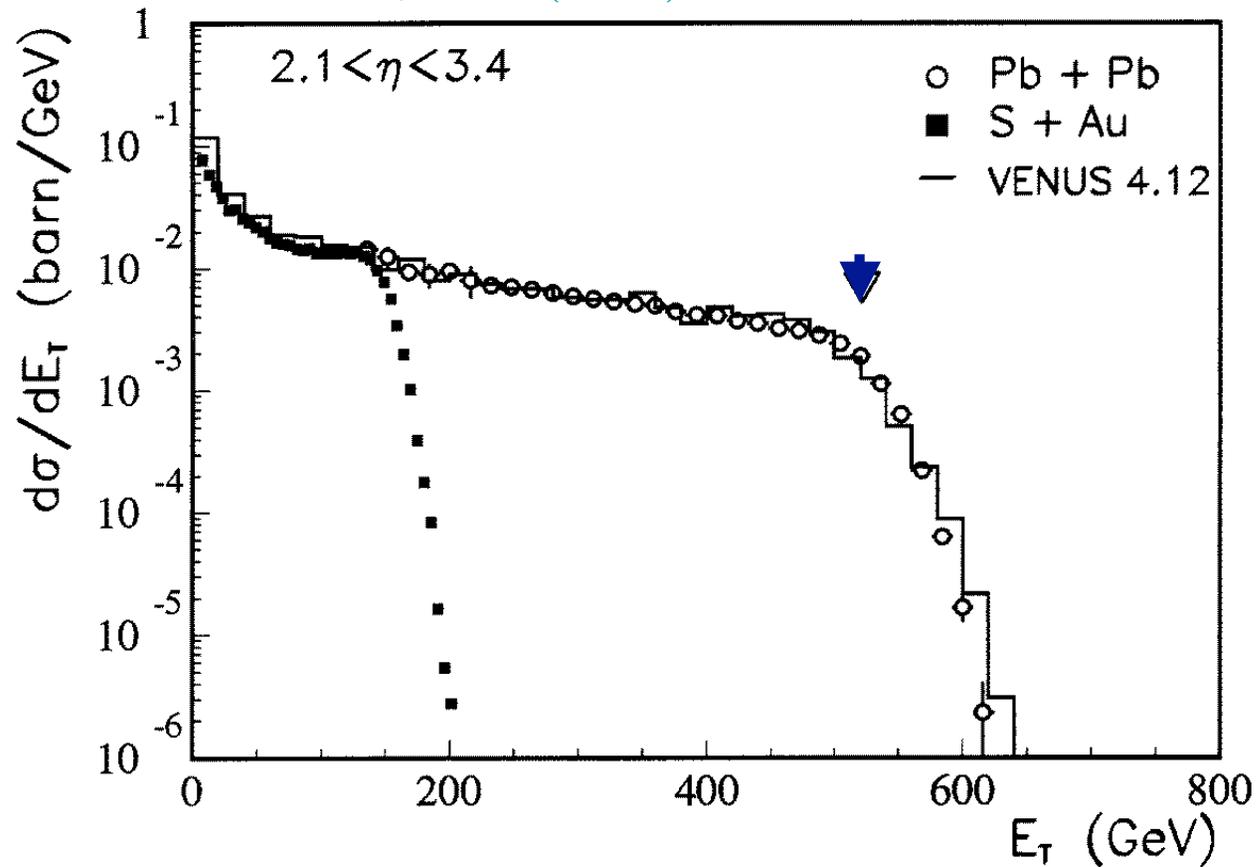


E802 W(p)NM calc Au+Au--details



NA35-->NA49 Pb+Pb $\sqrt{s_{NN}}=17$ GeV

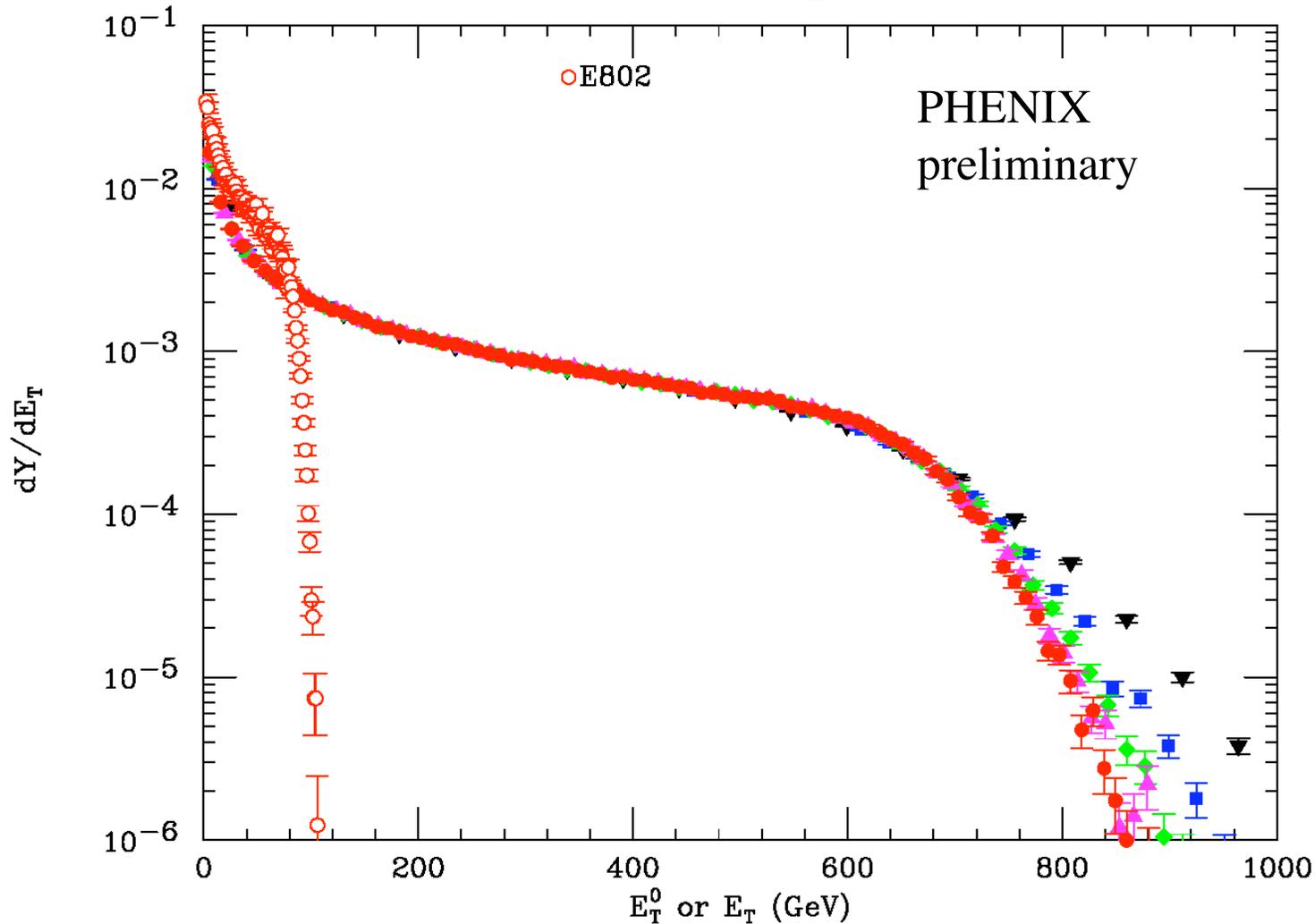
PRL 75, 3814 (1995)



$E_T(2.1-3.4) \rightarrow dE_T/d\eta = 405$ GeV @ $\sqrt{s_{NN}}=17$ GeV

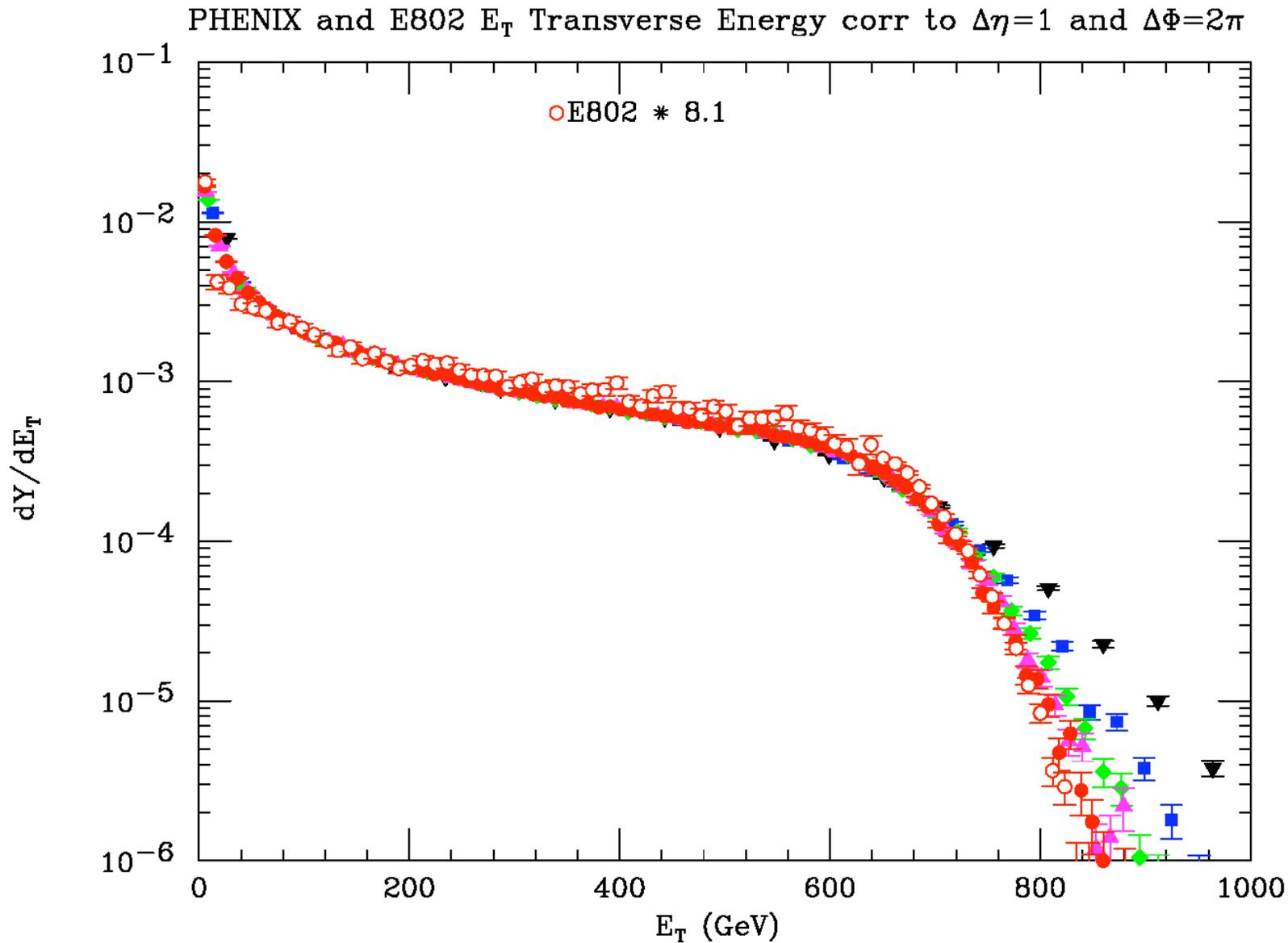
PHENIX and E802 E_T compared

PHENIX and E802 E_T Transverse Energy corr to $\Delta\eta=1$ and $\Delta\Phi=2\pi$

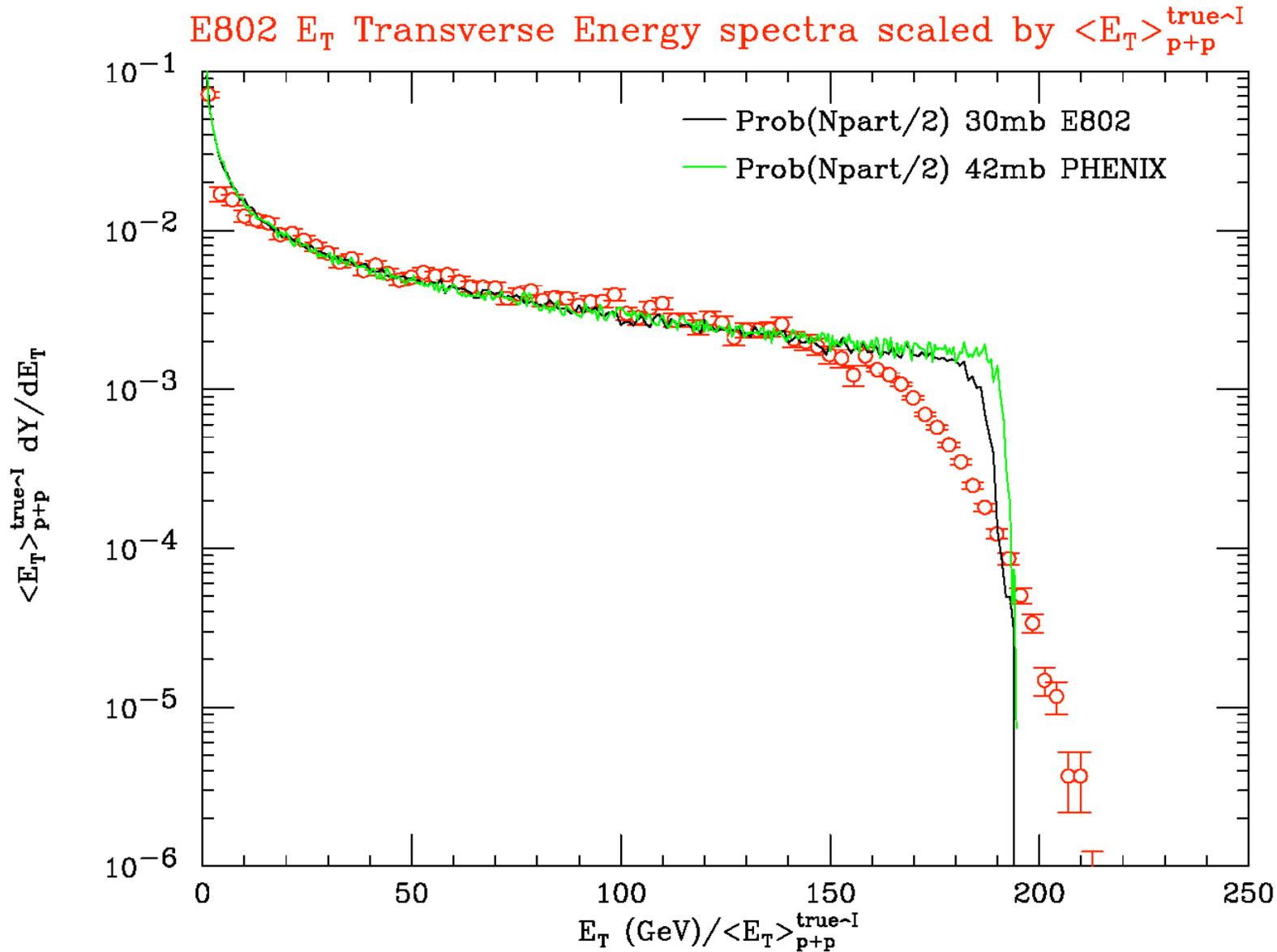


E877 $dE_T/d\eta=200$ GeV @ $\sqrt{s_{NN}}=4.8$ GeV PHENIX $dE_T/d\eta \sim 680$ GeV @ $\sqrt{s_{NN}}=200$ GeV

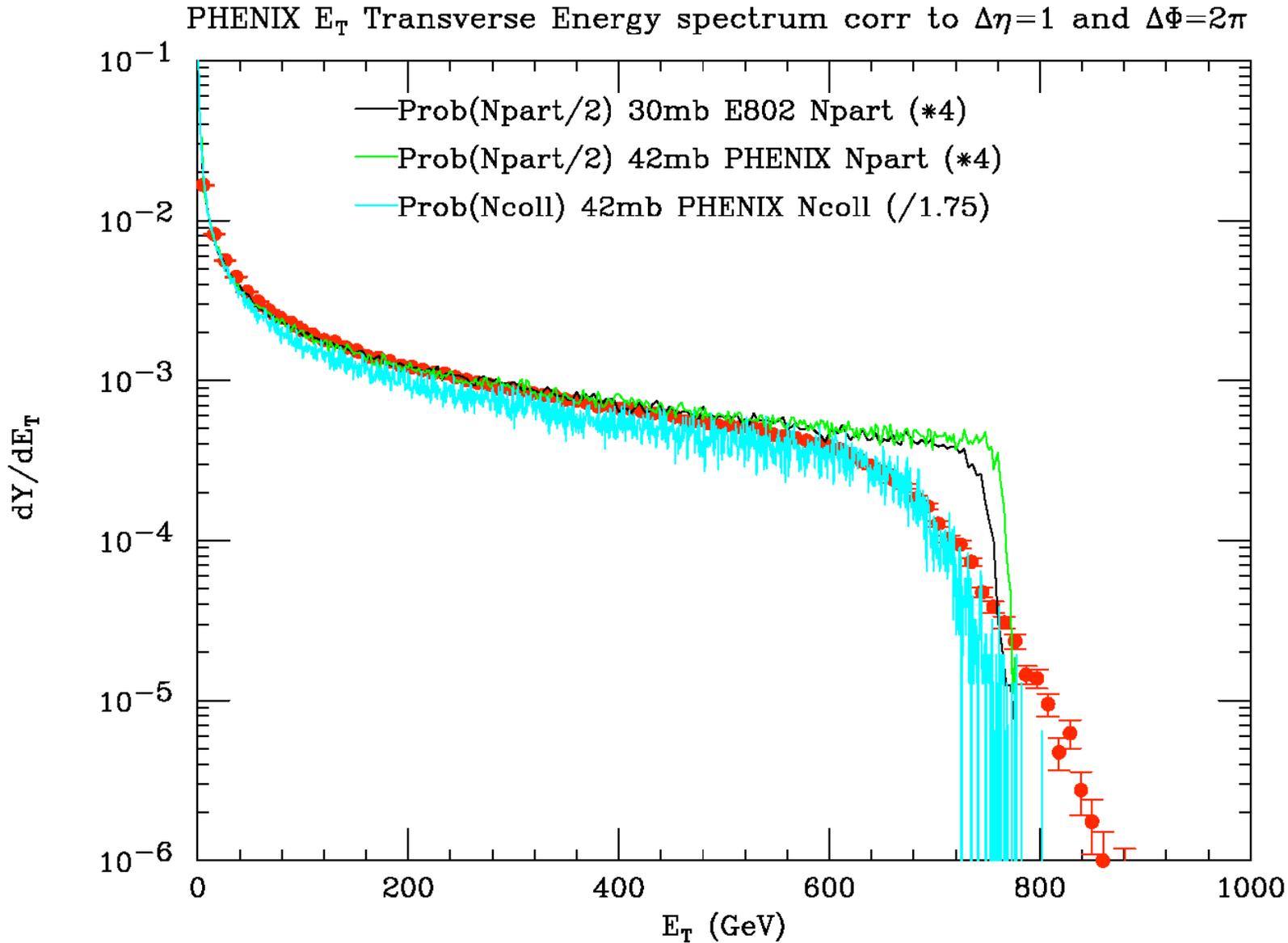
Au+Au E_T spectra at AGS and RHIC are the same shape!!!



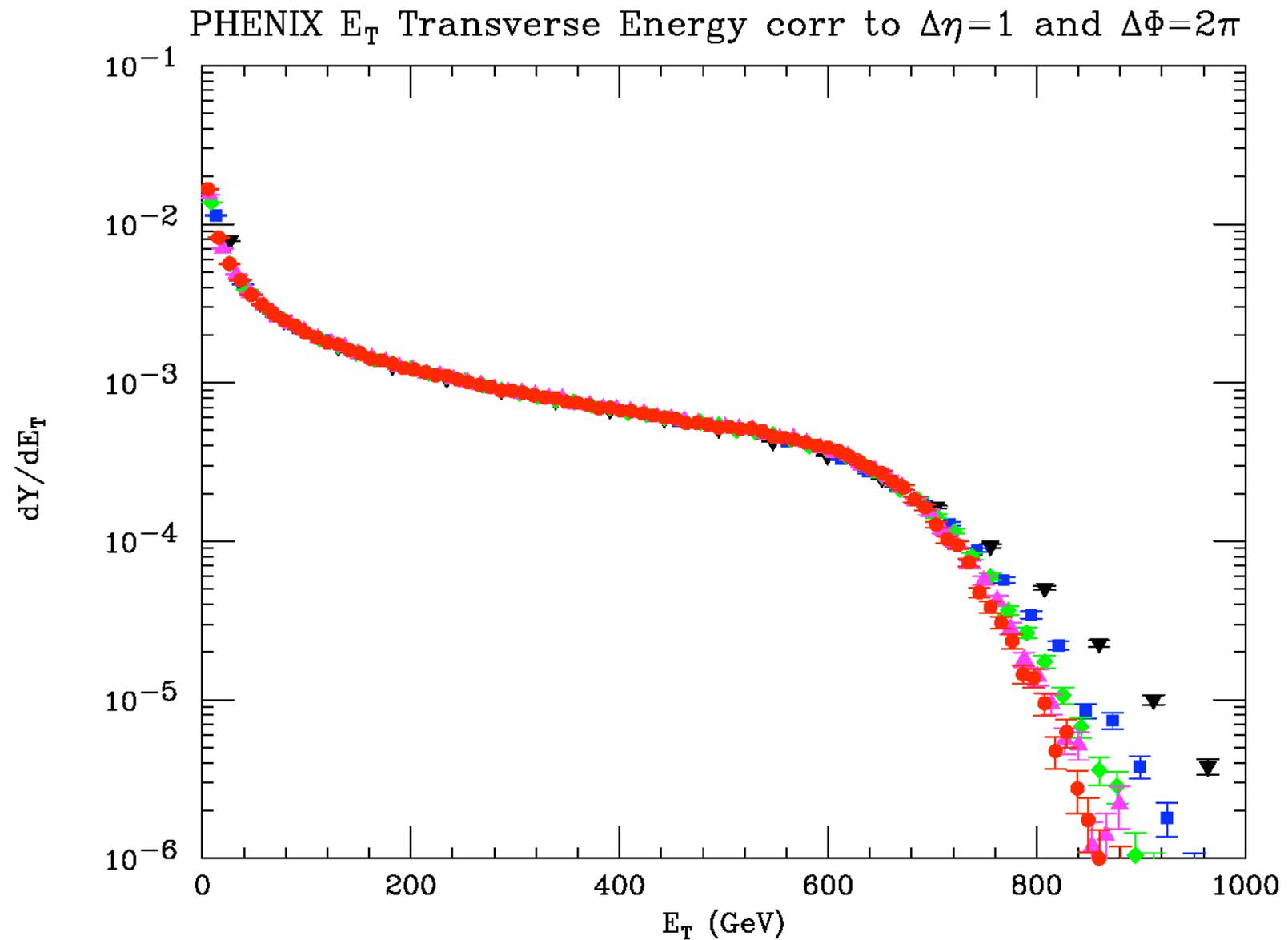
AuAu spectrum at AGS follows WN geometry



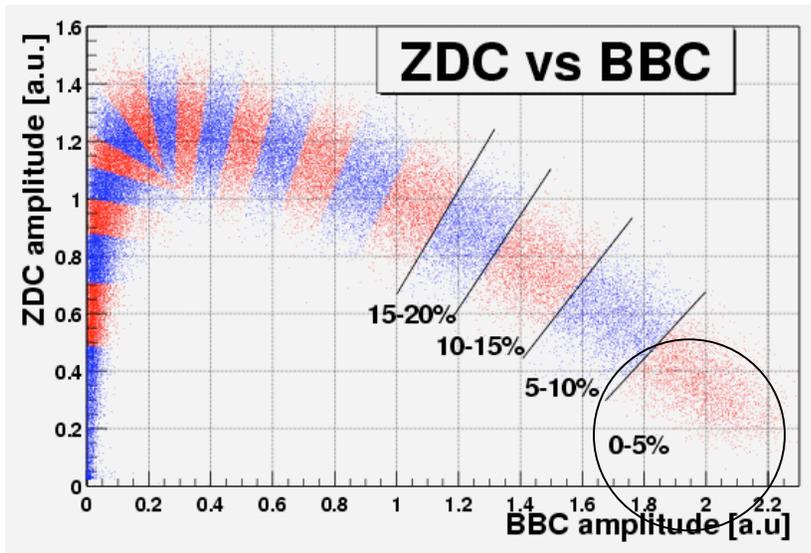
RHIC 2-3 times more E_T than WNM but:



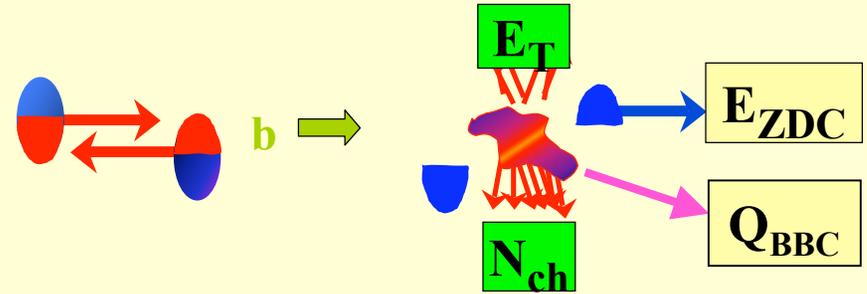
Are upper edge fluctuations random?



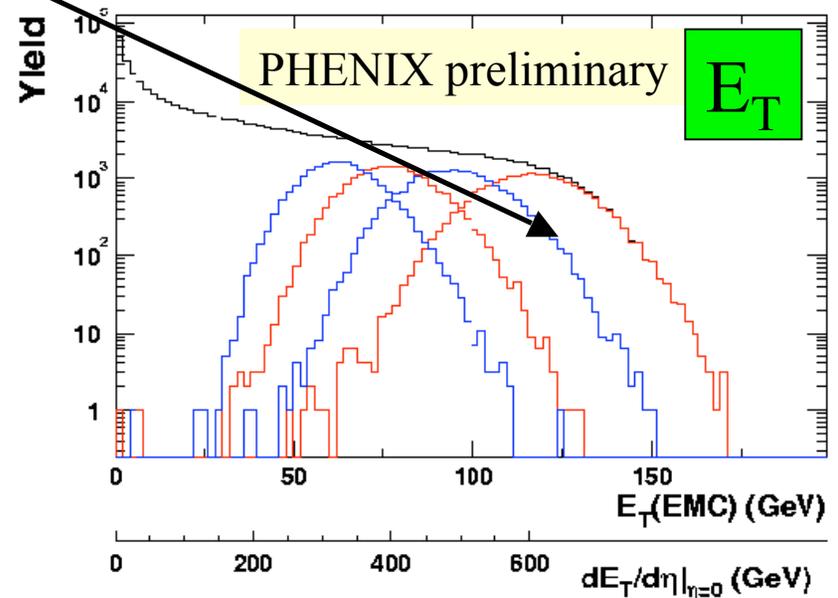
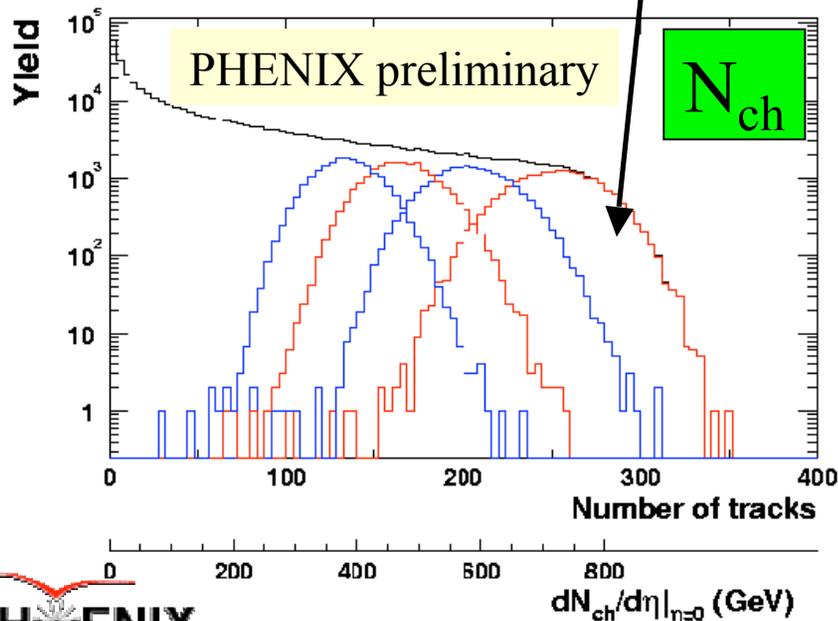
Fluctuations for central collisions



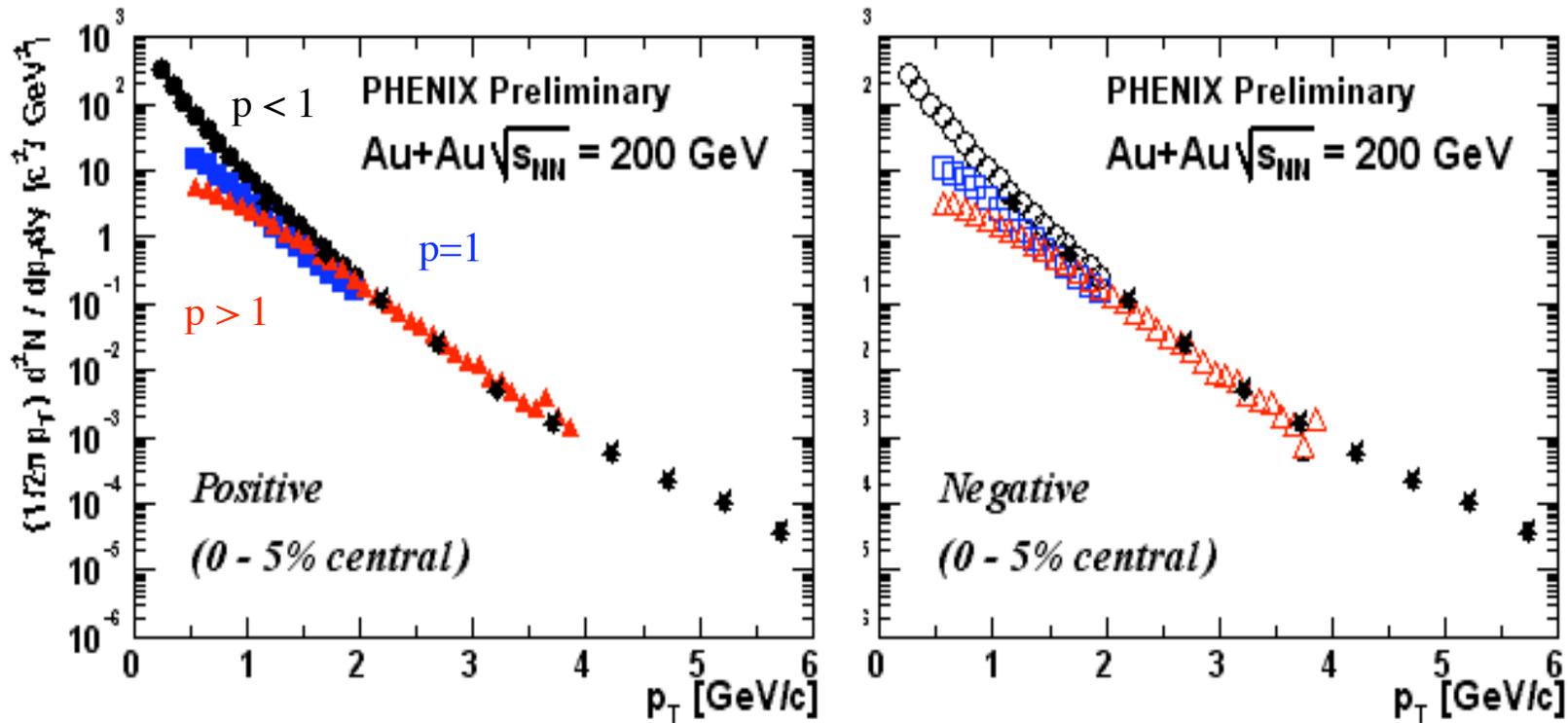
Define centrality classes: ZDC vs BBC



Extract N participants: Glauber model



Inclusive p_T spectra are Gamma Distributions

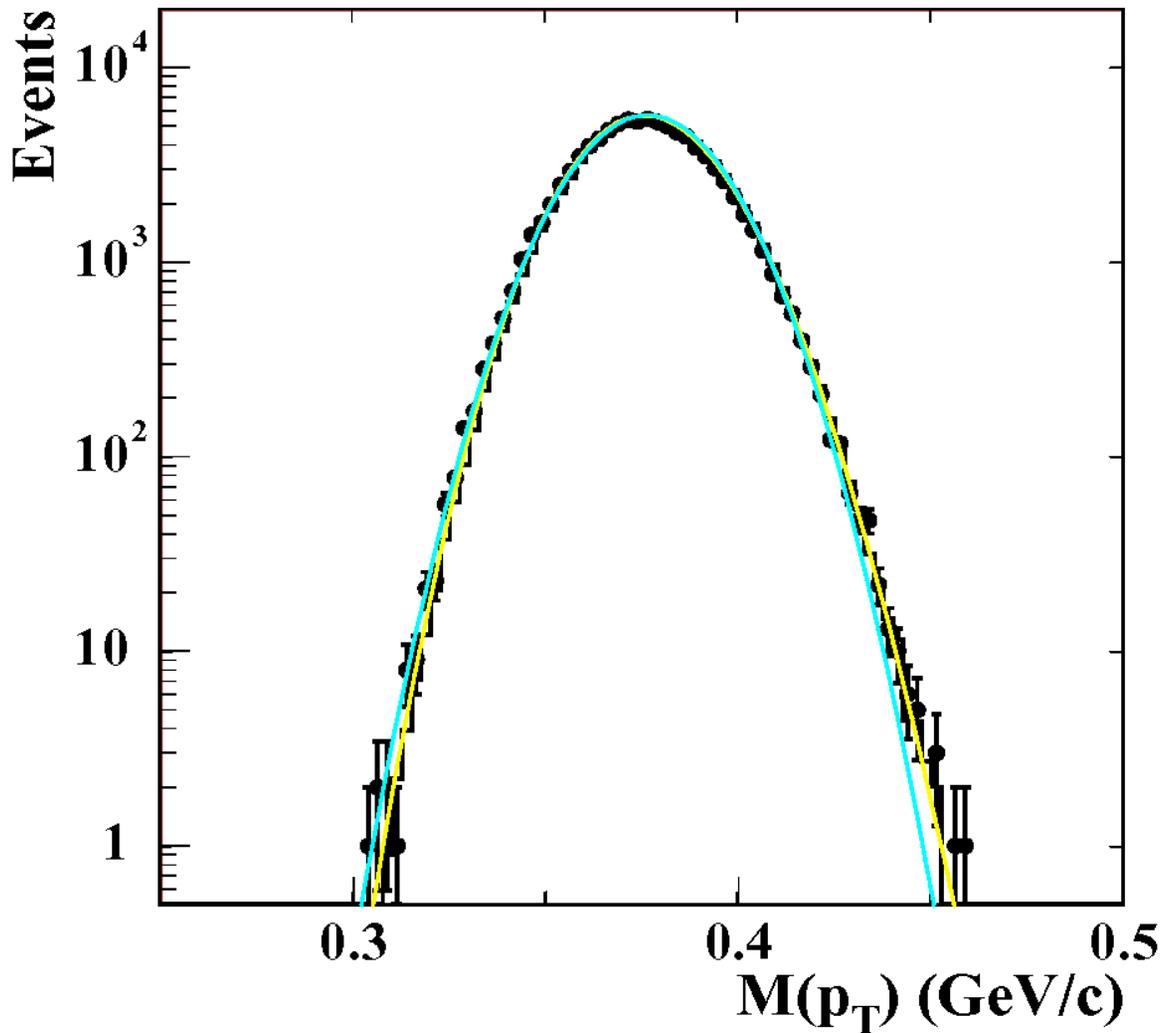


Note: $d\Omega = p_T dp_T \Omega(p_T, p) = dp_T \Omega(p_T, p+1)$

Event by Event average p_T (M_{p_T}) Distributions

It's not a gaussian, it's a gamma distribution

Analytical formula for statistically independent emission (the sum of independent x_i)



$$M_{p_T} = \overline{p_T(n)} = \frac{1}{n} \sum_{i=1}^n p_{T_i} = \frac{1}{n} E_{Tc}$$

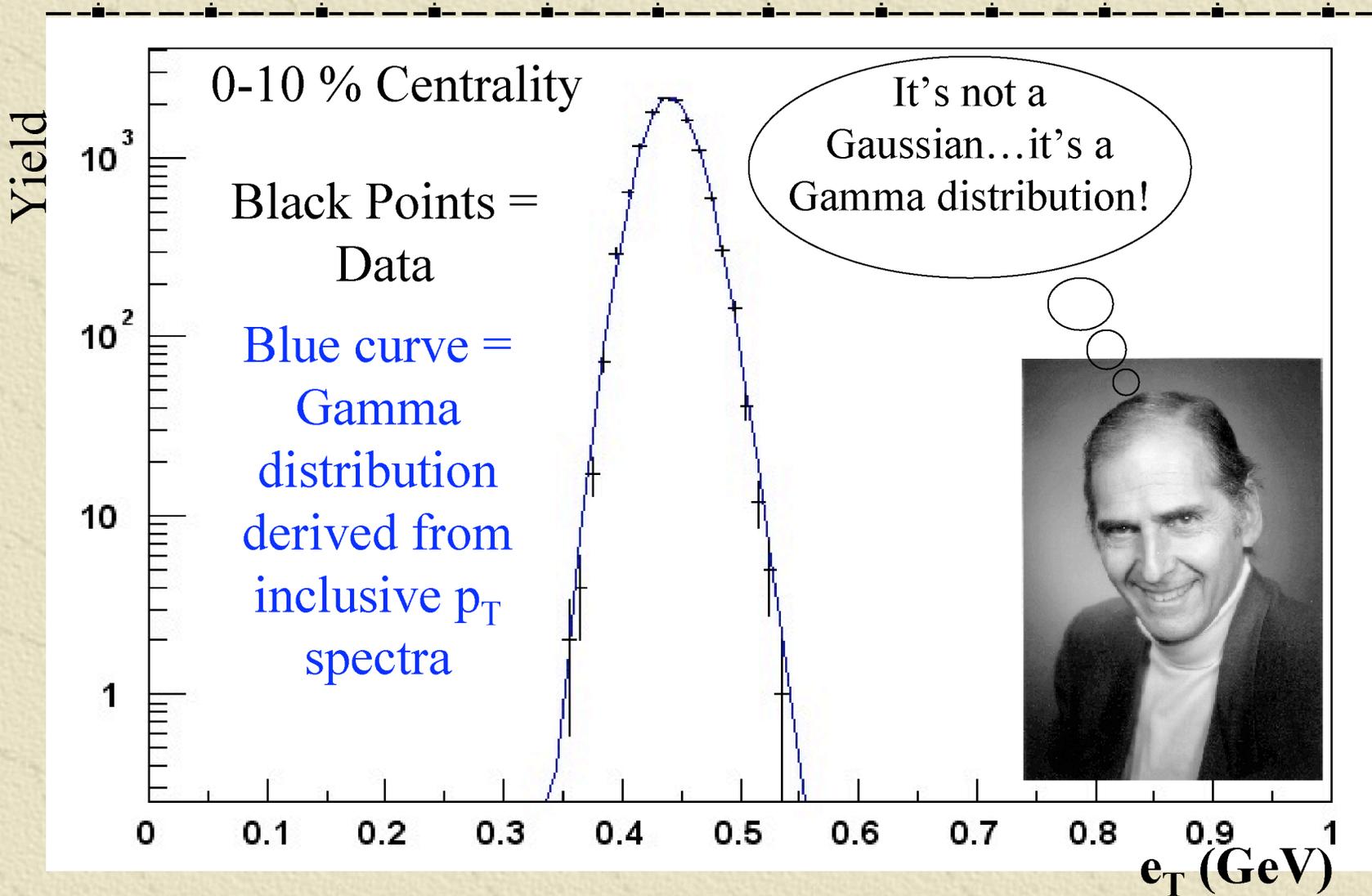
It depends on the 4 semi-inclusive parameters b , p , $\langle n \rangle$, $1/k$ (NBD), which are derived from the quoted means and standard deviations of the semi-inclusive p_T and multiplicity distributions. The result is in excellent agreement with the NA49 Pb+Pb central measurement [PLB 459, 679 \(1999\)](#)

See [M.J.Tannenbaum PLB 498, 29 \(2001\)](#)

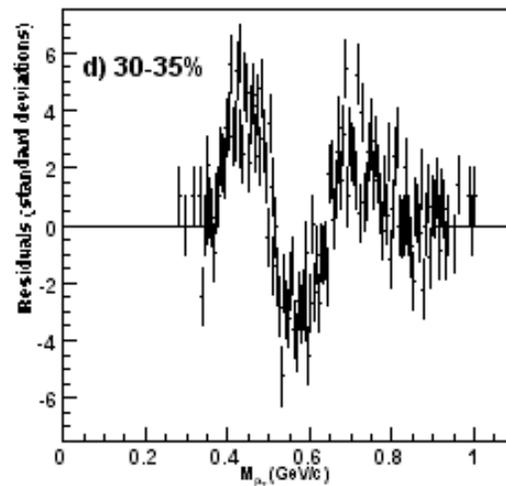
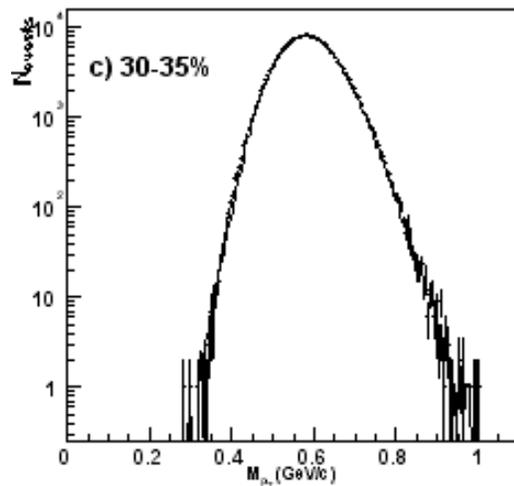
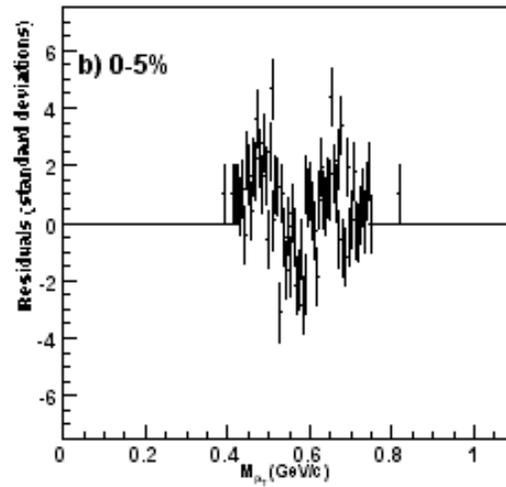
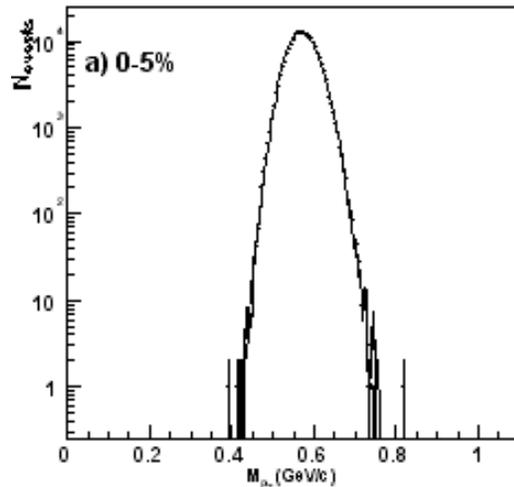
Average e_T Fluctuations

Run-2

Jeffery Mitchell



PHENIX M_{p_T} vs centrality

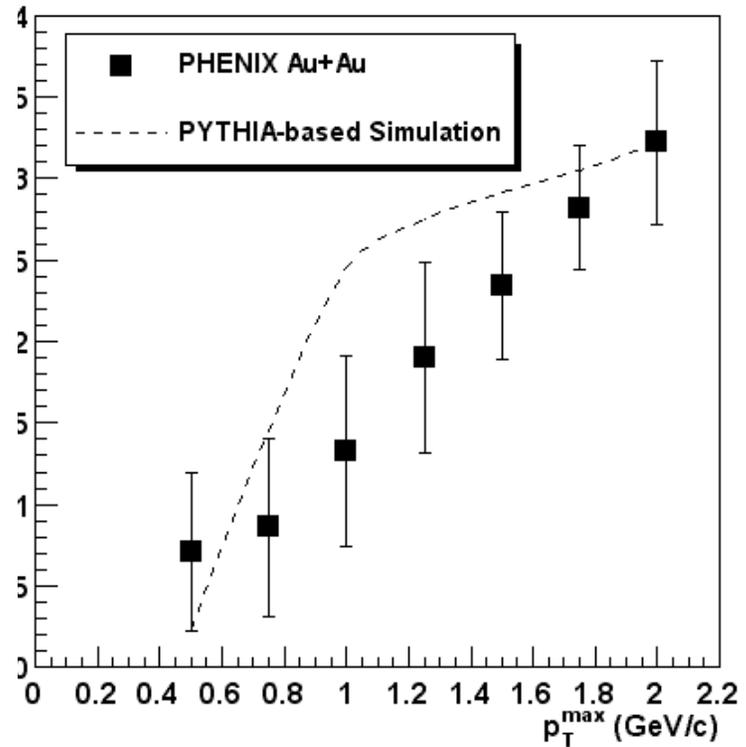
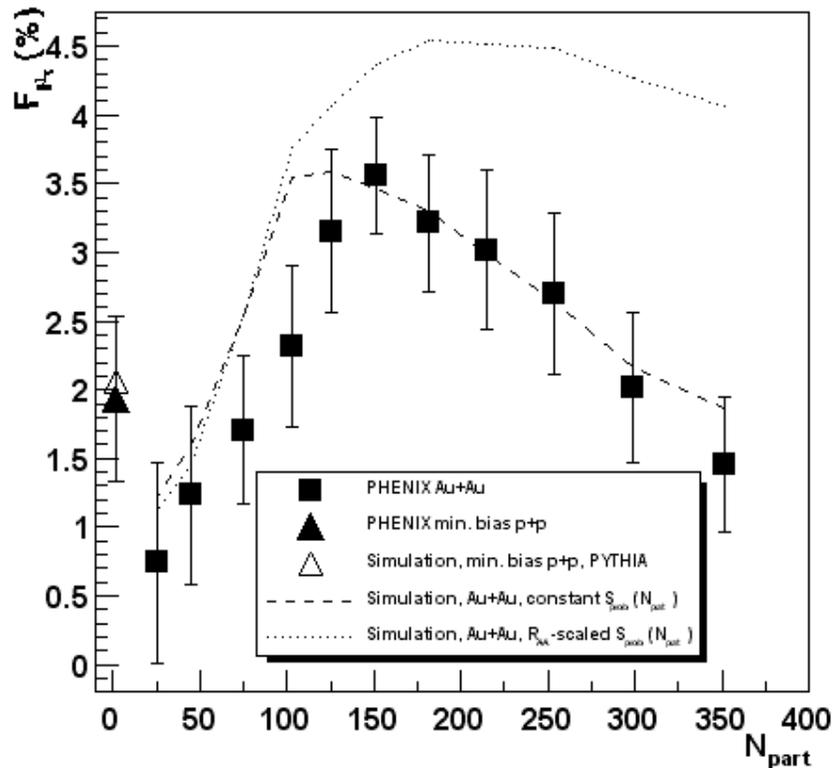


- compare Data to Mixed events for random.
- Must use **exactly** the same n distribution for data and mixed events and **match** inclusive $\langle p_T \rangle$ to $\langle M_{p_T} \rangle$
- best fit of real to mixed is statistically unacceptable
- deviation expressed as:

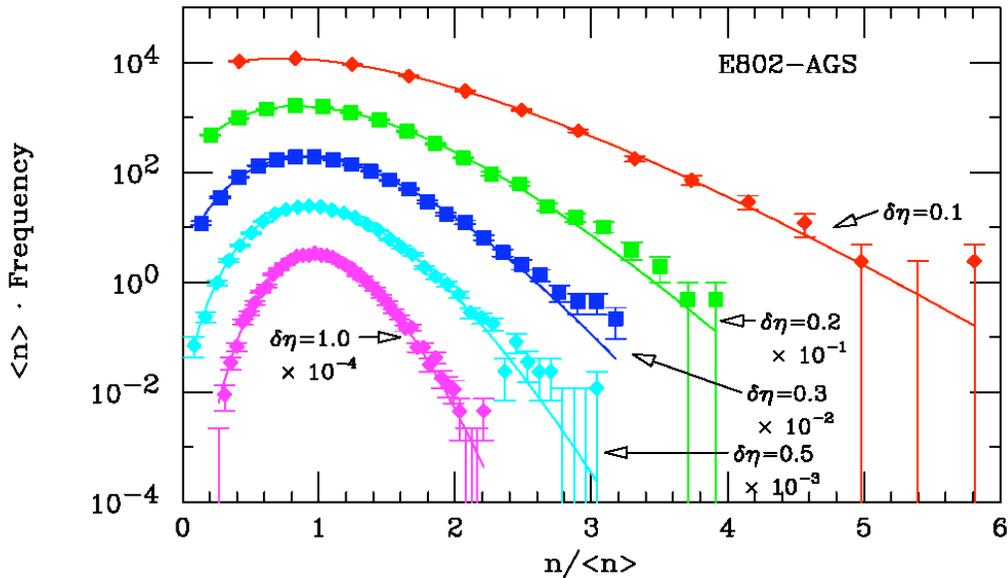
$$F_{p_T} = \frac{\sigma_{M_{p_T} \text{ data}}}{\sigma_{M_{p_T} \text{ mixed}}} - 1$$

Corelation is a few percent of \square_{MpT} : Due to jets

PHENIX nucl-ex/0309xxx



E802 O+Cu Central Multiplicity data in eta bins



Large fluctuations
in AGS O+Cu
central *multiplicity*:
distribution is NBD
(correlations due to
to B-E

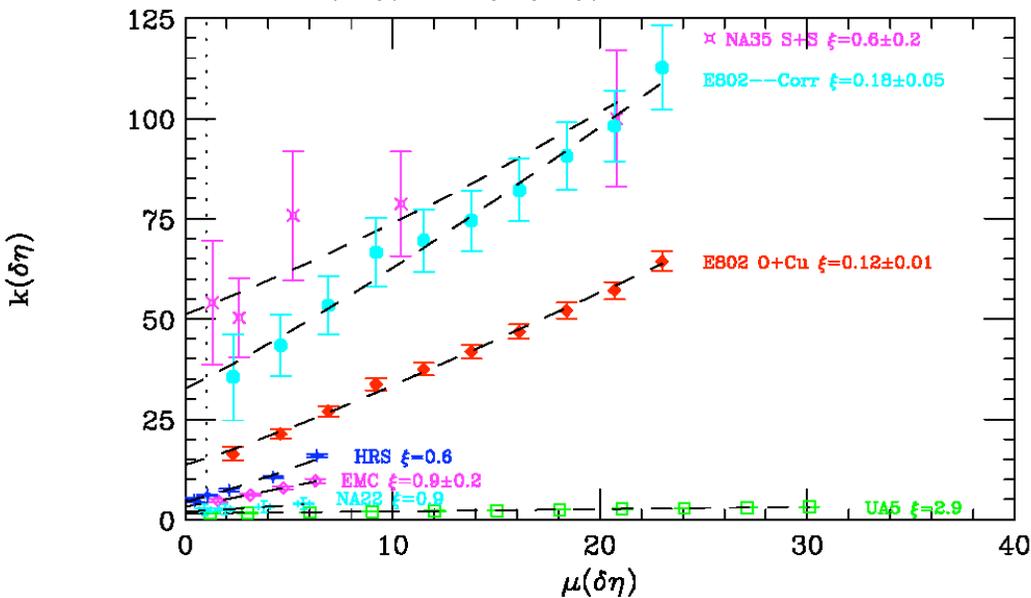
don't vanish)
PRC **52**, 2663 (1995)

$$\sigma^2/\langle n \rangle^2 = 1/\langle n \rangle + 1/k$$

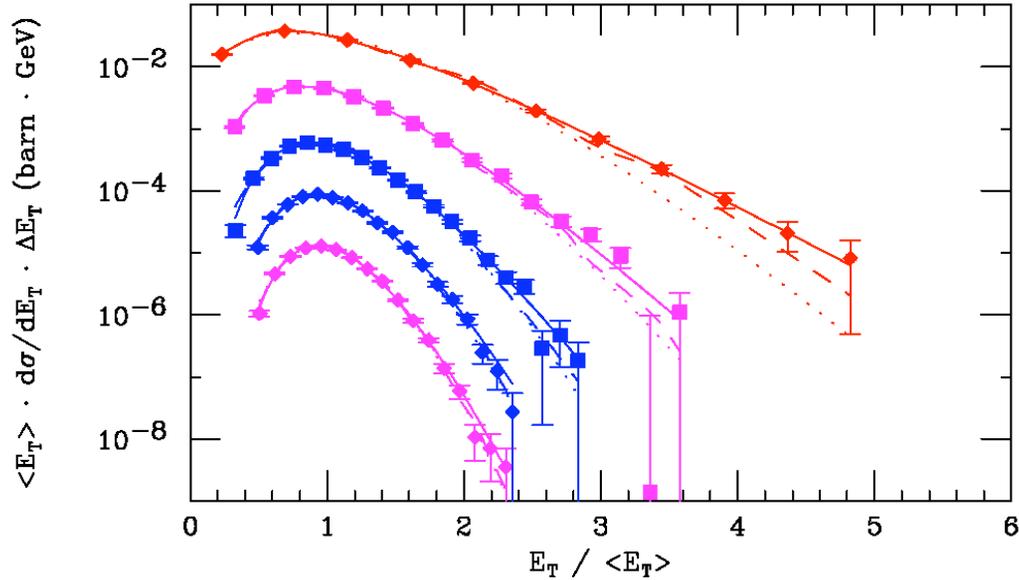
What happens at RHIC?

- Au+Au
- Lighter Ions

$k(\delta\eta)$ vs $\mu(\delta\eta)$ from NBD fits



E802 0+Cu Central E_T data in eta bins



Is E_T or
multiplicity
primary?
Which fluctuates
more?

PRC **63**, 064602 (2001)

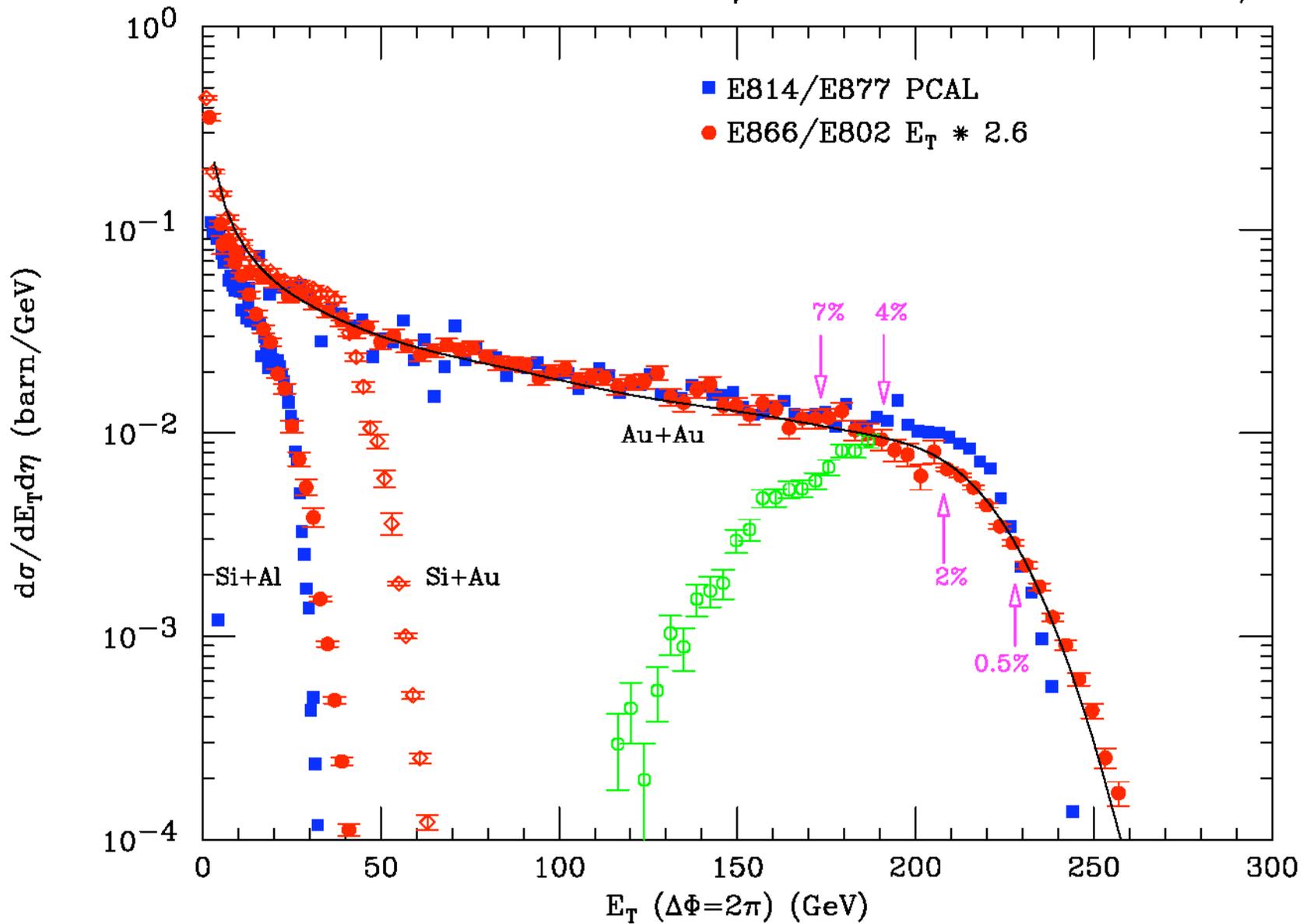
- QCD-like: first create Energy which fragments to multiplicity
- as in Mp_T , create multiplicity which then gets p_T according to inclusive distribution.

- solid lines fit to simple \square distribution
- dotted line mult NBD followed by \square for E_T per particle

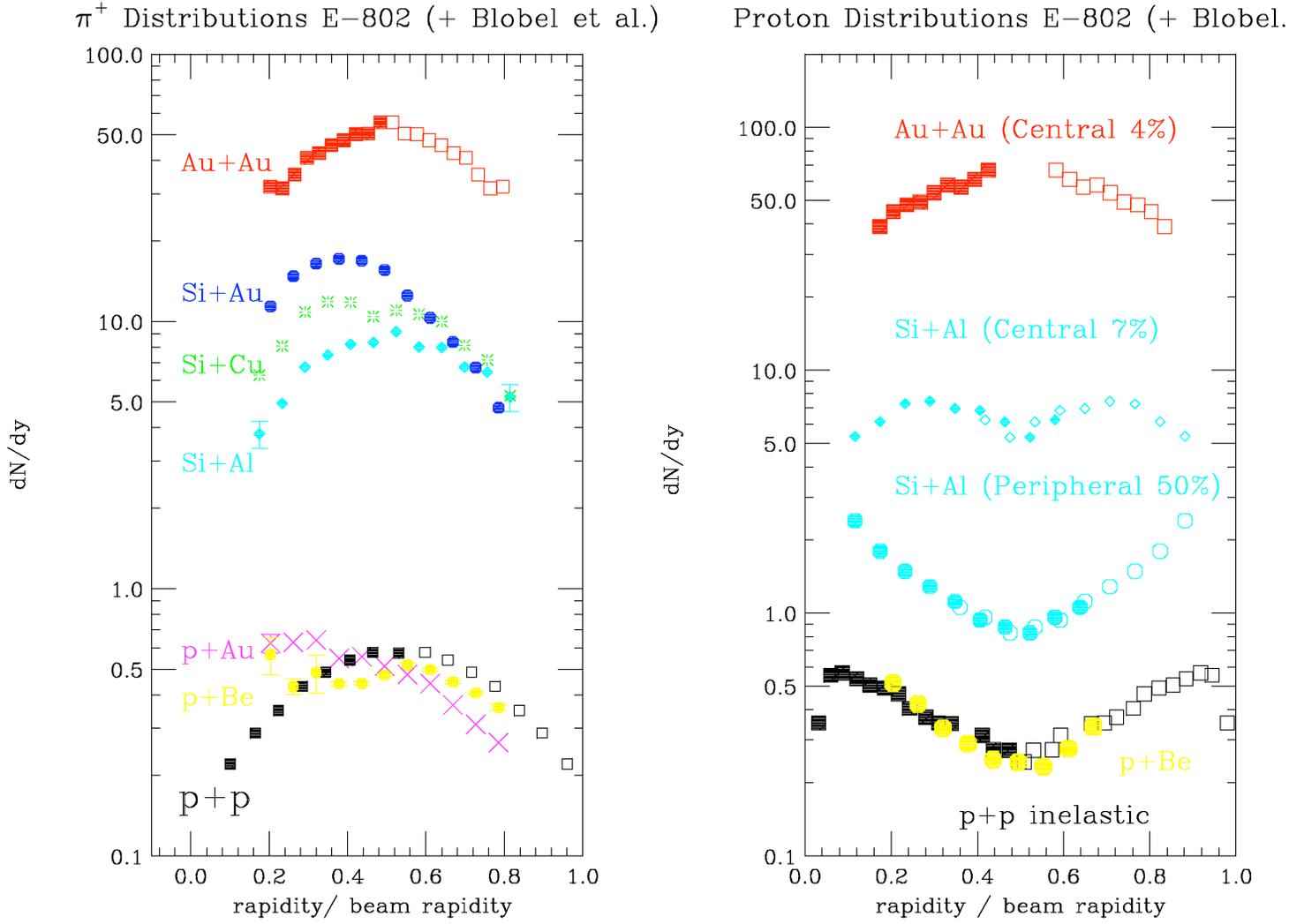
$$\frac{d\sigma}{dE_T} = \sigma \sum_{n=1}^{n_{\max}} f_{\text{NBD}}(n, 1/k, \mu) f_{\Gamma}(E_T, np, b)$$

??

E866 & E877-AGS corr to $\Delta\eta=1$ and $\Delta\Phi=2\pi$ 14.6 A GeV/c



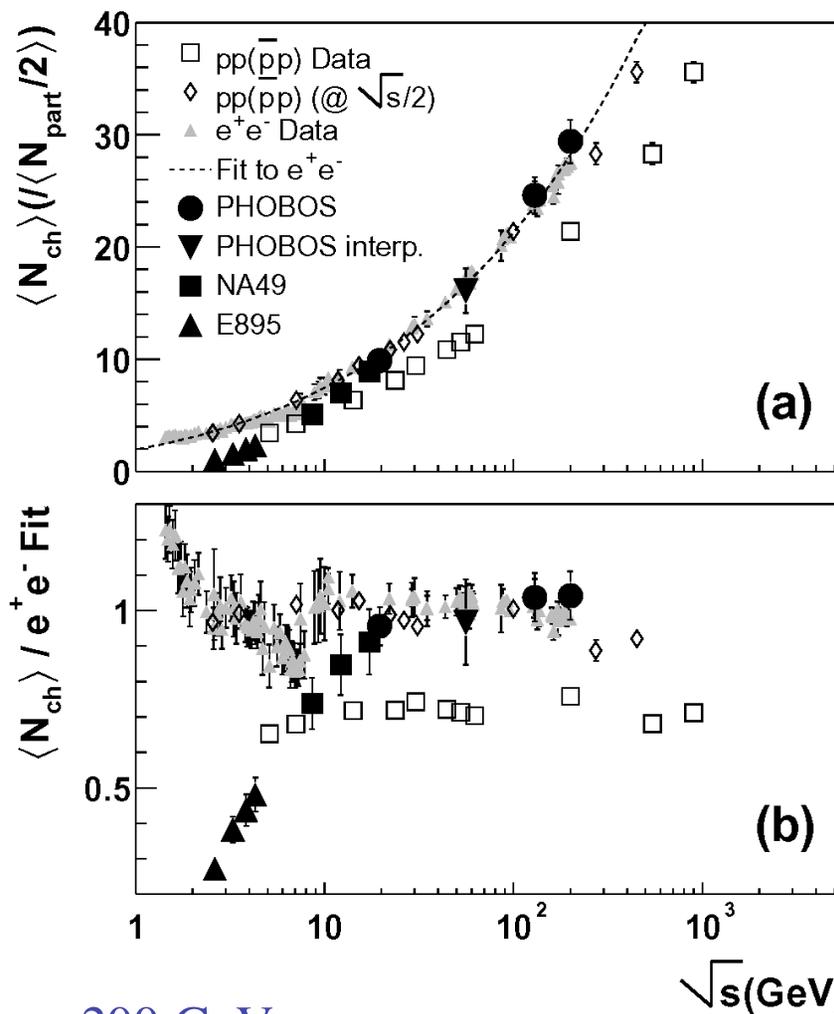
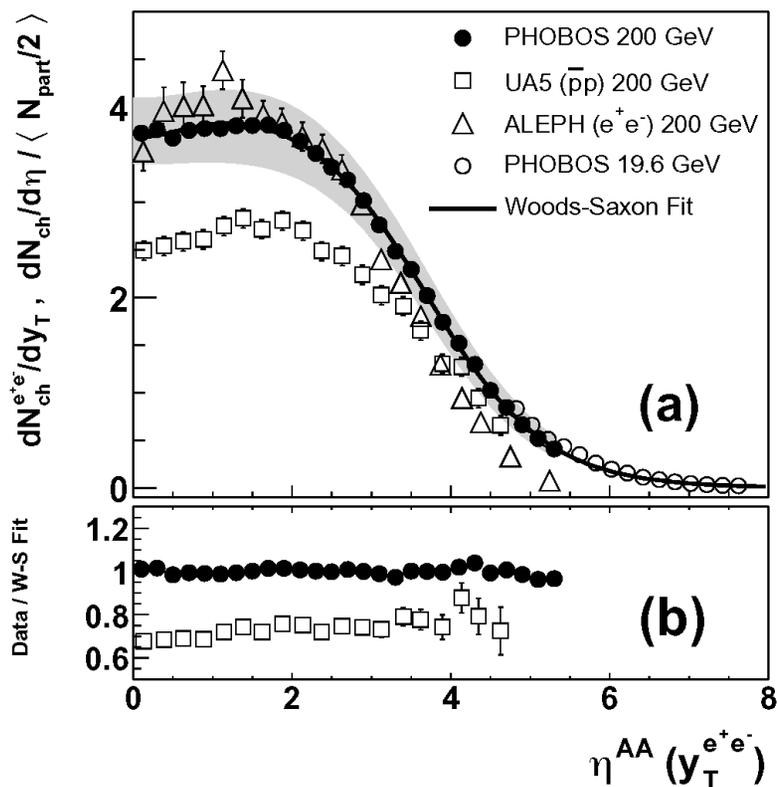
E802 measurements of dn/dy π^+ and p confirm stopping



For Nino

PHOBOS $dn/d\eta$

cf. M.Basile, et al, **PLB92**, 367 (1980);
B95, 311 (1980)



$dn/d\eta / \langle N_{part} / 2 \rangle$ Au+Au $\sim e^+e^- > pp @ \sqrt{s_{NN}}=200$ GeV