Experimental Studies of Particle Production in Ultra-Relativistic Heavy Ion Collisions

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Abstract

This thesis is about high energy heavy ion physics. Two major experiments, the WA98 at CERN and the PHENIX at Brookhaven National Laboratory, are described. Results from data analysis from both experiments are presented.

The WA98 chapter presents particle identification methods and phi-meson measurements. A procedure for calculating efficiency maps for tracking chambers is discussed.

The PHENIX chapter gives a technical overview of the Pad Chamber reconstruction software and ends with an analysis of single particle spectra based on 65+65 A GeV Au+Au data.

The thesis is based on five papers. The first four describe the development of measurement and analysis methods, while the last paper was the first physics paper from the PHENIX experiment and is about charged particle multiplicity.

Key words
Pad chambers, pad readout, PHENIX, WA98, phi-meson, tracking, single particle spectra, heavy-ion collisions, inverse slope parameters, particle production, charged particle multiplicity, RHIC, SPS

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**Populärvetenskaplig sammanfattning
(Popularised summary in Swedish)**

Denna lilla skrift behandlar ett område inom fysiken som kallas högenergetisk tungjonsfysik. Kort förklarat betyder detta att vi studerar kärnmateria under extrema förhållanden, såsom höga temperaturer och höga tryck. För att åstadkomma dessa, kolliderar vi bly- och guldkärnor med mycket höga energier vid partikelacceleratorer, i USA (New York) och i Schweiz (Genève). När kärnorna kolliderar smälter dessa till en soppa av sina innersta beståndsdelar, kvarkar och gluoner. Är energin tillräckligt hög hoppas vi kunna skapa ett s.k. kvark-gluon plasma, där kvarkarna och gluonerna uppträder som fria partiklar, om än för en kort stund; en miljarddels miljarddels sekund eller så. Man säger att en fasövergång har ägt rum, ungefär som då kokande vatten övergår i vattenånga. Liknande när vattenångan kondenserar tillbaka till flytande form då den kyls av, kommer kvark-gluon plasmat att kondensera till kärnpartiklar, såsom protoner och neutroner men även andra mer exotiska partiklar som pioner och kaoner. Detta studium visar sig även vara viktigt för vår förståelse av universums tidiga utveckling. Fram till en miljondels sekund efter Big Bang (den stora smällen), bestod universum av ett kvark-gluon plasma som snabbt expanderade. I våra experiment försöker vi återskapa de förhållanden som rådde i universum när det var ungt. För att göra detta har vi konstruerat detektorer, stora som flervåningshus, för att kunna registrera de ytterst små partiklarna. Denna avhandling beskriver bl.a. utvecklingen av mätinstrument, metoder, och hur informationen som kommer från detektorn tas om hand när den används i experimentet. Under experimentets gång genereras gigantiska datamängder, omkring 20MB i sekunden, som sedan analyseras. Avhandlingen innehåller även analys av sådan data, där en uppskattning görs av partikeltätheten och temperaturen hos kärnmaterien skapad i guld-guldkollisioner, vid tio gånger högre energi än tidigare.

Avhandlingen är indelad i fem kapitel, utöver de fem artiklar som avhandlingsarbetet har resulterat i. Första kapitlet är en introduktion till ämnesområdet riktat till icke experter. Den käre läsaren ledsagas genom universums historia som vi känner den, till hur vi studerar Big Bang i laboratoriet. De senare delarna är av mer teknisk natur, där andra kapitlet behandlar WA98 experimentet som kördes på CERN i Schweiz under flera år fram till senhösten 1996. Tredje kapitlet avhandlar PHENIX experimentet på Brookhaven National Laboratory i USA, medan det fjärde innehåller en sammanställning av vad vi har lärt oss vad gäller de mer tekniska delarna av avhandlingen. Sista kapitlet är en kort sammanfattning av artiklarna som återfinns i sin helhet i appendix I-V.

Pål Nilsson
Lund, 27:e augusti, 2001.

This thesis is based on the following publications:

I **A large-acceptance spectrometer for tracking in a high multiplicity environment, based on space point measurements and high resolution time-of-flight**
Nucl. Instr. and Methods **A** *431 (1999) 123-133.*
L. Carlén et al.

II **The pixel readout system for the PHENIX pad chambers**
Nucl. Phys. **A** *661 (1999) 665c-668c.*
P. Nilsson.

III **Event reconstruction in the PHENIX central arm spectrometers**
Accepted for publication in Nucl. Instr. and Methods.
J. T. Mitchell et al.

IV **Construction and performance of the PHENIX pad chambers**
To be submitted to Nucl. Instr. and Methods.
K. Adcox et al.

V **Centrality dependence of charged particle multiplicity in Au+Au collisions at $\sqrt{s_{NN}}$ = 130 A GeV**
Phys. Rev. Lett. **86** *(2001) 3500-3505.*
K. Adcox et al.

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[1] Brookhaven National Laboratory, New York, USA.
[2] Vanderbilt University, Tennessee, USA.
[3] See http://fysik.lu.se and http://fysikilund.lu.se

To my mother and
the memory of my father

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Chapter 1

Introduction

"If it were possible to experiment with neutrons or protons of energies above a hundred million volts, several charged or uncharged particles would eventually leave the nucleus or as a result of the encounter; with particles of energies of about a thousand of million of volts, we must even be prepared for the collision to lead to an explosion of the whole nucleus."

- Niels Bohr, Nature 137(1936)351.

What once was only dreamed about, in the minds of giants, is now possible at several facilities around the world. At CERN in Switzerland and at Brookhaven National Laboratory (BNL) in the United States, nuclei from heavy atoms are accelerated to extremely high energies colliding with a fixed target or to a beam of nuclei coming from the opposite direction. When two such ions collide, literally thousands of particles are created, flying out of the collision zone together with the remnants of the incoming nuclei.

As it turns out, these studies are important for understanding the origin of our universe. The small can be seen in the large, and the large in the small. In the following pages, two major experiments are described that were built for exploring the universe within the atom.

1.1 The history of the universe

The creation of the universe took place some 12-15 billion (10^9) years ago. At the very first instant of the newly born universe, it is believed that all forces of nature were united in a single force. After 10^{-43} seconds, the gravitational force was suddenly decoupled, space-time became four-dimensional and the era of super unification ended. The era of grand unification began. The physics of this epoch is far from certain, but is believed to be described by a grand unified theory. The mediators of this force are the X, $\overline{X}$, Y and $\overline{Y}$ particles. Together with quarks, leptons and exchange particles they populated the universe. At the end of this epoch, the X and Y, as well as their anti-particles, disappeared from the universe. As they decayed, they left behind a small excess of quarks over anti-quarks with the ratio of $(10^9 + 1) : 10^9$.

When the universe was about 10^{-34} seconds old, it was blown up well beyond the observable edge of the universe. Despite this enormous expansion, the epoch of inflation was very short. At 10^{-33} seconds, it was over and the expansion settled into its "normal" rate. It has, however, recently been discovered from studies of distant supernovae that the expansion is slowly accelerating [1, 2, 3], implying a possible existence of a repulsive component in the gravitational force.

In the primordial fireball, the average energy per particle exceeded 10^{19} GeV [1] and most of it was in the form of radiation. As the universe continued to expand, it gradually cooled off.

Up until 10^{-10} seconds, when the average energy had sunk to around 100 GeV, the particle soup also contained W and Z-bosons in thermal equilibrium. Quarks and anti-quarks were too close to each other for the strong force to bind them into hadrons. Any system that was temporarily bound together, would immediately be

[1] Being the original seed of our universe, a quantum bubble the size of Planck length dimensions, $l_P = ct_P = \sqrt{\frac{\hbar G}{c^3}}$, had an energy density of $\rho_P = \frac{m_P}{l_P^3}$ where the Planck mass is given by $m_P = \sqrt{\frac{\hbar c}{G}}$. The energy can then be calculated from $E = m_P c^2 = \rho_P l_P^3 c^2$.

blasted apart by the high energy particles that also were present. Also mixed up in this soup were the mediators of the strong force, the gluons, which together with the quarks are often referred to as partons.

The free state of partons, called a *quark-gluon plasma* (QGP), lasted until the first few microseconds had passed [4]. At this time it is believed that a phase transition took place which imprisoned the once free quarks and gluons into the particles we now call hadrons. In particular, protons and neutrons were formed.

As the temperature dropped below the threshold for proton and neutron production, virtually all anti-matter was removed from the universe leaving a small excess of matter behind. After the annihilation, there was one proton for every 10^9 photons. The surviving protons and neutrons were kept in thermal equilibrium with the rest of the universe via neutrino interactions. This lasted until one second after the Big Bang when the density became too low for maintaining the equilibrium. The neutrinos decoupled from the rest of the universe which, for the neutrinos, was transformed into a thin fog, barely observable.

When the universe was three seconds old, the remaining anti-matter in the form of positrons was removed by the excess of electrons. Three minutes later, the nucleosynthesis started and the first compound nuclei, deuterons, were formed by protons and neutrons. The deuterons quickly formed the helium nuclei ^{3}He and ^{4}He. Nucleosynthesis was halted some 13 minutes after the Big Bang, when the temperature had dropped below that required for nuclear fusion reactions.

After some 300 000 years, the average photon energy was not sufficient any more to prevent electrons from being captured by protons. The first atoms were formed. When this happened, the large number of free electrical charges disappeared. The photons had nothing to interact with and thus decoupled from matter. The universe became transparent. At approximately the same time, the radiation density dropped below the density of matter. The radiation era was over and the matter era began.

The first galaxies were not born until a billion years had passed.

The actual dynamics are not known. Huge clouds of gas and dust contracted to form stars, in which nucleosynthesis restarted and produced heavier elements. Solar winds swept through the local systems and removed most of the debris. Smaller lumps of matter became planets on which life itself was finally formed. From a simple molecule in a puddle of water, plants, fish and complex beasts with growing intelligence evolved. Eventually man was created, declared himself the crown of creation and began to speculate about his origin.

At this point in the history of the universe we are capable of reconstructing the conditions of the early universe, although on a much more modest scale. This is where our story begins.

1.2 The Big Bang in the laboratory

Looking deeply into matter, require probes of high energy. Until the 1950's, the only source of such probes was cosmic rays. These can be extremely energetic, but since the rate is low it makes statistical analysis difficult. Today, the overwhelming majority of experiments use large machines, known as accelerators, to produce beams of particles or nuclei.

1.2.1 Accelerators

There are different kinds of accelerators in use today. Relevant for this thesis are the fixed-target and the colliding-beam machines. In a fixed-target machine, the particles or ions are accelerated to the desired operating energy after which the beam is extracted from the machine and steered on to a stationary target. In colliding-beam accelerators, two beams coming from opposing directions are made to collide with each other. An advantage of using a collider is that all of the beam energy is available for particle production, in contrast to a single beam accelerator where a large fraction of the

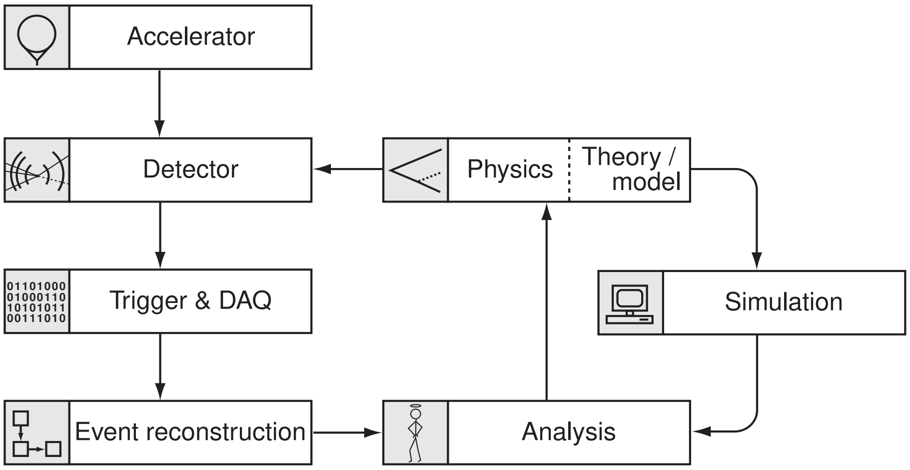

Figure 1.1: *The various steps of a physics experiment.*

energy is consumed by moving the center-of-mass frame. A disadvantage is however that the collision rate in a heavy ion collider is usually smaller than can be achieved with a fixed target.

An example of an accelerator for fixed-target experiments is the Super Proton Synchrotron (SPS), at CERN. It started as a proton accelerator but was later upgraded for acceleration of the nuclei ^{16}O and ^{32}S, with center-of-mass energy $\sqrt{s} = 19.4$ A GeV, and ^{208}Pb with $\sqrt{s} = 17.4$ A GeV.

In the summer of 2000, the Relativistic Heavy Ion Collider (RHIC) at BNL started to deliver ^{197}Au beams. The beams can each be accelerated up to 100 A GeV, which gives a center-of-mass energy of 200 A GeV. It was initially run at a beam energy of 65 A GeV.

With a startup in 2007, the Large Hadron Collider (LHC) at CERN will deliver Pb beams with a maximum center-of-mass energy of about 5 500 A GeV.

1.2.2 Experiments and detectors

The collisions leading to particle production take place on several stations (experiments) around the accelerator. In order to deduce details of the produced particles and eventually gain further knowledge of the laws of nature, each experiment has a set of devices known as detectors placed around a collision region. The detectors contain volumes of certain materials that are easily ionized by the traversing particles. The ionization is recorded by amplifying electronics for later hit reconstruction and analysis.

In a collision we have to track and identify the out coming debris as well as to understand exactly how it was created. For this purpose, very advanced detector systems have to be utilized. Two such systems, the WA98 and the PHENIX experiments, will be discussed in this thesis. Common to both of these experiments is the desire to study nuclear matter under extreme conditions such as high temperature and high pressure.

1.2.3 Experimental challenges

Heavy ion collisions present additional experimental difficulties compared to *pp*-collisions since the particle multiplicities are several orders of magnitude higher. Any detection aiming at observing individual particles must increase the granularity of the detector accordingly. Many of the interactions of the particles with the detector material have very small spatial extension and the appropriate granularity can be achieved by making the readout more granular. This is the case for e.g. gaseous ionization detectors which we deal with in this thesis. If the physical dimension is large, e.g. an electromagnetic shower, the granularity can only be achieved by moving the detectors out to a suitable distance (like the LEDA in WA98, see chapter 2) and increase the number of channels.

1.2.4 Detector construction

Although there are many different kinds of detectors, most of them are based on the principle of ionization. An obvious choice of a suitable detector medium is a gas, because of its great mobility of electrons and ions. When a charged particle traverses a gaseous volume it will ionize the gas. If an electric field is established, it will accelerate the positively charged gas atoms towards the cathode plane and the electrons towards the positive anode. If the electric field is strong, the accelerated electrons can ionize gas atoms resulting in an avalanche process which gives an amplification.

In a multi-wire proportional chamber (MWPC), the anode consists of many thin and closely spaced wires. The strong field is created by the fact that $E \sim \frac{1}{r}$, resulting in a very high E in the vicinity of the wire. It is possible to either read out the wires directly, or to divide the cathode plane into small electrodes (pads) and read out these instead. This has several advantages over wire readout; the most important being the removal of ambiguous signals caused by more than one particle passing close to the same wire.

The Pad Chambers in the PHENIX experiments are MWPC:s read out on finely segmented pad planes. It has partly been the responsibility of the Lund group to develop and fund these chambers, as well as their readout systems and analysis tools.

In the Multi-Step Avalanche Chamber (MSAC), used in the WA98 experiment, a homogeneous E-field is created by successive parallel meshes. The cloud of electrons liberated by the avalanches can be collected on the anode side of the chamber.

Thorough details about the PHENIX Pad Chamber construction can be found in Paper IV in this thesis, while the MSAC:s in the WA98 experiment are described in paper I.

1.2.5 Electronic readout

As mentioned, the granularity problem in gas chambers is reduced to finding a suitable electronic readout. In a typical central collision at RHIC, thousands of particles traverse the detector system. This generates a tremendous amount of data from the sub system electronics that need to be taken care of before the next collision takes place on the average a few hundred microseconds later. Events are buffered to cope with fluctuations in the event rate. In the PHENIX experiment, the data rate at full luminosity is about 20 MB/s. In the future ALICE experiment at the LHC, the rate is expected to be about 1.2 GB/s. Since the total time of data collection is on the order of several years, this puts extreme demands on storage systems and processors, not to mention the programmers that need to write fast code.

Due to the immense number of particles traversing the chambers, the readout electronics must have many readout channels in order to resolve the tracks. Since the signals from the sensing electrodes are very small and thus won't survive undisturbed any longer trace distances, it is often a requirement that the readout electronics are mounted on the chamber, in the path of the traversing particles. This in turn means that the electronics must be made very thin, in order not to disturb the particles passing through.

The many channels can potentially be very expensive, both in cost and power consumption. The concepts presented in this thesis (papers II and IV) describe a readout system that reduces this cost drastically. By introducing a simple readout channel, close to the pad that has a complex structure made to reduce the number of channels, we achieve low cost, high performance, low power and minimal data volume. The minimal material thickness is achieved by realizing the front-end amplifying and digitizing electronics in specially designed integrated circuits mounted with the chip-on-board technique on ultra-thin printed circuit boards.

1.2.6 Event reconstruction

The raw data coming from the detectors, or sub systems, is orders of magnitudes larger than the numbers given above. Most of this data is not worth saving and is thus removed by the hardware in the

Data Acquisition system (DAQ). Special triggers select the events that are considered to contain interesting data. The level-one triggers select events that are considered to be real events not triggered by cosmic radiation or other background signals, while the level-two trigger algorithms have been written to select special events, such as coherent peripheral collisions, events containing a detected J/ψ candidate, etc.

In the PHENIX experiment, the event data is stored in PRDF (PHENIX Raw Data Format) files. The selected events are processed by the reconstruction algorithms. These are described in detail for the WA98 and PHENIX experiments in papers I and III in this thesis. The reconstructed events are stored in DST (Data Summary Tape) files for later analysis. A typical DST file is quite large and usually contain more information than necessary for a certain analysis. They are thus often compressed into much smaller μDST files that are more easy to analyze, e.g. in the PAW [5] or ROOT [6] environments.

Analysis will eventually lead to physics results that either can be understood with existing theoretical models or to improve them. Theory and models form the basis of simulation. The entire experiment can be simulated by software using input from event generators, such as Fritiof [7], RQMD [8], HIJING [9], NEXUS [10] etc. The GEANT tool [11] was developed for this purpose and is heavily used in the particle physics community, but also in other areas of science. With this tool, the detectors are first modeled in 3D. Functions are then added to simulate the response of the modeled detectors. When the simulation is run, the input data from an event generator, i.e. the generated particles, are stepped through the detector system. The output file is ideally of the same format as real data in order to read it seamlessly into the same analysis package.

1.2.7 Analysis

Before the physics analysis can begin, enough statistics (data) has to be collected. In addition to this, the response of the detector subsystems must be understood and the cuts that should be applied to the data in order to remove background etc need to be investigated. Normally, a detector system can not record all emitted particles. This limitation, called the detector acceptance, can be due to geometrical coverage and/or limitations in momentum outside of which particles may not enter the detector (low momentum) or may not be identified (high momentum). When these things are understood by simulation, the corrections can be determined, the raw spectra can be produced and the physics analysis can start.

1.3 The quark-gluon plasma

If hadrons were subjected to the same energy densities as the state of the universe during the first microsecond, one would recreate the conditions of the early universe in a system of nuclear size and generate a phase transition back to the free parton state. Assuming that this is indeed possible, how would we know that a phase transition really took place? It is generally recognized that there is no single unique signal to look for that guarantees that a QGP has been formed. What can be achieved may be a set of signals that, when taken together, may indicate the presence of the deconfined state. A few such probes will now be discussed briefly.

1.3.1 Thermodynamical variables

A phase transition can be classified according to the discontinuity of the free energy F as a function of the temperature T [12]. The phase transition is said to be of the nth order if the nth partial derivate $\partial F/\partial T$ is discontinuous while all the lower derivatives in

T are continuous. Since the free energy can be written

$$F = E - TS \tag{1.1}$$

where E is the internal (total) energy and S is the entropy, we have that

$$S = -(\frac{\partial F}{\partial T})_E \tag{1.2}$$

According to the definition above, a first order phase transition is thus characterized by a discontinuous entropy and a continuous free energy as a function of temperature. An example of a first order phase transition is the liquid to gas process of boiling water.

In a system of colliding nuclei, the particle density ($dN/d\eta$) can be related to the entropy [13]. Furthermore, it is usually assumed that a measure of the system temperature can be obtained from the $< p_T >$ or the exponential slope of the p_T-distribution of the produced particles. During the collision process, the system will be heated until it reaches a critical temperature, T_c. At this point the system remains in a mixed state at a constant temperature until the phase transition is complete. A first order phase transition should therefore be seen in a $< p_T >$ versus $dN/d\eta$ plot as a plateau and a rise. T_c could then be determined from the $< p_T >$ at the level of the plateau.

Another interesting variable is the energy density, ϵ, which can be derived from Eq. 1.1 above by dividing the total energy with the volume,

$$\epsilon = \frac{E}{V} = \frac{F + TS}{V} \tag{1.3}$$

It is experimentally most commonly estimated from the Bjorken formula [14]

$$\epsilon_{Bj} = \frac{1}{\pi R_0^2 c\tau_0} \frac{dE_T}{d\eta} \tag{1.4}$$

where E_T is the transverse energy, R_0 is the transverse radius of the system and τ_0 is the formation time, usually taken as 1 fm/c.

1.3.2 Strangeness enhancement

As predicted by Rafelski et al. [15] and later confirmed by lattice QCD calculations [16], the strangeness content in a quark-gluon plasma in thermal and chemical equilibrium is believed to be enhanced from that of normal hadronic matter. The strange quark production is expected to be enhanced since the up and down quarks are Pauli blocked due to their high concentration and to the high Fermi energy they would have to overcome. The creation of $s\overline{s}$ pairs might therefore be energetically favourable over the creation of $u\overline{u}$ or $d\overline{d}$ pairs, if the available energy is above $2m_s$c^2. In a gluon rich plasma, the strange quark pairs will predominantly be produced by gluon-gluon interactions. The estimated equilibration time is of the order of a few fm/c, i.e. of the same order as the interaction time of a high energy heavy ion collision.

An enhancement in the strangeness production can also happen in a hot and dense hadronic gas. The relative abundance of strange particles in that scenario could gradually grow in a chain of rescattering processes. In this case however, the characteristic time scale is on the order of 100 fm/c, i.e. much longer.

It should be pointed out that any observed strangeness, whether it originates from a QGP or not, is a powerful probe of the reaction dynamics as there is no strangeness in the initial state. All strange hadrons are created in the interaction.

The enhancement should be more pronounced for multi-strange hadrons, e.g. ϕ-mesons ($s\overline{s}$) and Ω-baryons (sss), than for singly strange hadrons, like the K-mesons ($K^+ = u\overline{s}$, $K^0 = d\overline{s}$, $\overline{K_0} = \overline{d}s$, $K^- = \overline{u}s$). To avoid ambiguities, it would be even better to study anti-hyperons (e.g. $\overline{\Lambda}$) since they are not likely to be produced by hadronic rescattering.

Several experiments at the AGS (BNL) and SPS (CERN) have observed strangeness enhancement. At AGS energies (14.6 A GeV) it was first observed by the E802 experiment in Si-induced central collisions as an increase in the K^+/π^+ ratio [17]. At 158 A GeV/c $Pb + Pb$ collisions at the SPS, WA97 has reported a large excess for particles of higher strangeness content, as compared to $p + A$. For $\Omega^- + \overline{\Omega}^+$ the factor is about 20 [18].

1.3.3 Light vector mesons

It has been suggested (e.g. in [19]) that the mass, width, branching ratios and yields of light vector mesons (such as ρ, ω and ϕ) differ from normal nuclear matter when produced in hot, dense and strongly interacting matter. E.g. the branching ratios between

$$\phi \to \begin{cases} K^+ + K^- & (49.1 \pm 0.8)\% \\ e^+ + e^- & (2.99 \pm 0.08) \times 10^{-4}\% \end{cases} \tag{1.5}$$

are very sensitive to mass changes. If chiral symmetry is restored, which it is believed to be inside a plasma [20, 21], the quark masses go to zero. Since the mass of the ϕ-meson is very close to twice the kaon mass, $m_\phi - 2m_K = 32$ MeV/c^2, and the kaon channel is the dominant decay mode, the ϕ decay will be highly susceptible to changes in the strange quark mass. The kaon decay mode would be prohibited and the branching ratios between the kaon and electron channel be modified.

The PHENIX experiment at RHIC will measure both of these channels, while the NA49 [22] and WA98 experiments (see section 2.5) at the CERN SPS have measured the $K^+ + K^-$ channel.

1.3.4 Direct photons

Photons created in the collision do not interact strongly with the plasma and can thus escape from the hot zone carrying information about the early stages of the interaction. Since the production rate and momentum distribution of these direct photons, also known as thermal photons, are determined by the momentum distributions of the quarks and gluons, they can give information about the thermodynamical state of the environment in which they are produced. However, it is not a trivial task to distinguish the direct photons from the photons coming from other sources such as $\pi^+ + \pi^- \rightarrow \gamma + \rho_0$. A copious source of photons come from hadronic decay, e.g. abundantly produced neutral mesons that decay into photon pairs forming an experimental background.

The WA98 experiment has published results constituting the first observation of direct photons in ultra-relativistic heavy-ion collisions [23].

1.3.5 J/ψ suppression

The J/ψ particle is the bound state of a charm and anti-charm quark. It has been suggested that the production of this resonance will be suppressed in a QGP due to Debye screening of the color charges [24]. It will thus be more probable for the $c(\overline{c})$ to combine with one of the more abundant $\overline{u}(u)$ or $\overline{d}(d)$ quarks in the hadronization process to form a D-meson rather than a $\overline{c}(c)$ to form a J/ψ.

A suppression has indeed been observed when going from lighter $(p + U)$ to heavier $(S + U)$ systems [25]. However, it has been shown that this is consistent with suppression due to absorption in a dense medium [26]. In heavier systems still $(Pb + Pb)$, the suppression has been found to be much stronger than what can be expected from an extrapolation from the lighter systems [27, 28, 29]. It is as of yet not clear whether this is due to the formation of a QGP or if it could be explained by other mechanisms.

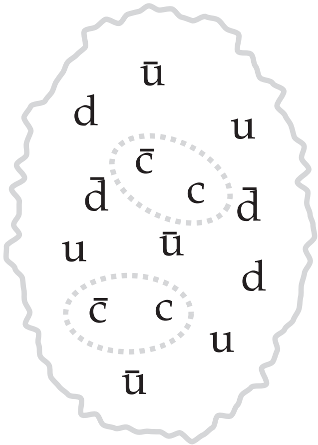

Figure 1.2: *Debye screening in a plasma.*

1.3.6 Neutron stars

Quark matter may not only be formed in particle physics facilities on earth, but also in the far reaches of space, deep inside the remnants of old and once massive stars.

It is a known fact that rotating neutron stars slow down with time due to interactions between the electromagnetic field and the surrounding matter. When the centrifugal forces become smaller, the pressure and density in the core region increase, and a phase transition of the first order may occur in the center. Glendenning et al. [30] found that a drastic softening of the equation of state, e.g. by a phase transition to a QGP, can lead to a rapid contraction of the neutron star at a critical angular velocity and that this will show up as a back bending moment of inertia as a function of frequency. As the star slows down, it will become unstable. The sudden decrease of inertia due to the dramatical changes in the core would create a large glitch which in principle could be observed by careful monitoring of young pulsars.

Chapter 2

The WA98 Experiment

This CERN SPS based, fixed-target experiment, was a large acceptance spectrometer for measurements of photons and hadrons in $^{208}Pb+^{208}Pb$ collisions at 158 A GeV. The aim was to study highly excited and compressed matter, and to search for possible quark-gluon plasma signals predominantly by the direct photon signals. The WA98 experiment ended its data taking in November 1996.

This chapter presents the experiment, a procedure for calculating efficiency maps for tracking chambers, as well as particle identification methods and ϕ-meson measurements.

2.1 Physics objectives

Since several of the suggested quark-gluon plasma signals were covered, the strategy was to measure both hadrons and photons and to study their correlations. Observables studied in this experiment are: charged particle and photon multiplicities, thermal photon emission, energy flow, single particle spectra, charged particle correlations and the production of hadronic resonances such as the Δ^{++} and the ϕ-meson.

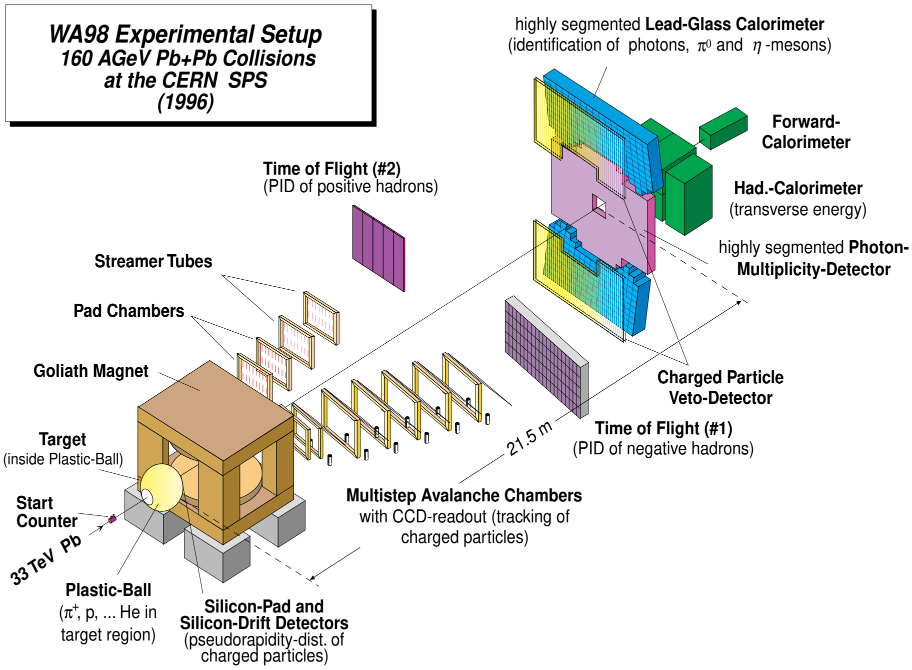

Figure 2.1: *The WA98 experimental setup for the 1996 Pb run.*

2.2 Experimental setup

The Beam Counter and veto detectors. The *Pb* beam provided by the CERN SPS, enters the experimental setup from the lower left in Fig. 2.1. Before it hits the target, it passes the nitrogen gas Čerenkov beam counter which provides a fast start signal for the time-of-flight measurements discussed below. The time resolution is about 30 ps. At a distance of 5.7 m upstream of the target, a halo detector was placed to reject interactions before the beam counter which in turn reject beam particles with charges different from that of the lead nucleus. At a distance of 2.7 m upstreams, a veto detector was placed to limit the beam profile for accepted collisions. A clean beam trigger was obtained by the beam counter in anti-coincidence with the inner halo and the veto detector.

The Plastic Ball. This detector system identified charged particles such as π^+, protons and light ions and was mainly used to determine the reaction plane of the collision [31]. The target during the 1996 Pb run was a 0.2 mm thick lead foil placed in the center of the Plastic Ball detector.

The Silicon Drift and the Silicon Pad Multiplicity Detectors. The multiplicity and pseudorapidity of charged particles were measured by the Silicon Drift Detector (SDD) and the Silicon Pad Multiplicity Detector (SPMD) [32], located 12.5 and 30 cm downstream of the target. The pseudorapidity range of the SDD was $2.0 < \eta < 3.4$, and $2.35 < \eta < 3.75$ for the SPMD.

The Goliath magnet. A magnetic field with a bending power of 1.6 Tm was provided by the Goliath magnet. The center of the magnet was located 3.28 m downstream of the target. The polarity of the field was such that negatively charged particles were bent into the first tracking arm (described in section 2.2.1 below), while positively charged particles were bent into the second tracking arm (section 2.2.2). During the 1996 run, the polarity of the field was reversed during shorter periods.

The Mid-Rapidity and the Zero Degree Calorimeters. The last part of the trigger system consists of the Mid-Rapidity Calorimeter (MIRAC) and the Zero Degree Calorimeter (ZDC). MIRAC provided a complete coverage of the transverse energy, E_T, over the forward pseudorapidity interval of $3.5 < \eta < 5.5$ and was used to estimate the energy density, ϵ_{Bj} as well as to provide the centrality trigger. The remnants of the projectile was measured by the ZDC at angles less than $0.3°$ to the beam.

The Lead Glass Calorimeter. The Lead Glass Calorimeter or the Lead glass Detector Array (LEDA) was designed to detect thermal photons [33]. It consisted of 10 080 lead-glass modules each with a size of $4{\times}4{\times}40$ cm^3 and had a good energy resolution as well as fine granularity needed for a high resolution and a high reconstruction efficiency. The LEDA was placed 21.5 m downstream of the target, covering a pseudorapidity range of $2.28 < \eta < 2.95$ which corresponds to an angular range of $6° < \theta < 11.6°$.

The Charged Particle Veto Detector. In front of the LEDA, streamer tube detectors [34] equipped with pad readout were placed. Serving as charged particle veto detectors, it improved the photon reconstruction efficiency in the LEDA. Except for the pad size and tube length, these were of the same construction as the two planes of streamer tube detectors in the second tracking arm (described below).

The Photon Multiplicity Detector. Located 21.5 m from the target, the Photon Multiplicity Detector (PMD) [35] was a large and highly segmented pre-shower detector with a pseudorapidity coverage of $2.8 < \eta < 4.4$. It measured the photon multiplicity by conversion into electromagnetic showers in a 2 cm lead layer. The showers were detected by scintillator pads read out by CCD-cameras. Together with the SPMD, the PMD has been used to study fluctuations in the N_{γ}/N_{ch} ratio on an event-by-event basis, in the search for disoriented chiral condensates [36].

2.2.1 The first tracking arm

The first tracking arm consisted of six tracking planes followed by a time-of-flight wall used for the particle identification. During a normal run, the orientation of the magnetic field bent negatively charged particles into the first tracking arm.

The Multi-Step Avalanche Chambers. The gaseous Multi-Step Avalanche Chambers (MSAC:s) [37, 38, 39, 40] used as tracking planes, were equipped with an optical readout by CCD cameras. The chambers were located between 3.4 m and 9.4 m from the center of the Goliath magnet.

The Russian Time of Flight Wall. The time-of-flight wall, known as the RTOF, allowed particle identification of π^-, K^- and

$\overline{p}$. It was made of 480 plastic scintillators that were read out from both ends by photomultipliers. The time resolution was $\sigma_t = 120$ ps. It was located 13.5 m from the center of the magnet.

2.2.2 The second tracking arm

A second tracking arm was implemented in the experiment for the 1996 run, allowing the measurement of positively charged particles such as π^+, K^+ and p as well as the double charged Δ^{++} resonance [41] decaying into a π^+ and p. The combination of this arm with the first arm, rendered the possibility to reconstruct light vector mesons such as the ϕ-meson (see section 2.5).

The four tracking planes were two pad chambers (Fig. 2.2) and two streamer tube chambers (Fig. 2.3). Particle identification was achieved with a time-of-flight wall.

The development and construction of this tracking arm was the main contribution from the Lund group to the experiment in 1996. The chambers and their performance are described in more detail in paper I.

The Pad Chambers. The so-called pad chambers are in principle MSAC:s of the same type as in the first arm but are equipped with a pad readout [42, 43, 44] instead of CCD camera readout. The number of pads per chamber is approximately 36 000, each individually read out. The size of a pad is 1×17 mm^2. A hit is defined as two or more neighbouring pads that have received charge above a pedestal value. The pad layout is shown in Fig. 4.1.

The Streamer Tubes. The two streamer tube planes, located after the pad chambers, were also equipped with a pad readout. There are about 6 000 pads per plane with a size of 7×22 mm^2.

The Japanese Time of Flight Wall. The particle identification is done with a time-of-flight wall known as the JTOF and was constructed by the Japanese team. It consists of 480 slats of scintillators that are read out at both ends by phototubes. The