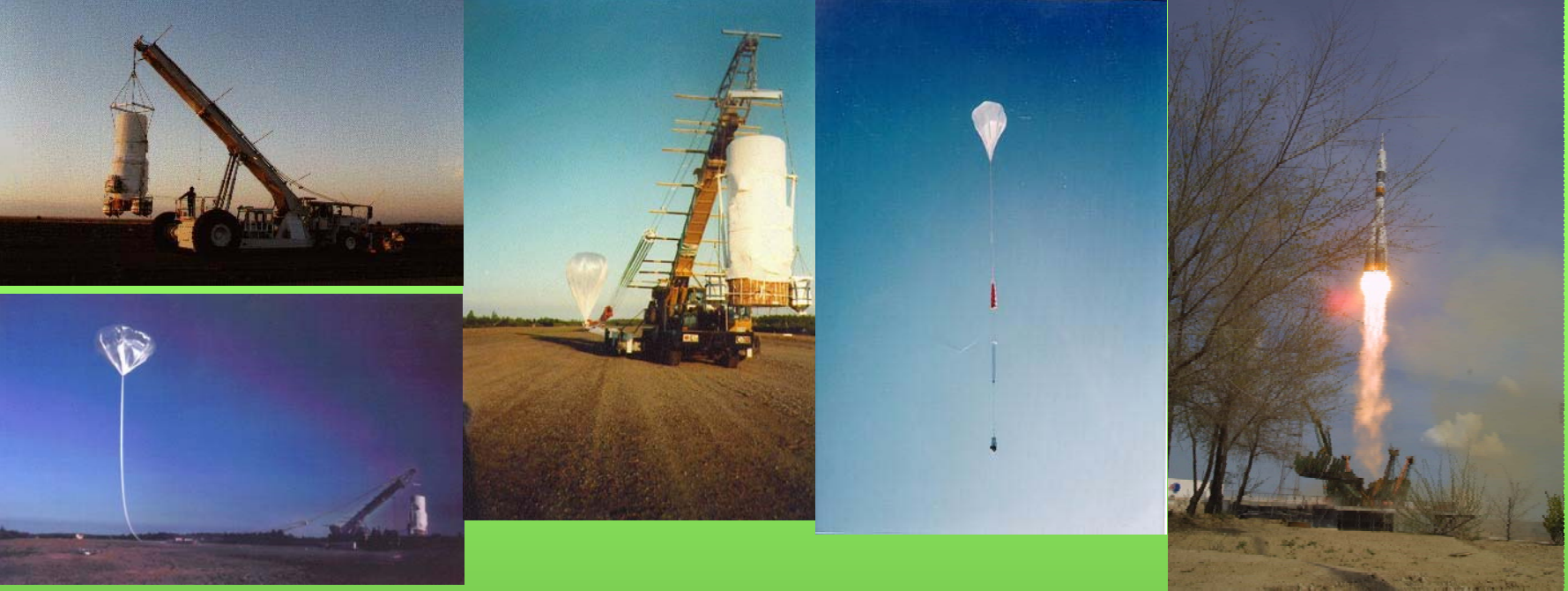




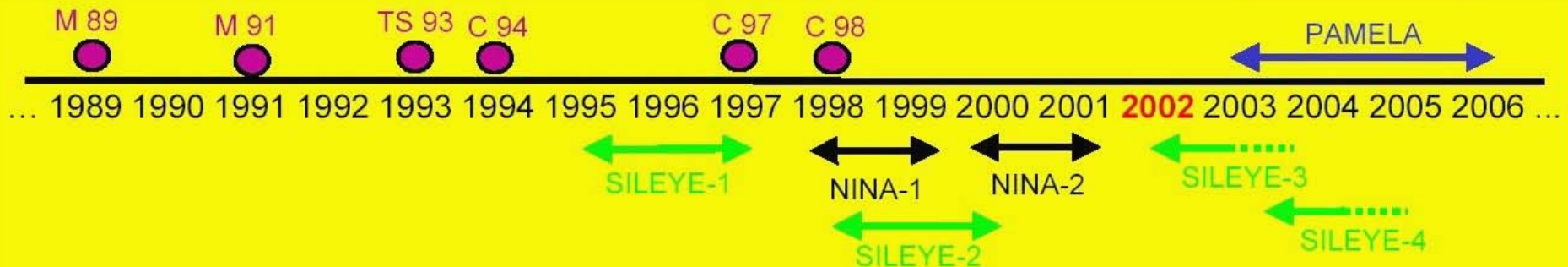
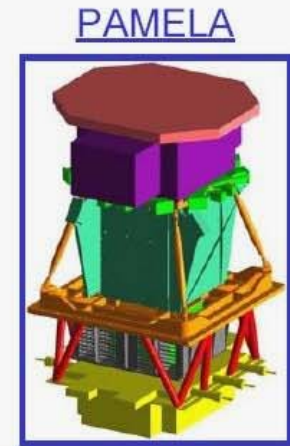
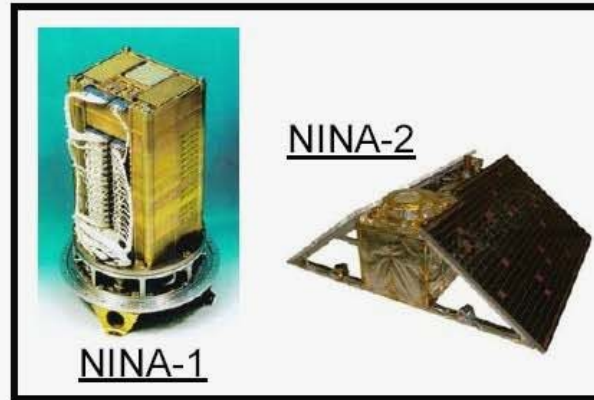
Silicon-Tungsten Calorimeters of the WiZard Collaboration

OUTLINE

- Brief History
- Calorimeters Description
- Simulation and Comparison with Experimental Data
- Identification Capabilities
- Front-End Electronics (*Gigi's talk*)
- Future developments (*Gigi's talk*)

The WiZard Collaboration experiments

MASS-89, 91, TS-93,
CAPRICE 94-97-98



SILEYE-1



SILEYE-2



ALTEINO: SILEYE-3



ALTEA:
SILEYE-4

Italy:



Bari



Florence



Frascati



Naples



Rome



Trieste



CNR, Florence

Russia:



Moscow
St. Petersburg

India:



Bombay

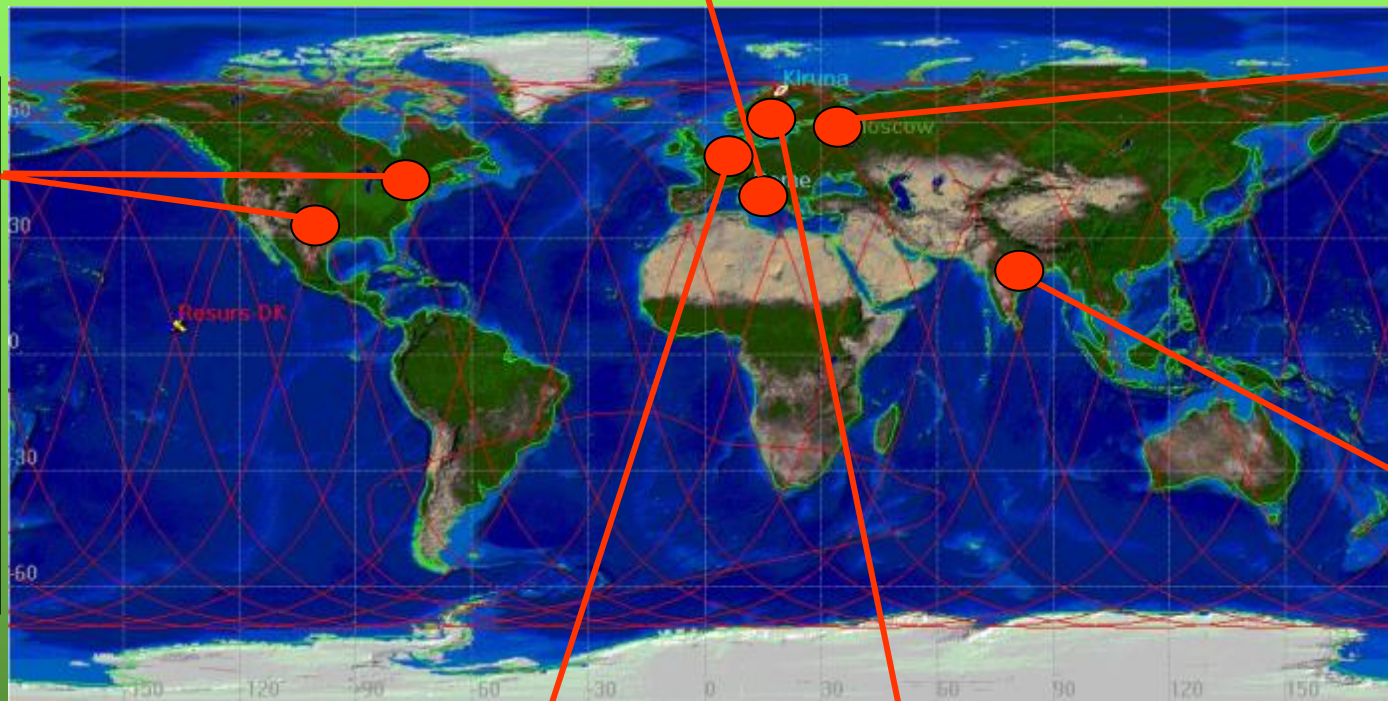
USA:



GSFC



NMSU



Germany:



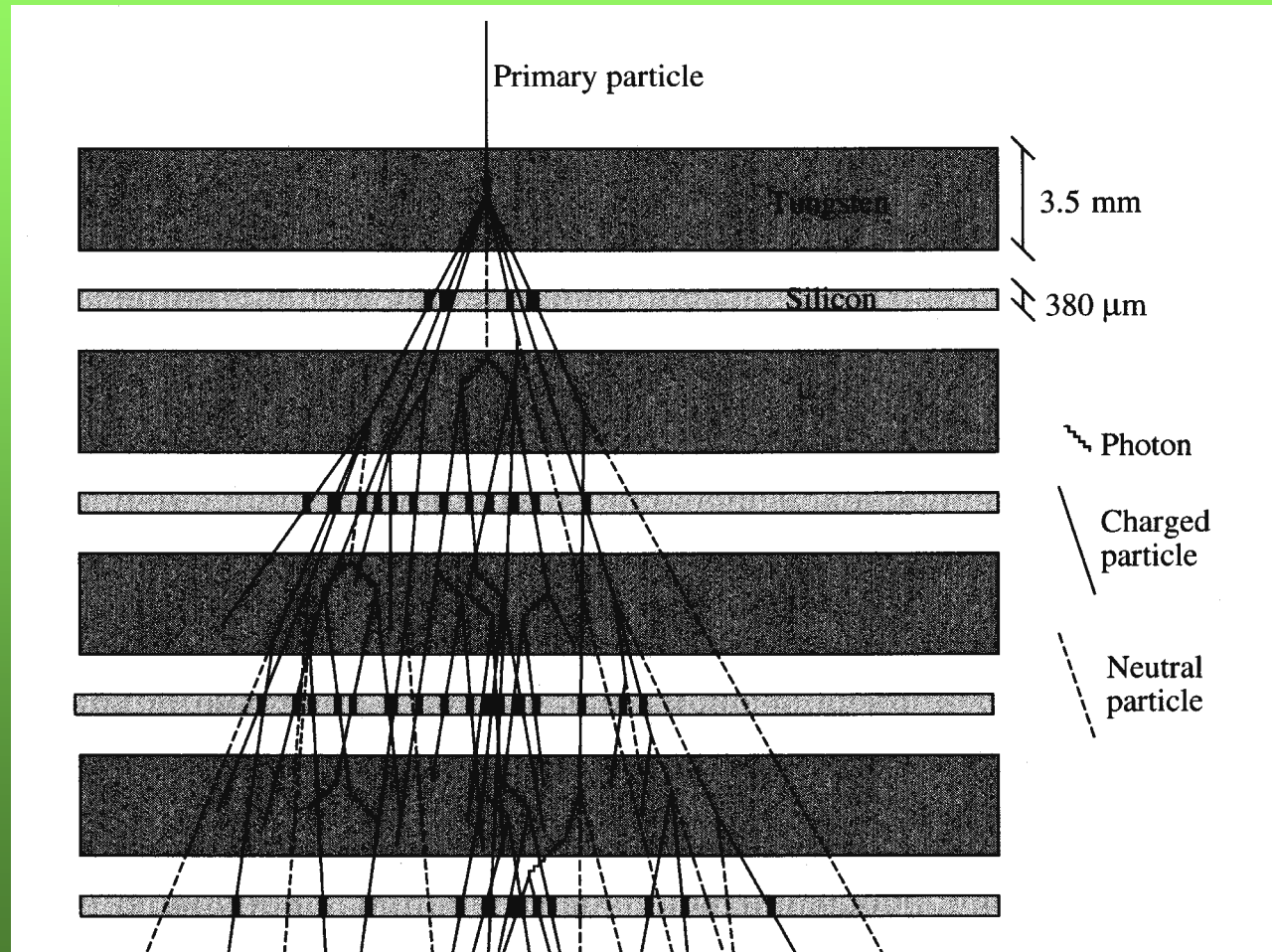
Siegen

Sweden:

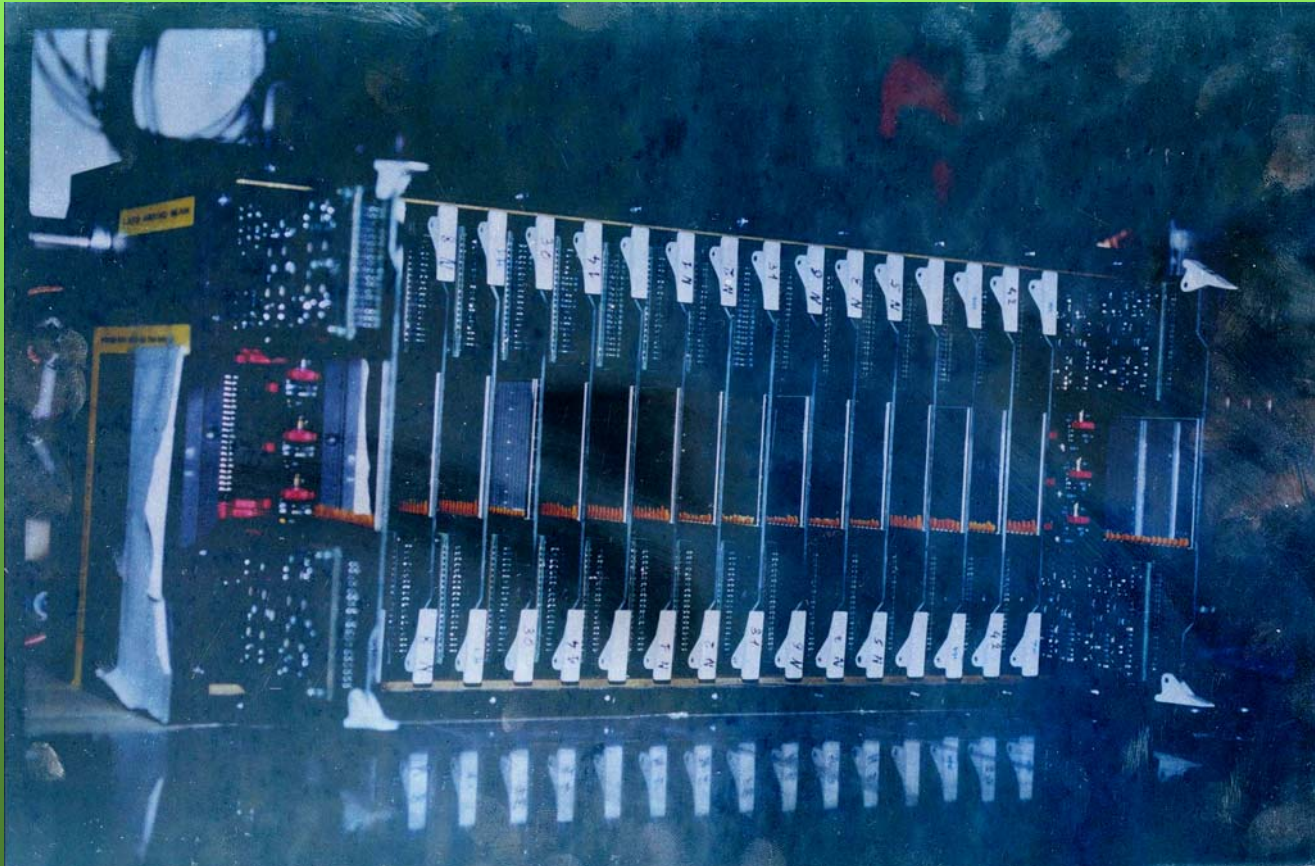


KTH, Stockholm

Calorimeter Principle

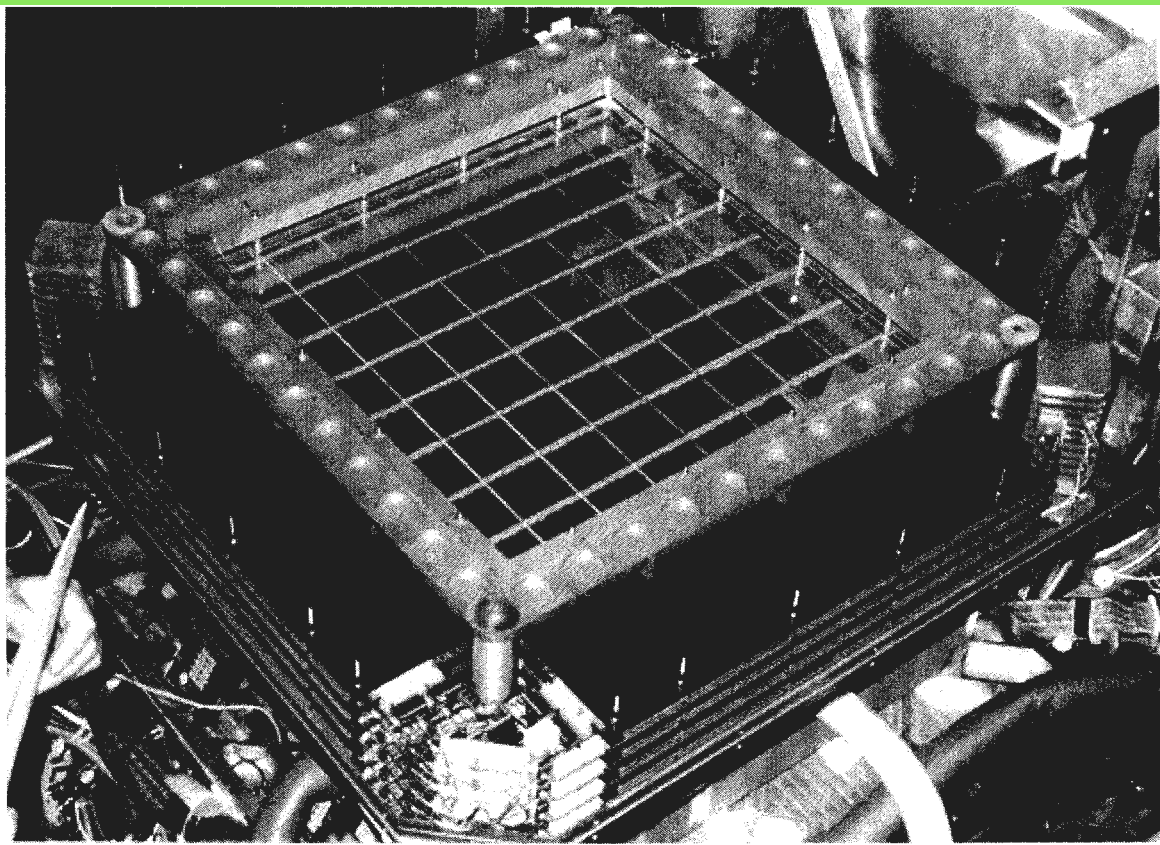


First Calorimeter Prototype



20x2 (X and Y views) 60x60 cm² Si detectors interleaved by 19 W layers of 0.5 X₀.
Test beams 1991-1992.

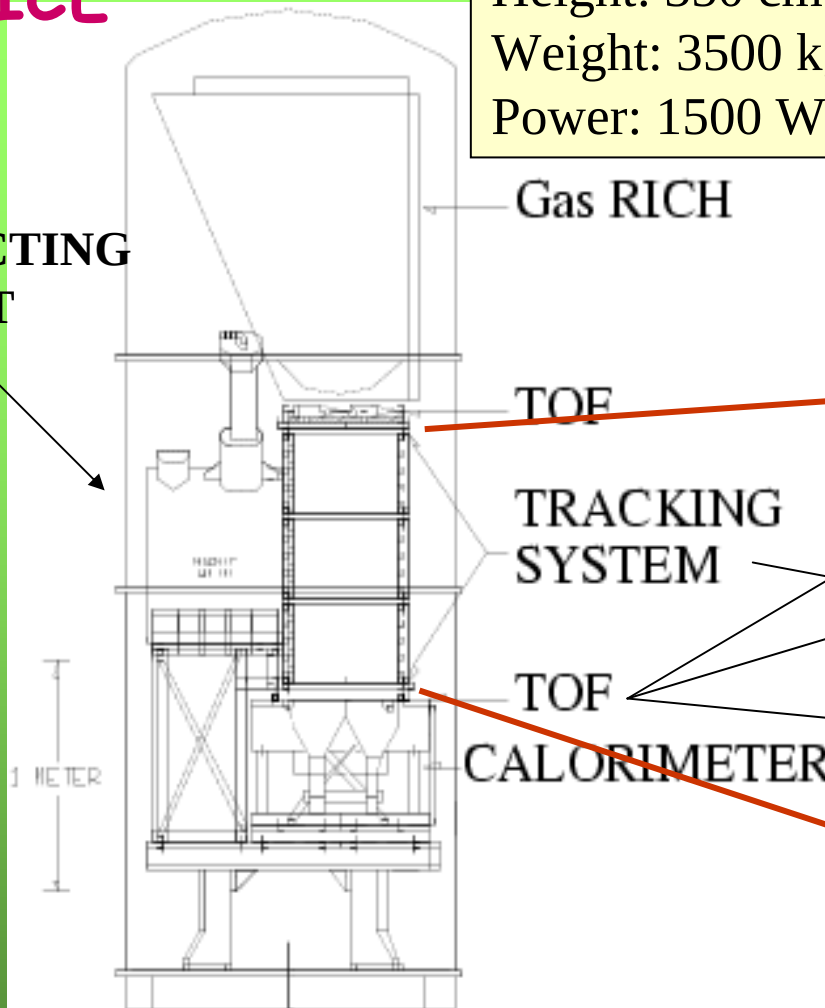
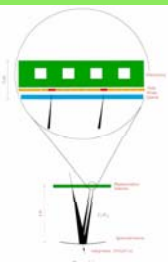
CAPRICE Calorimeter



8 planes made by
64x2 (X and Y views)
60x60 cm² Si detectors
interleaved by W layers
of 1 X_0 .
It flew on 1993, 1997
and 1998

CAPRICE

**SUPER
CONDUCTING
MAGNET**

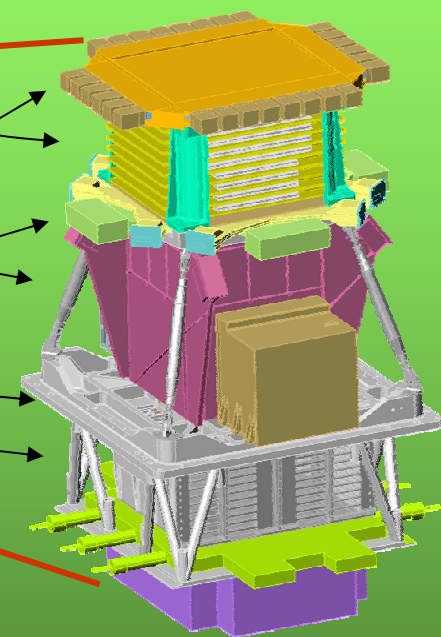


Height: 350 cm
Weight: 3500 kg
Power: 1500 W

PAMELA

Height: 120 cm
Weight: 450 kg
Power: 360 W

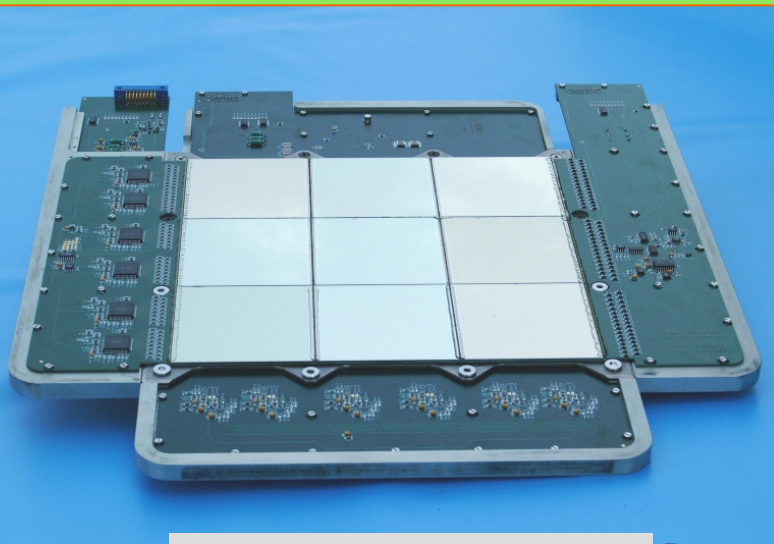
TRD



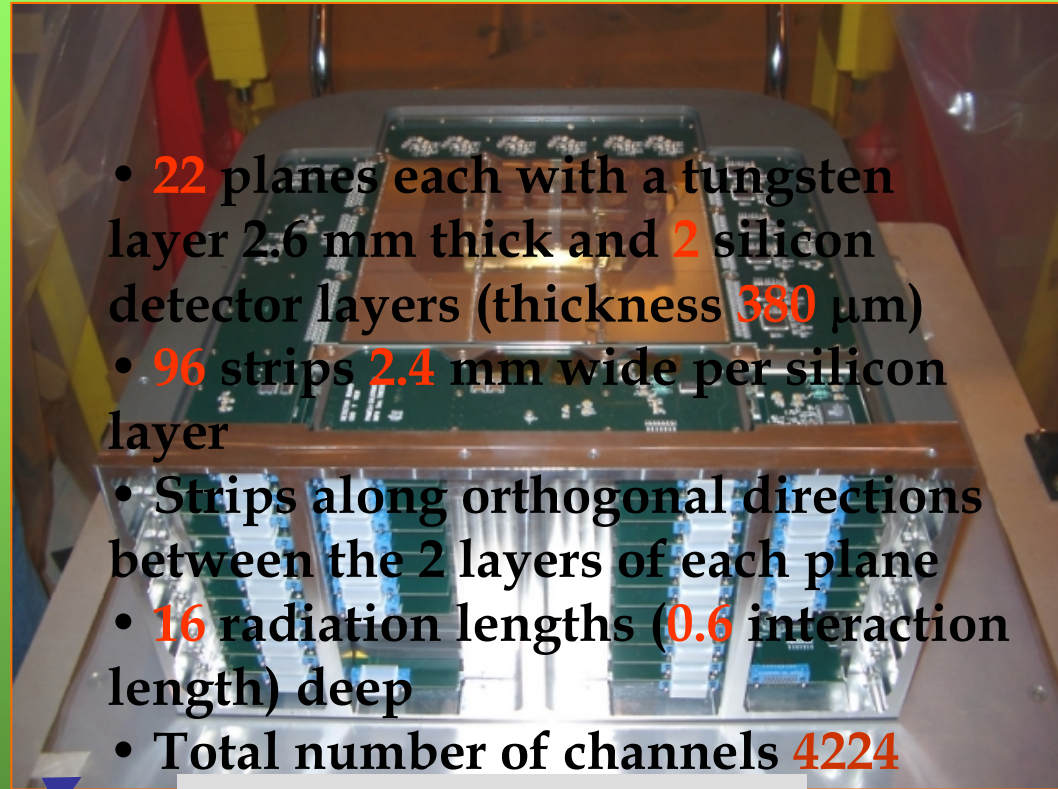
1.2 m

**from balloon-borne detectors
to satellite experiments**

The imaging calorimeter



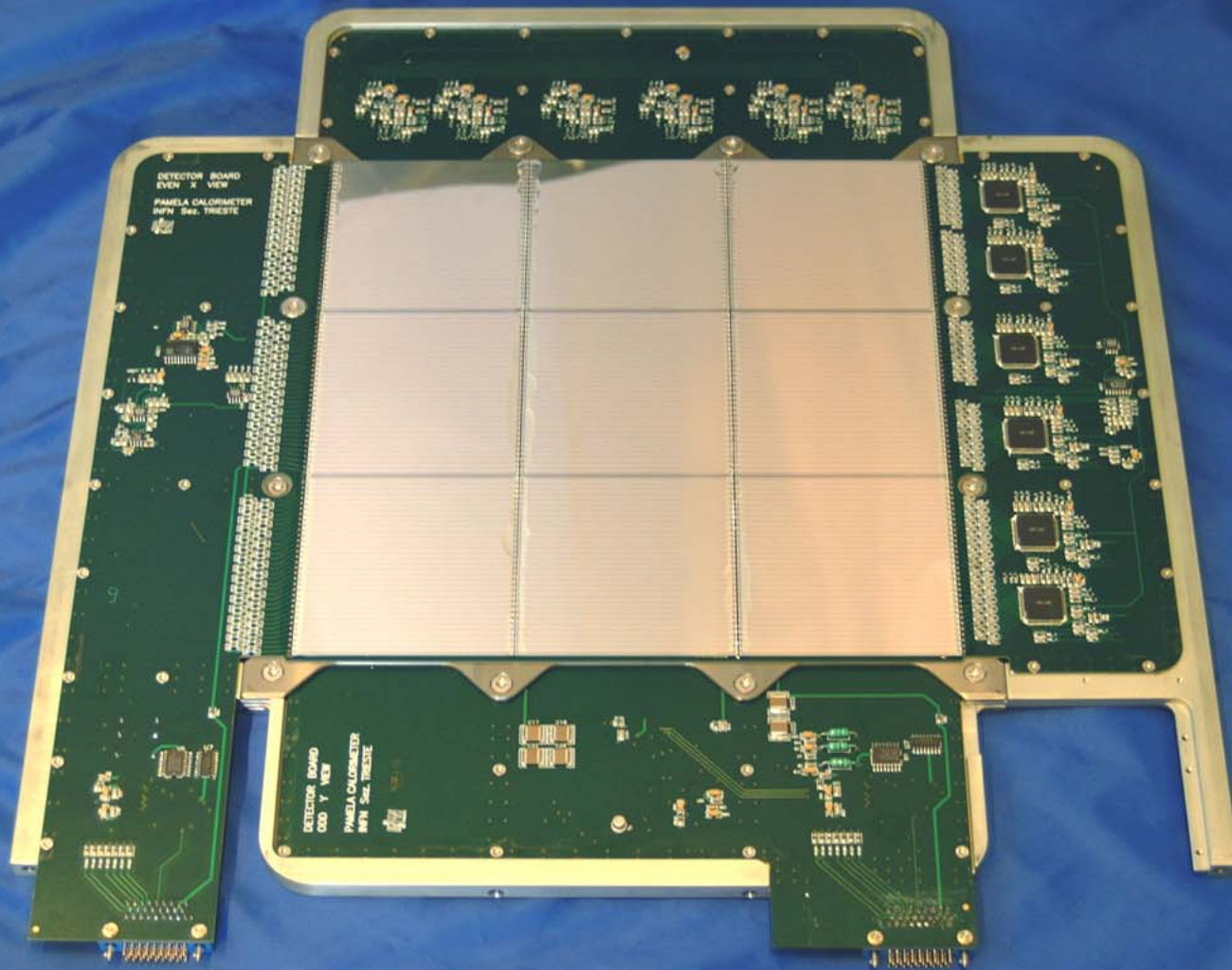
calorimeter plane



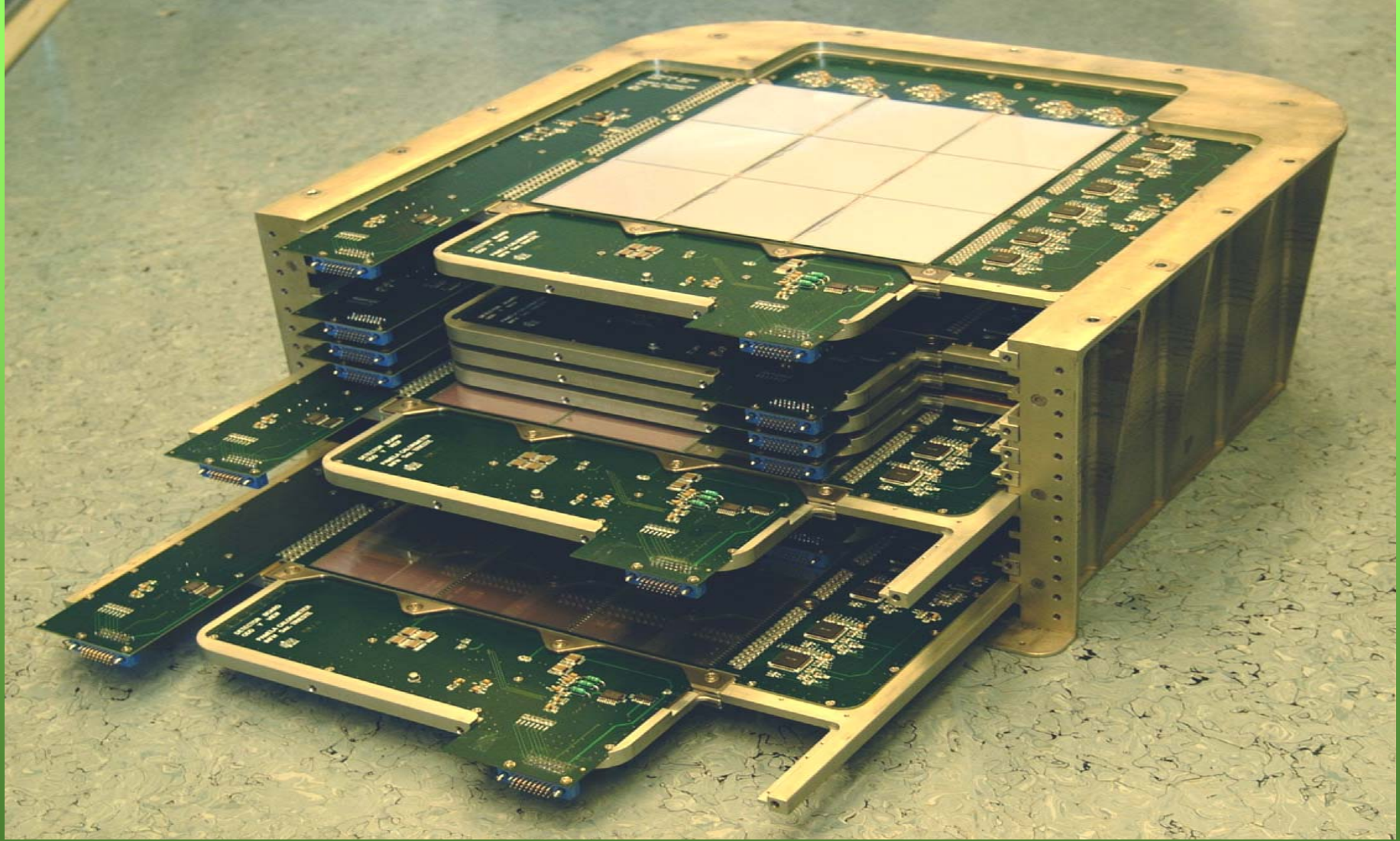
- **22** planes each with a tungsten layer 2.6 mm thick and **2** silicon detector layers (thickness **380** μm)
- **96** strips **2.4** mm wide per silicon layer
- Strips along orthogonal directions between the 2 layers of each plane
- **16** radiation lengths (**0.6** interaction length) deep
- Total number of channels **4224**
- Wide dynamic range $\approx 1 \div 1000$ mip

Assembled calorimeter

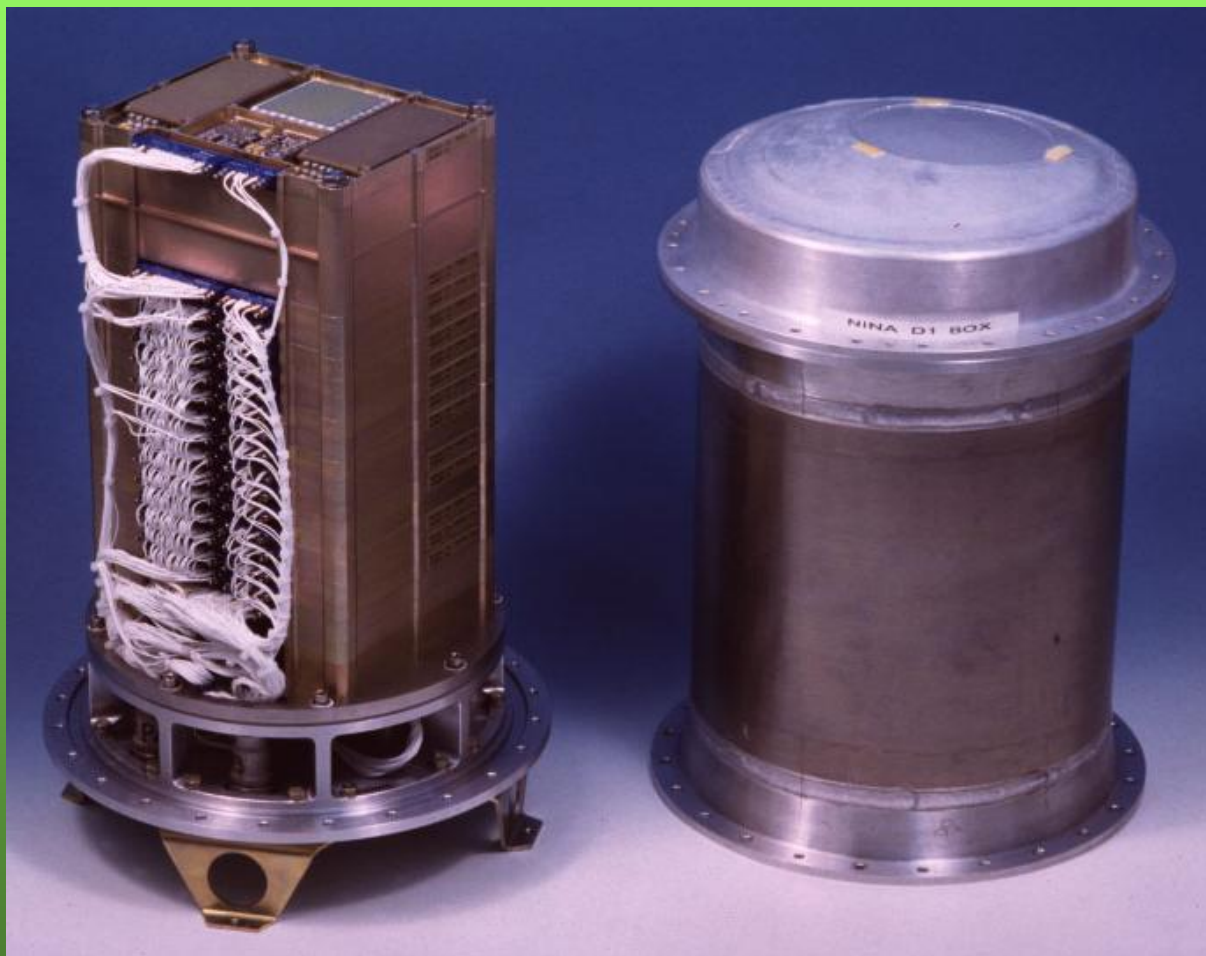
A calorimeter plane



The imaging calorimeter



NINA: a Si Space Instrument



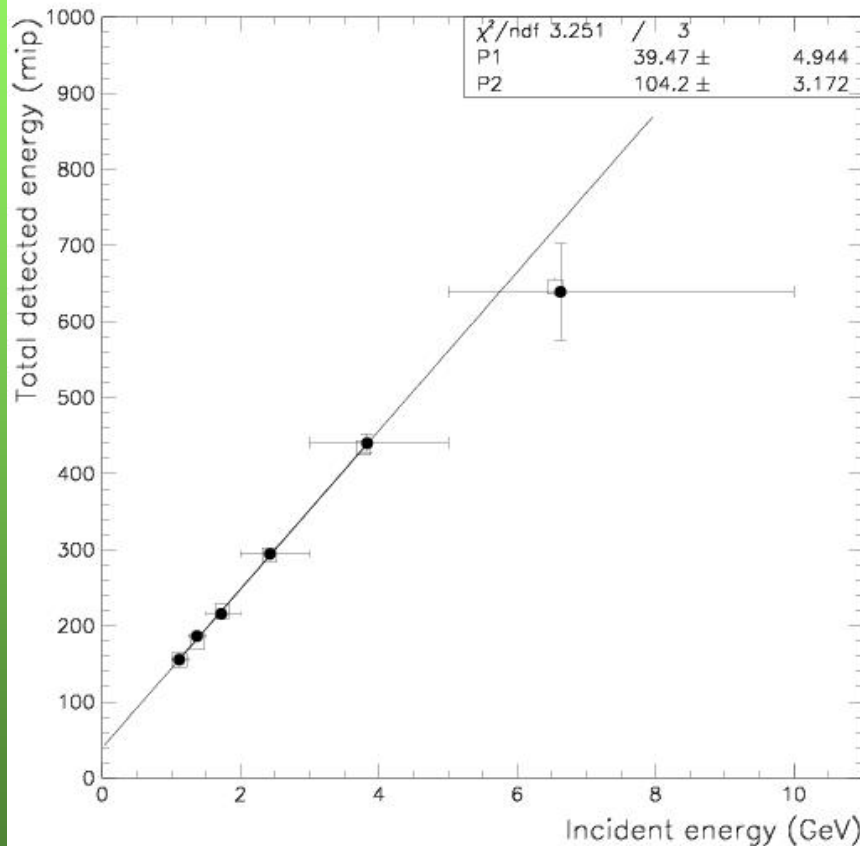
16 Si detectors 60×60 cm^2 to detect low energy particles via their energy losses. One version was launched with a satellite in 1998 and one in 1999.

Simulation Package

- Package used **GEANT** both for the description of the geometry and physics.
- Hadronic packages: **GHEISHA** & **FLUKA** (old version)
- Comparison performed both with test beam data and cosmic-ray data

Simulation - Cosmic-Ray Data

Caprice Calorimeter

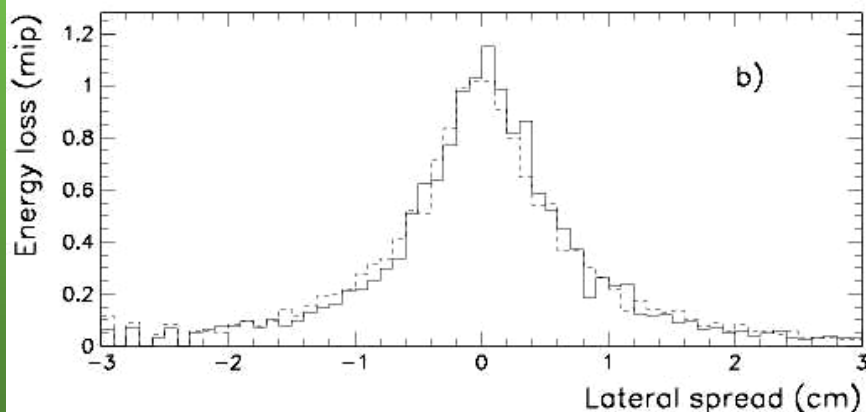
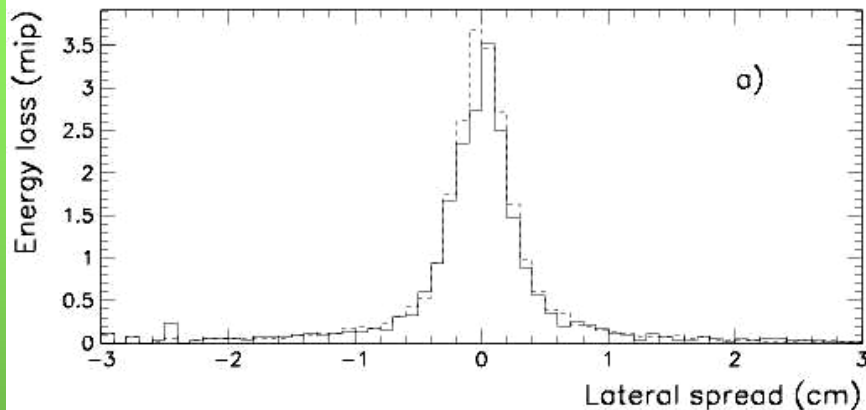


Energy
Measurement

- Flight data
- Simulation

Simulation - Cosmic-Ray Data

Caprice Calorimeter

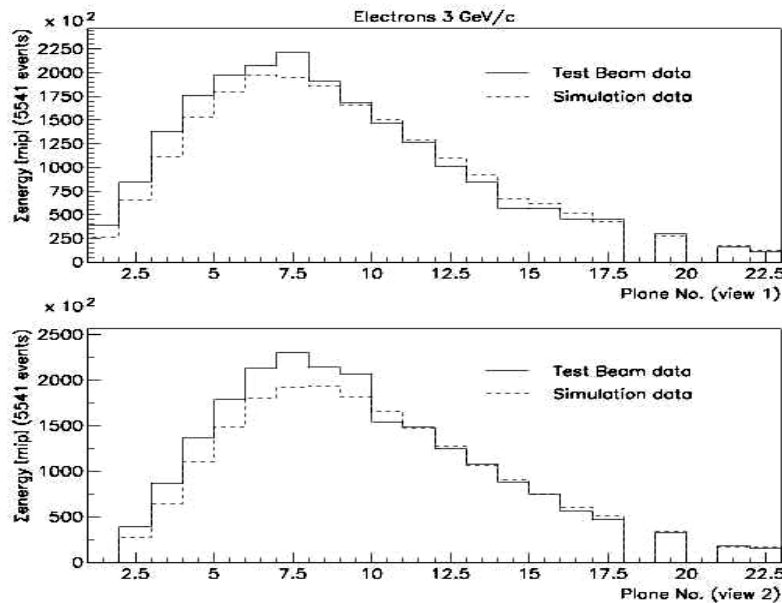


Lateral
Profile

- Solid line: Flight data
- Dashed line: Simulation data

Simulation - Cosmic-Ray Data

Pamela Calorimeter

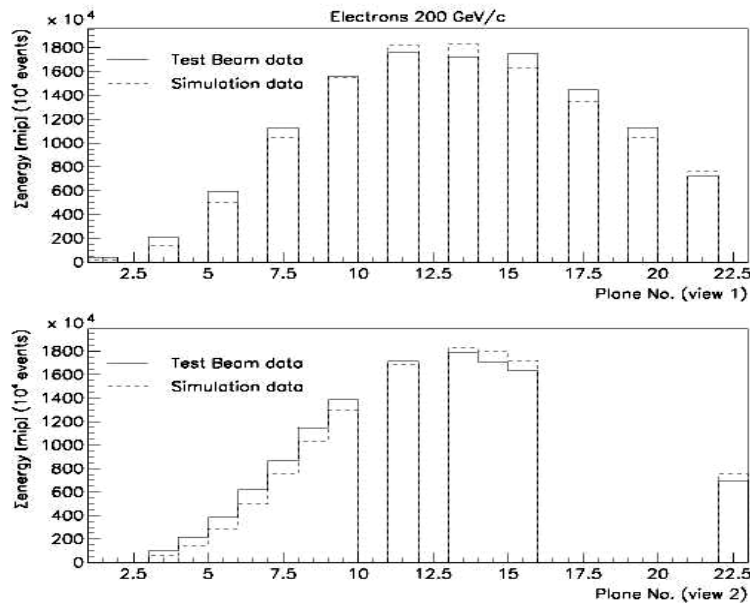


Longitudinal
Profile
(3 GeV/c e^-)

- Solid line: Test beam data
- Dashed line: Simulation data

Simulation - Cosmic-Ray Data

Pamela Calorimeter

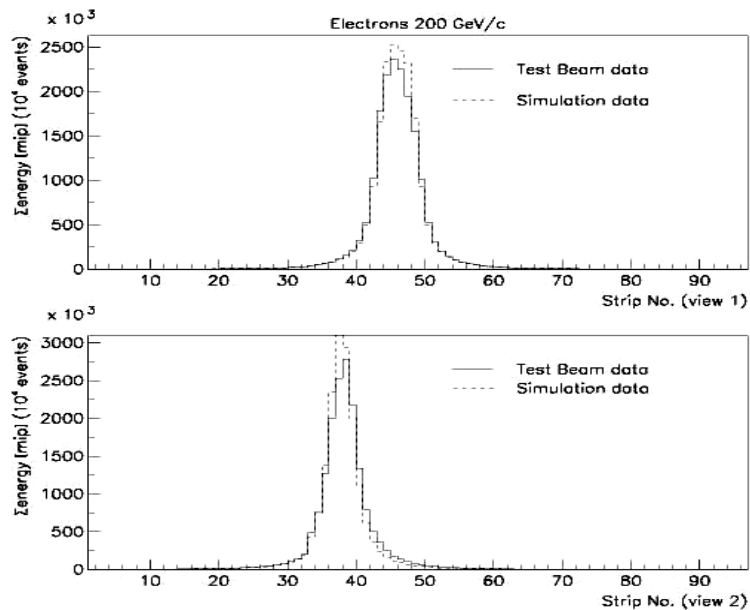


Longitudinal
Profile
(200 GeV/c e^-)

- Solid line: Test beam data
- Dashed line: Simulation data

Simulation - Cosmic-Ray Data

Pamela Calorimeter

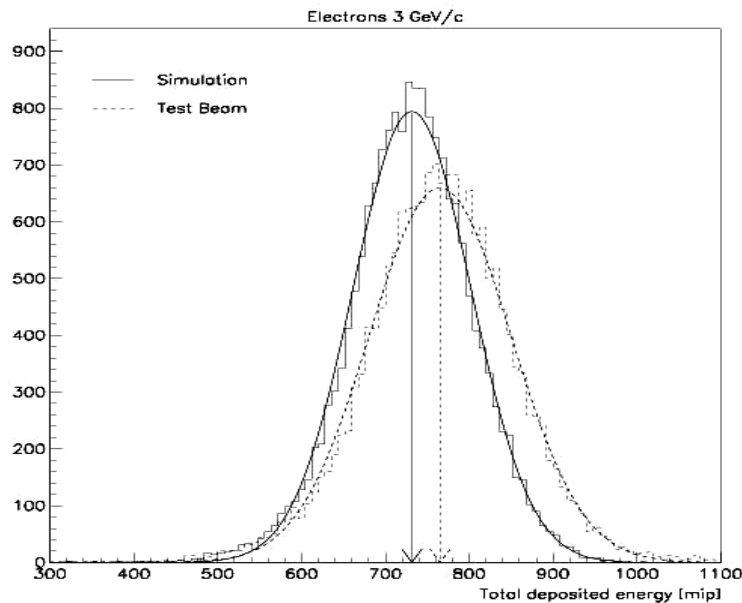


Lateral
Profile
(200 GeV/c e^-)

- Solid line: Test beam data
- Dashed line: Simulation data

Simulation - Cosmic-Ray Data

Pamela Calorimeter

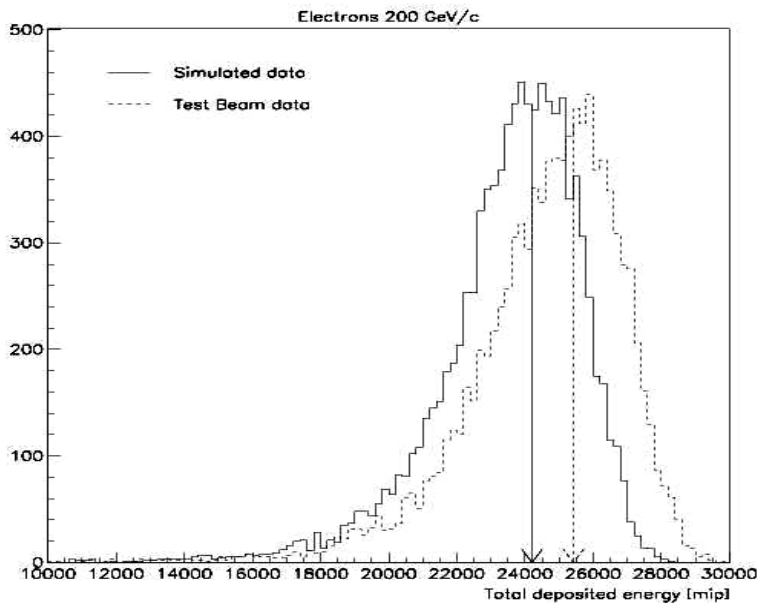


Energy
Measurement

- Solid line: Test beam data
- Dashed line: Simulation data

Simulation - Cosmic-Ray Data

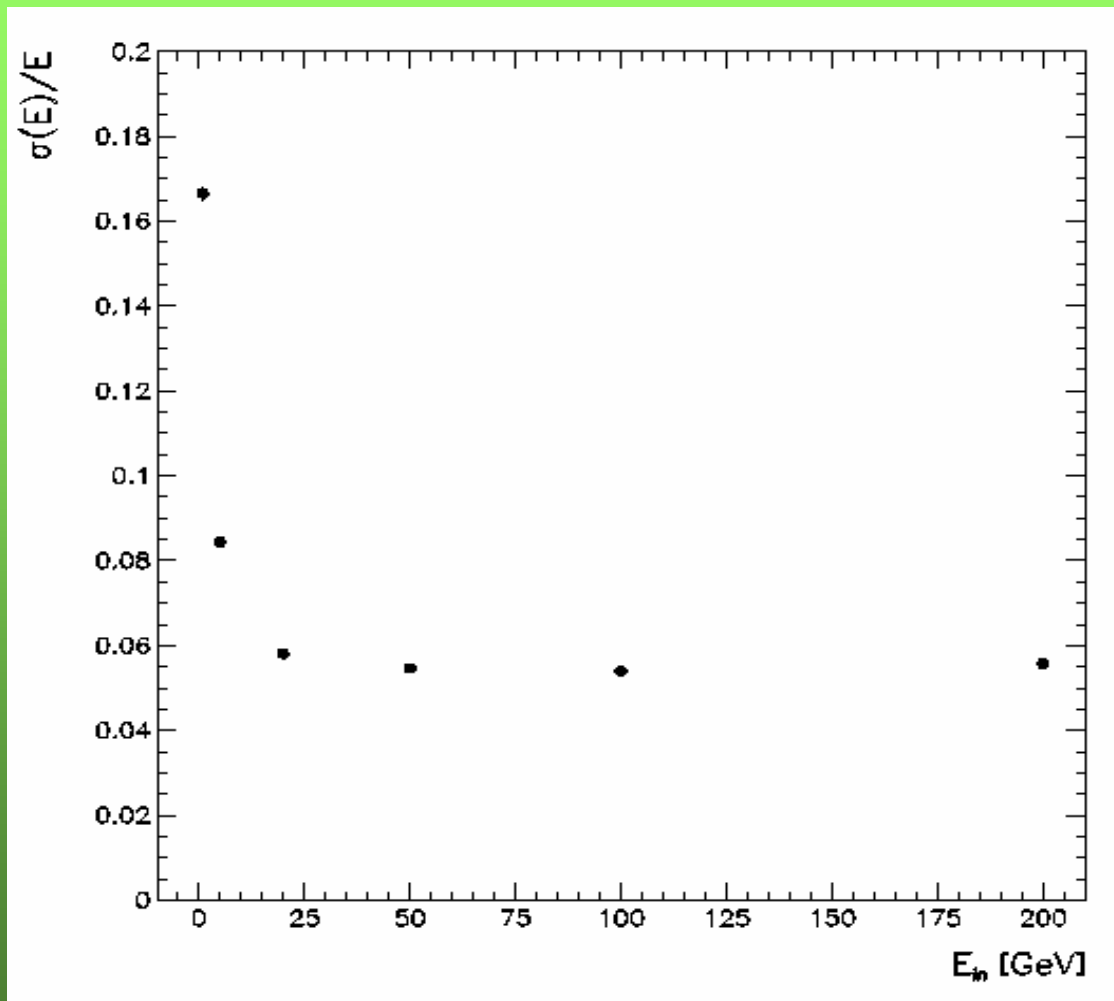
Pamela Calorimeter



Energy
Measurement

- Solid line: Flight data
- Dashed line: Simulation data

Calorimeter response (simulation)



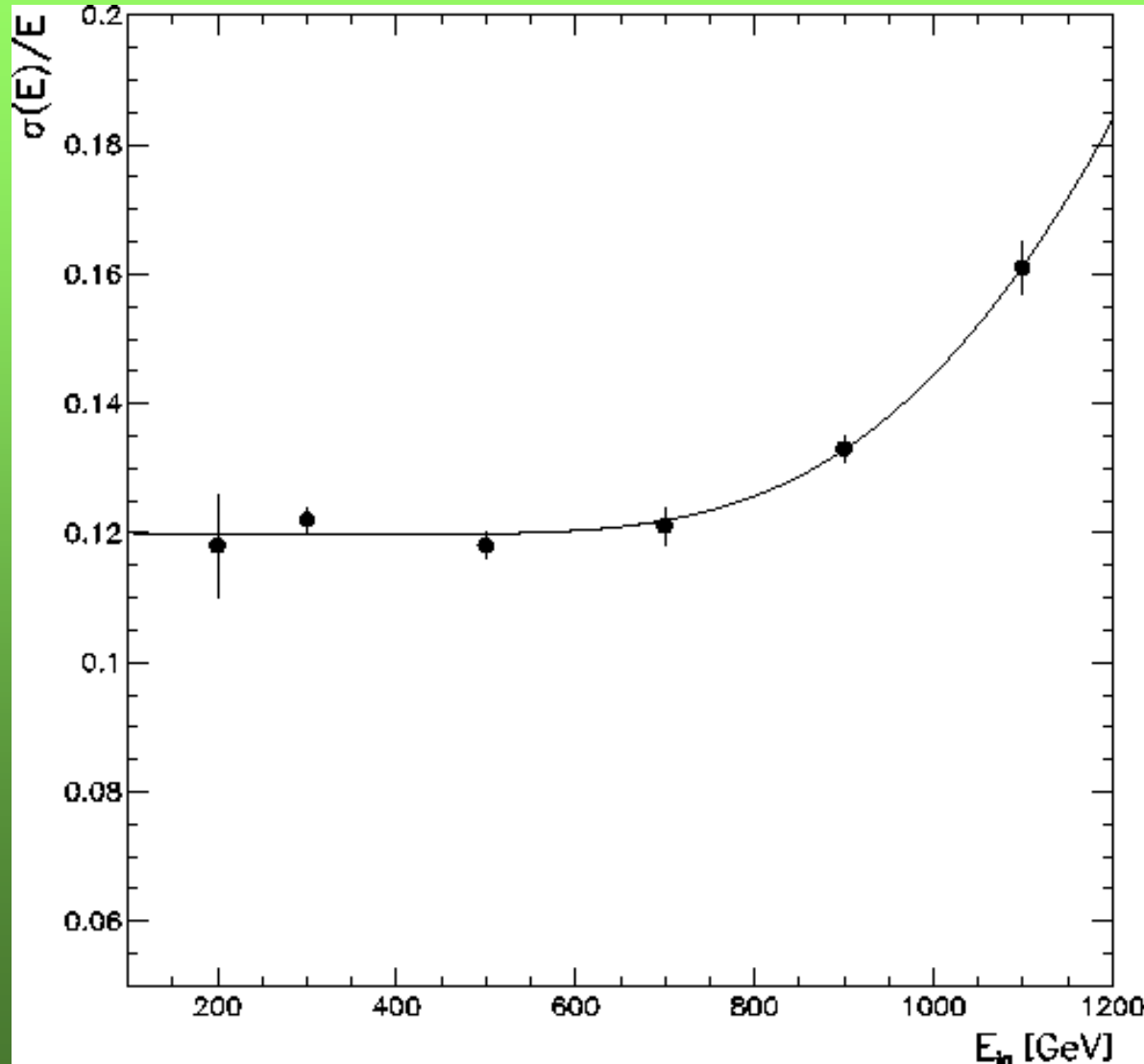
Expected energy resolution in the actual PAMELA configuration for cosmic-ray electrons

Calorimeter Response (simul.)

INFN Bologna, INFN Trieste, INFN Padova, INFN Ferrara, INFN Pisa, INFN Roma, INFN Torino, INFN Udine, INFN Venezia

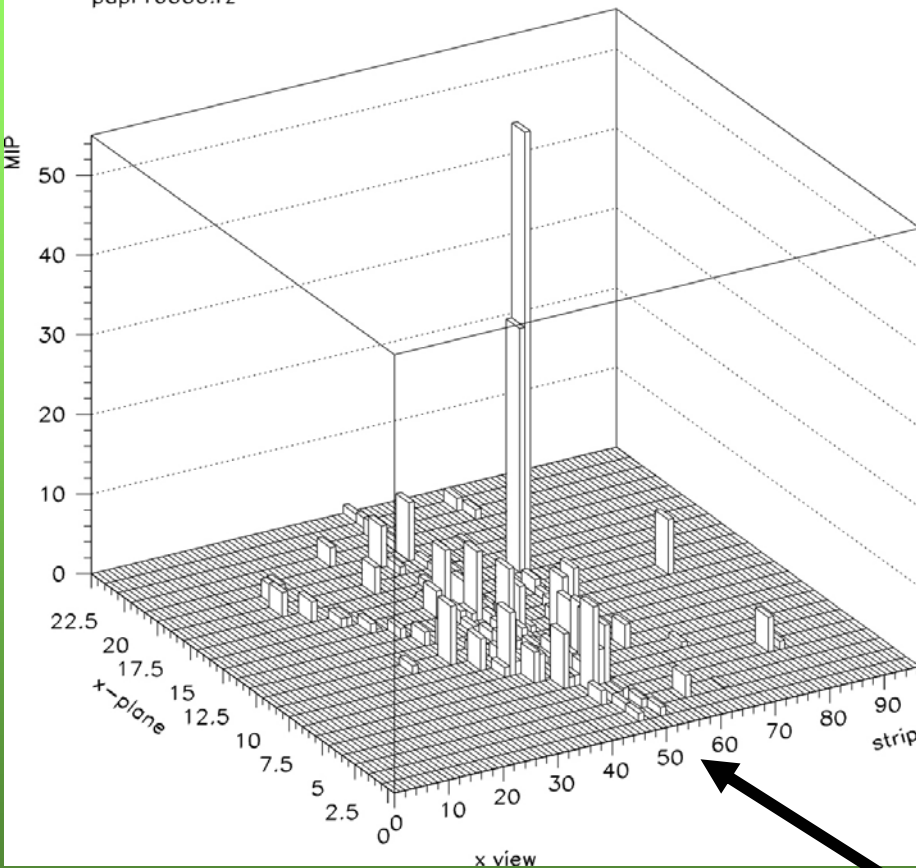
Calorimeter response

Simulated energy resolution in self-triggering mode



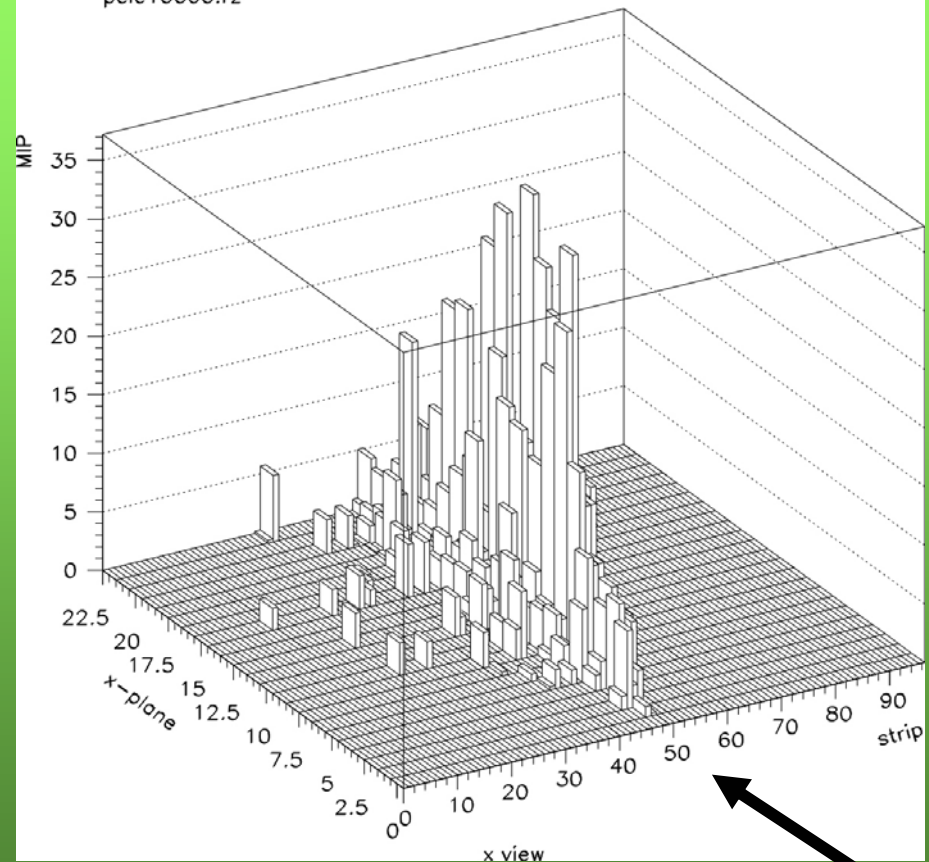
Calorimeter response (simulation)

papr10000.rz



10GeV $\bar{p}$

pele10000.rz

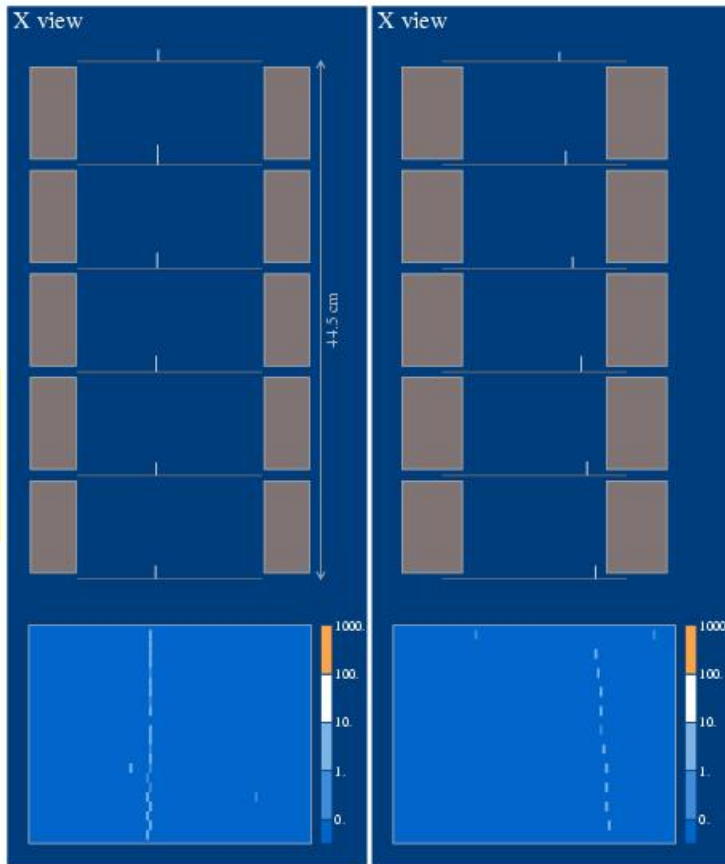


10GeV e^-

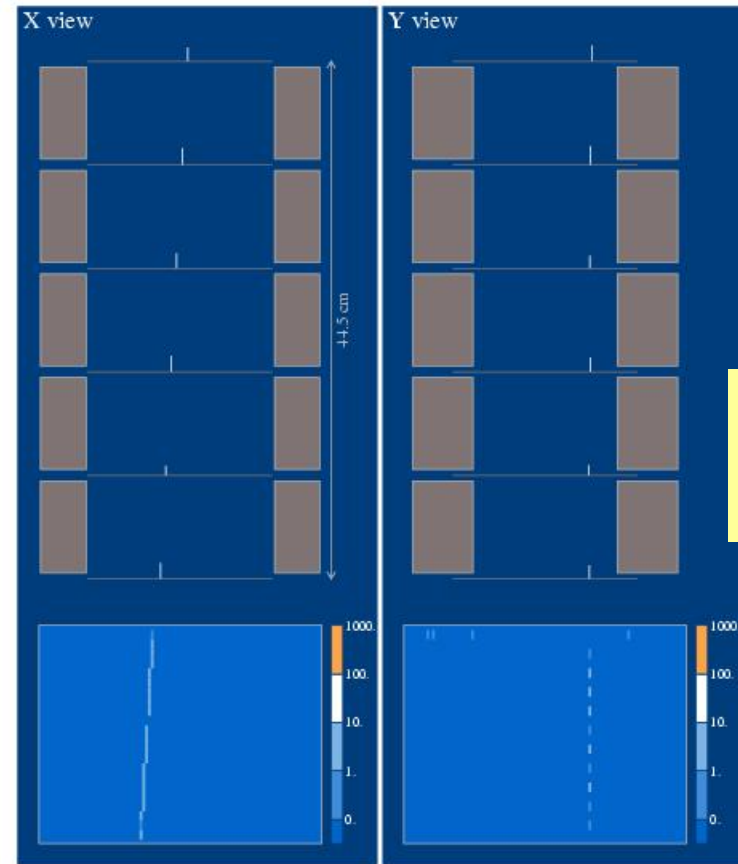
Conclusions about simulation

- A good agreement is found comparing simulation and experimental data for electromagnetic showers from 1 GeV up to 200 GeV and more. The simulation reproduces relatively well (differences of the order of few percent and less) the energy loss and its topological distribution.
- The simulation of hadronic showers does not give the same level of agreement obtained for the electromagnetic showers. However, differences are not large, of the order of 10-20%.

Real Events in the Calorimeter



Date 030921 File 293 Event 2069



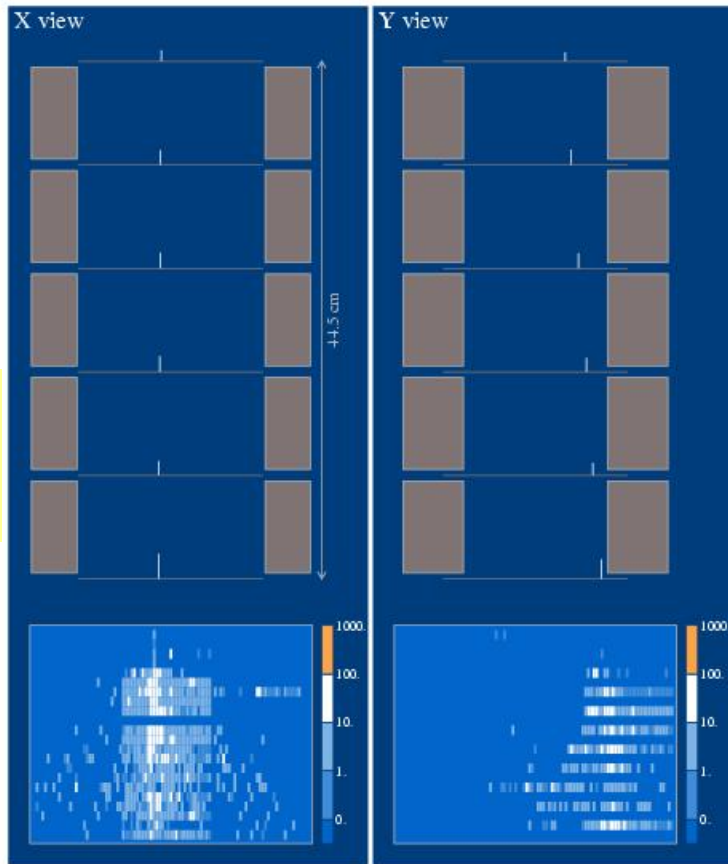
Date 030920 File 169 Event 115

SPS Test
Beam
Data: p
50 GeV/c

SPS Test
Beam Data:
p
100 GeV/c

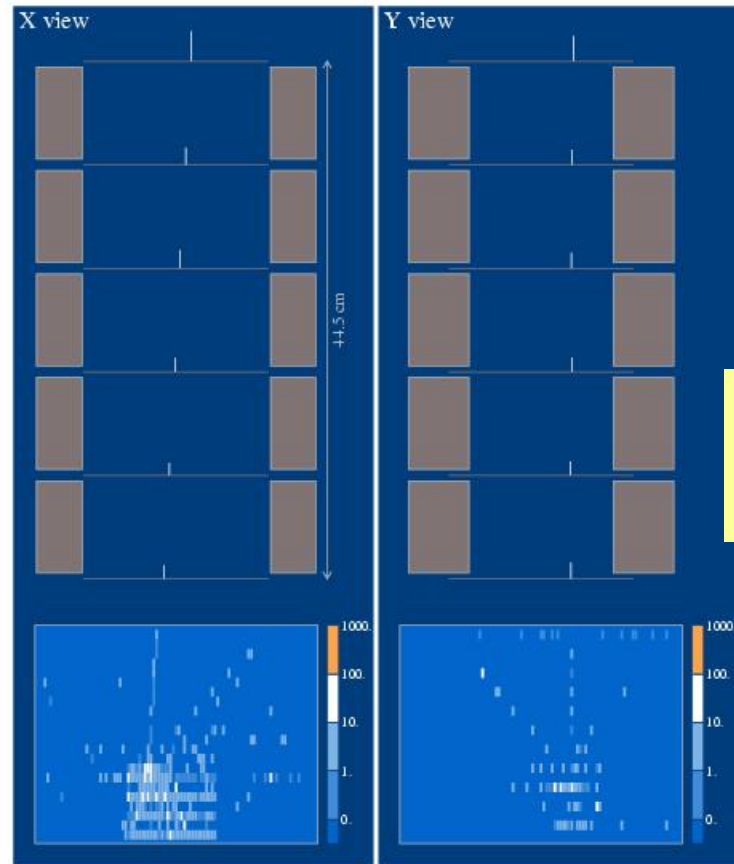
Real Events in the Calorimeter

SPS Test
Beam
Data: p
50 GeV/c



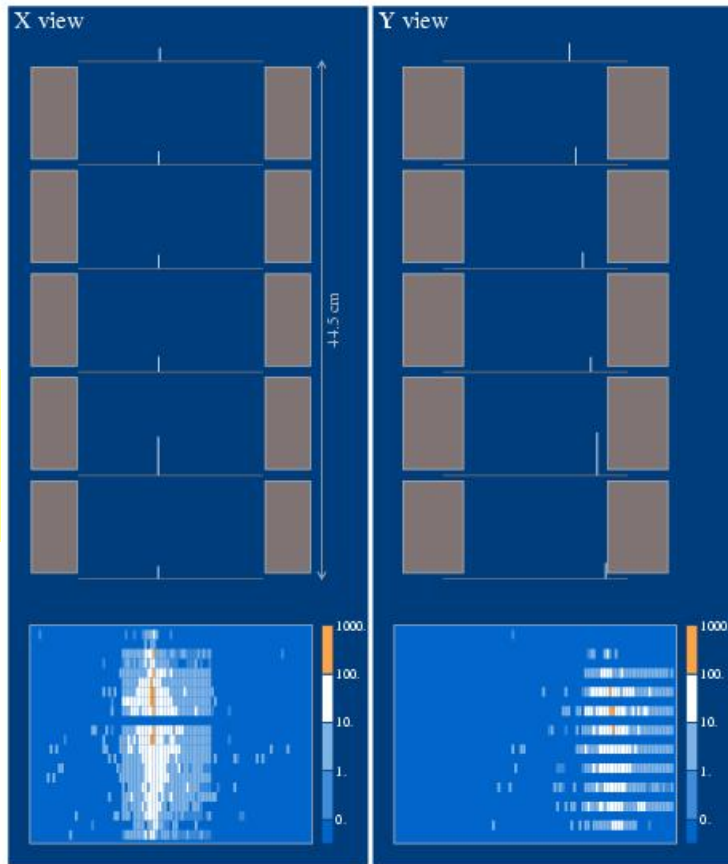
Date 030921 File 293 Event 2048

SPS Test
Beam Data:
p
100 GeV/c

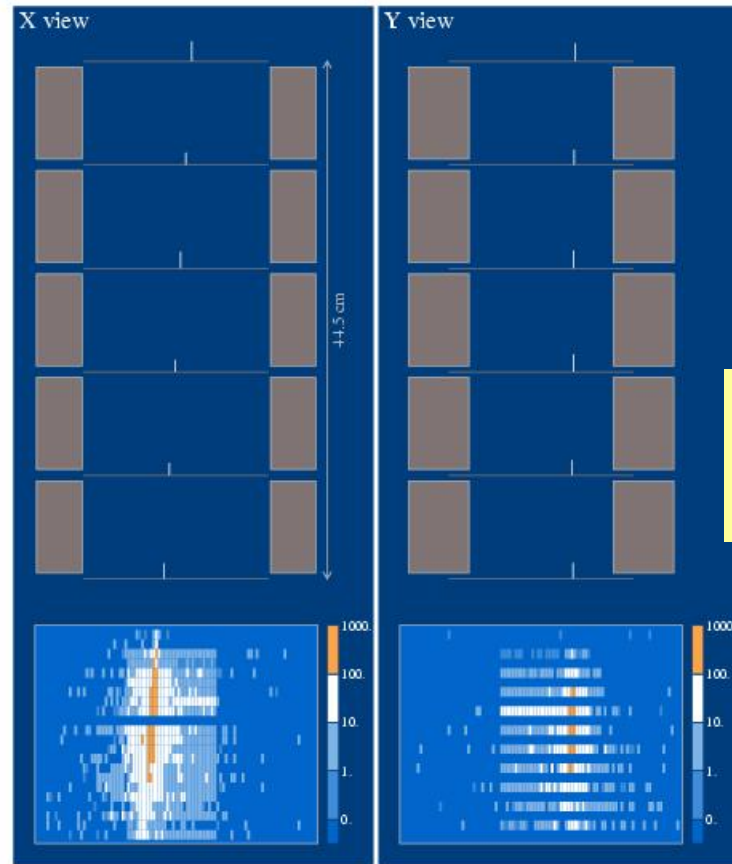


Date 030920 File 169 Event 1023

Real Events in the Calorimeter



Date 030921 File 323 Event 35



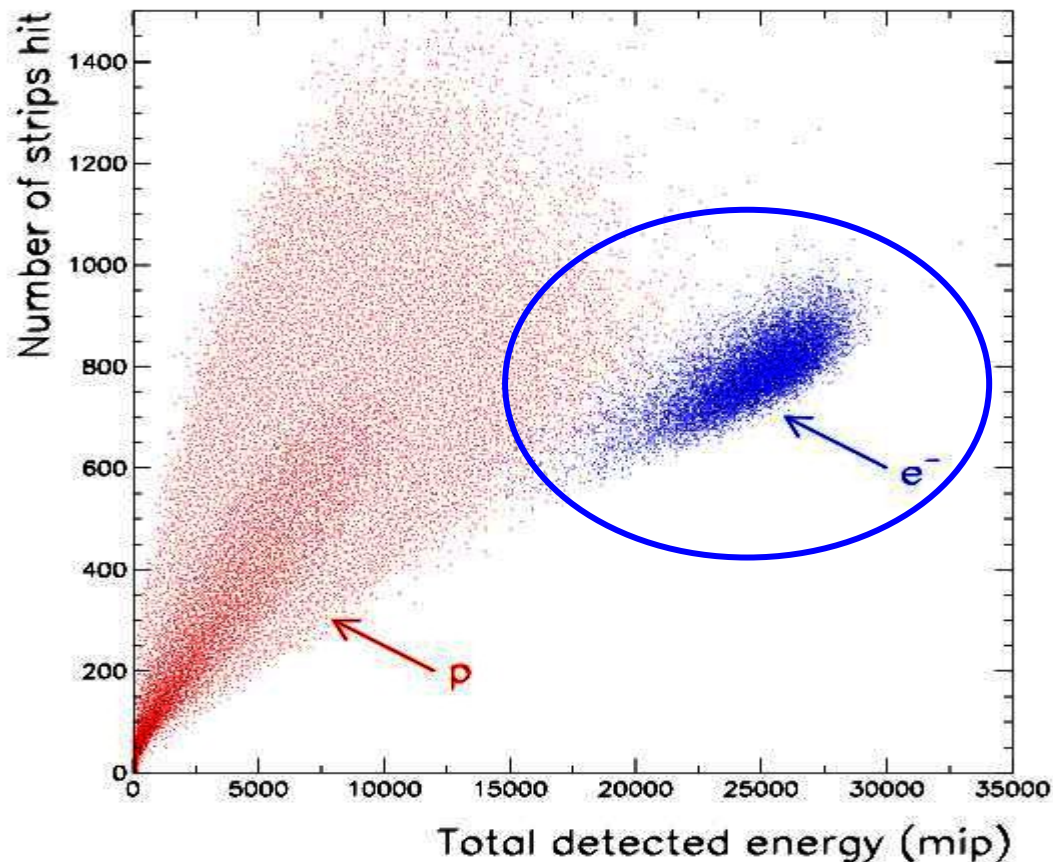
Date 030920 File 245 Event 6

SPS Test
Beam
Data: e^-
50 GeV/c

SPS Test
Beam Data:
 e^-
100 GeV/c

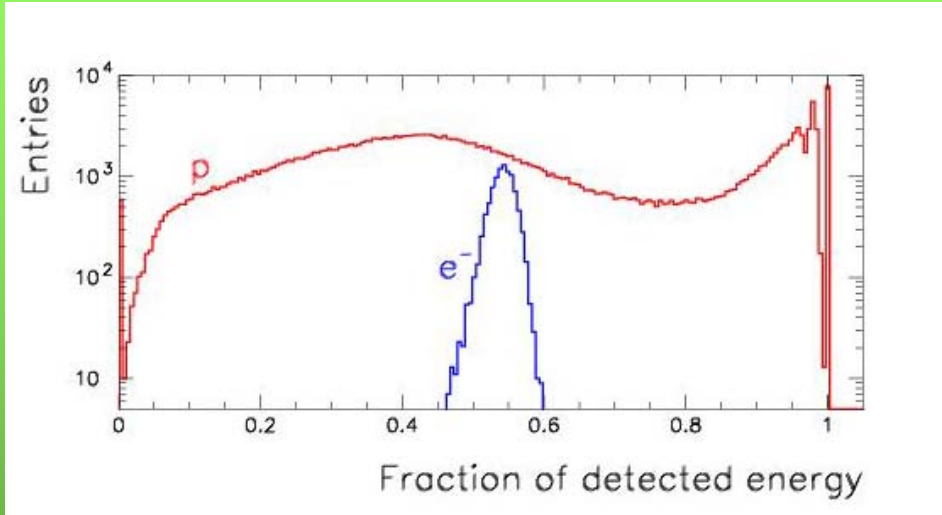
Electron-Proton separation (Calorimeter)

SPS Test
Beam Data:
 p & e^-
200 GeV/c



Electron-Proton separation (Calorimeter)

SPS Test
Beam Data:
p & e⁻
200 GeV



Proton rejection
factor $\cong 3 \cdot 10^{-4}$
Electron selection
efficiency $\cong 95\%$

