

# Cold matter affects J/ψ production

---

Raphaël Granier de Cassagnac  
LLR – École polytechnique / IN2P3

QM08 satellite meeting  
Tata Institute for Fundamental Research  
Mumbai, 2008, February 12<sup>th</sup>



## Reminder of the two striking behaviors of $J/\psi$ suppression at RHIC energy

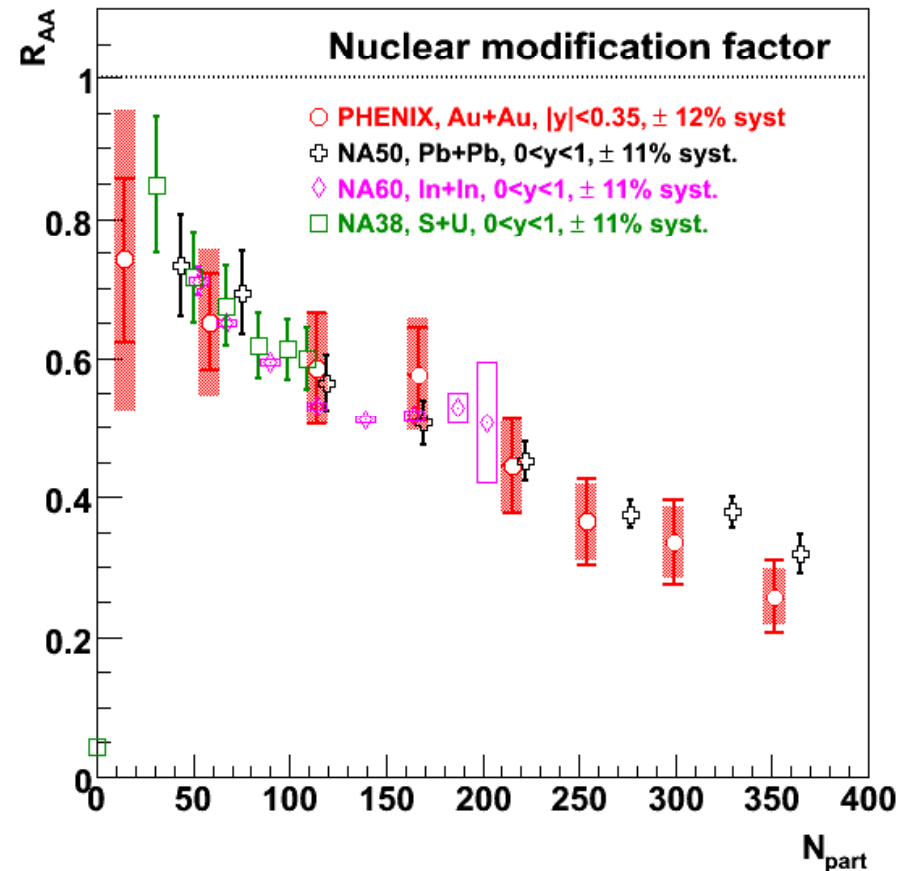
---

1.  $R_{AA}(\text{RHIC}) \approx R_{AA}(\text{SPS}) @ \gamma \approx 0$
2.  $R_{AA}(\gamma \approx 1.7) < R_{AA}(\gamma \approx 0)$

$$R_{\text{AuAu}} (y \approx 0 \text{ in PHENIX}) \approx R_{\text{PbPb}} (@ \text{ SPS})$$

- Midrapidity  $R_{\text{AA}}$  looks surprisingly similar, while there are obvious differences:

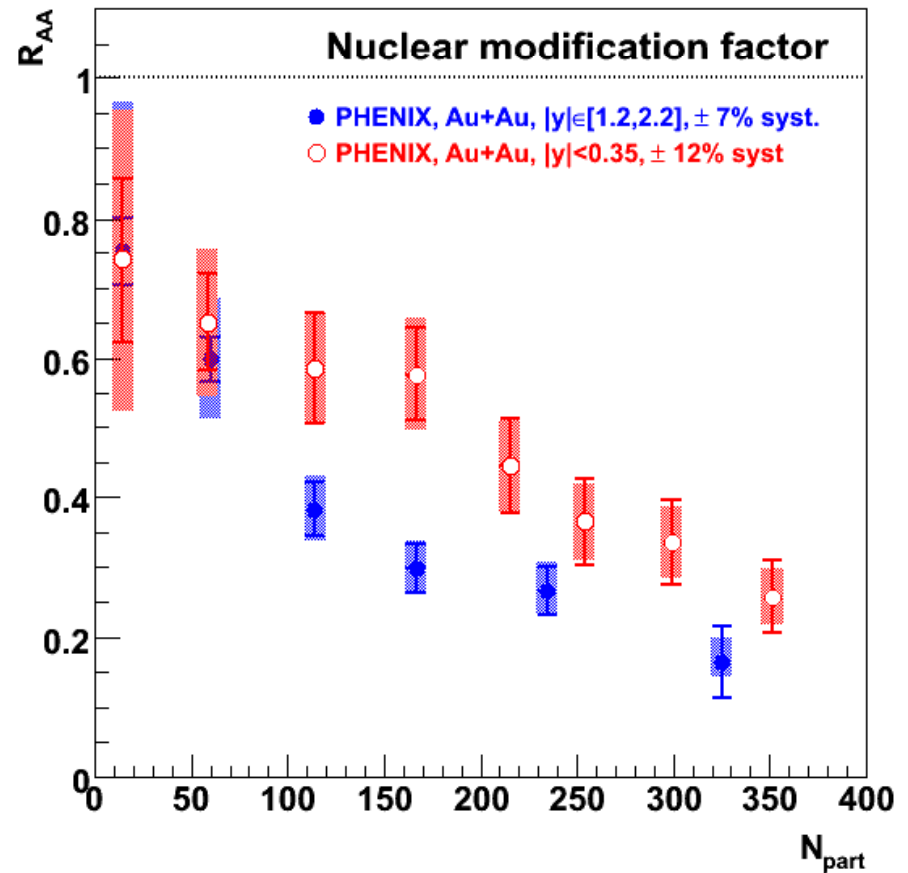
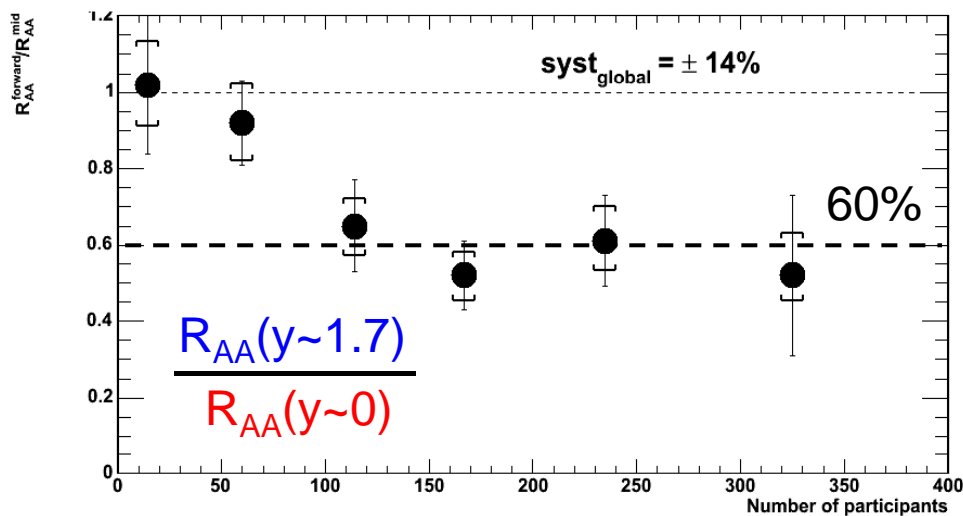
- At a given  $N_{\text{part}}$ , different energy densities...
- Cold nuclear matter effects ( $x_{\text{Bjorken}}, \sigma_{\text{abs}} \dots$ )
- ...



PHENIX, PRL98 (2007) 232301  
SPS from Scomarini @ QM06

# $R_{\text{AuAu}}(y \approx 1.7) < R_{\text{AuAu}}(y \approx 0)$ in PHENIX

- @ RHIC, more  $J/\psi$  suppression at forward rapidity !
- While energy density should be smaller...



PHENIX, PRL98 (2007) 232301



For the hot stuff, see  
Cesar Luis da Silva's  
talk on Wednesday

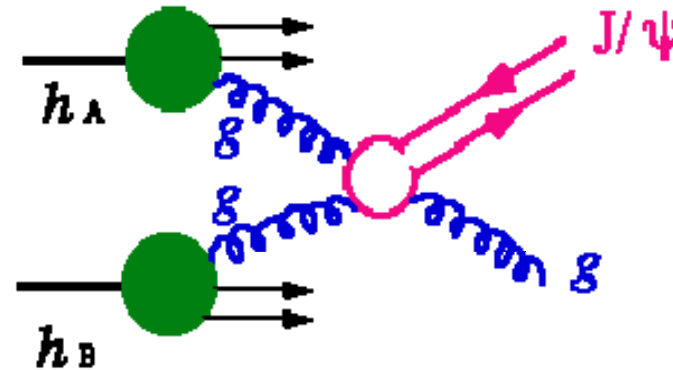
How much of this is due to  
normal nuclear matter ?

---

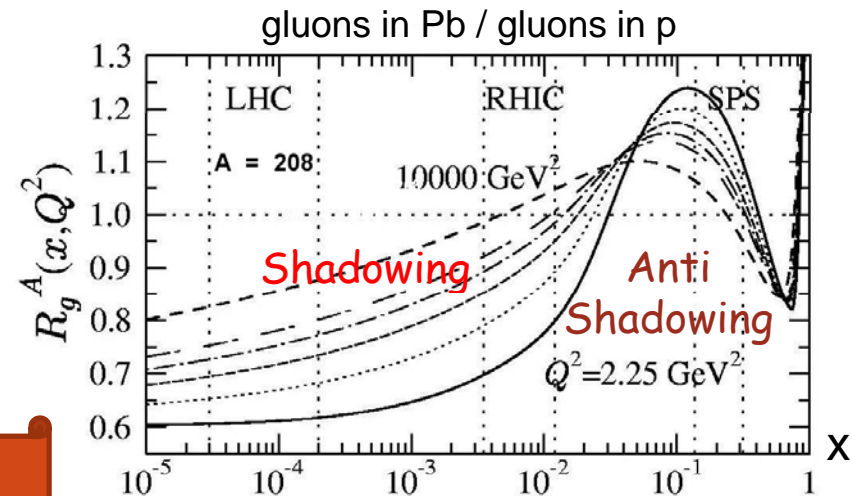
(what cannot claim anything about  
quark gluon plasma without first  
answering this question)

# Cold nuclear matter effects ?

- Many possible effects:
  - J/ $\psi$  (or  $c\bar{c}$ ) absorption/breakup
  - (Anti) shadowing  
(gluon saturation, CGC...)
  - Energy loss of initial parton
  - $p_T$  broadening (Cronin effect)
  - Complications from feed down  
 $\psi'$  &  $\chi_c$  ?
  - Intrinsic charm ?
  - Something else ?
- Absolute need for data !



An example of gluon shadowing prediction



Eskola, Kolhinen, Vogt  
NPA696 (2001) 729

# Nobody is perfect...

---

- ➔ @SPS: many pA ! High statistics ! But small kinematics ( $-0.1 < x_F < +0.1$ )
  - Nuclear absorption does a splendid job
- @FNAL: less pA... High statistics ! Large rapidity ( $x_F$ ) coverage... No AA...
  - Many cold nuclear effects needed!
- @HERAB: similar, negative  $x_F$  ( $-0.35$  to  $+0.15$ )
- ➔ @RHIC: only dAu, low statistics, but rapidity ( $-2.2$  to  $+2.2$ ) and centrality dependence
  - Absorption + (anti)shadowing

For the real stuff, see  
Carlos Lourenço's talk  
hereafter

## A snapshot of SPS

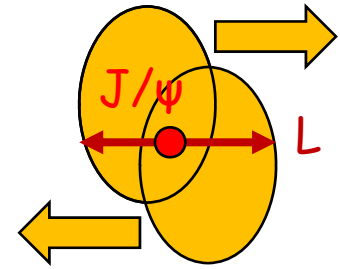
---

To first order, a simple and efficient  
description of nuclear matter effects

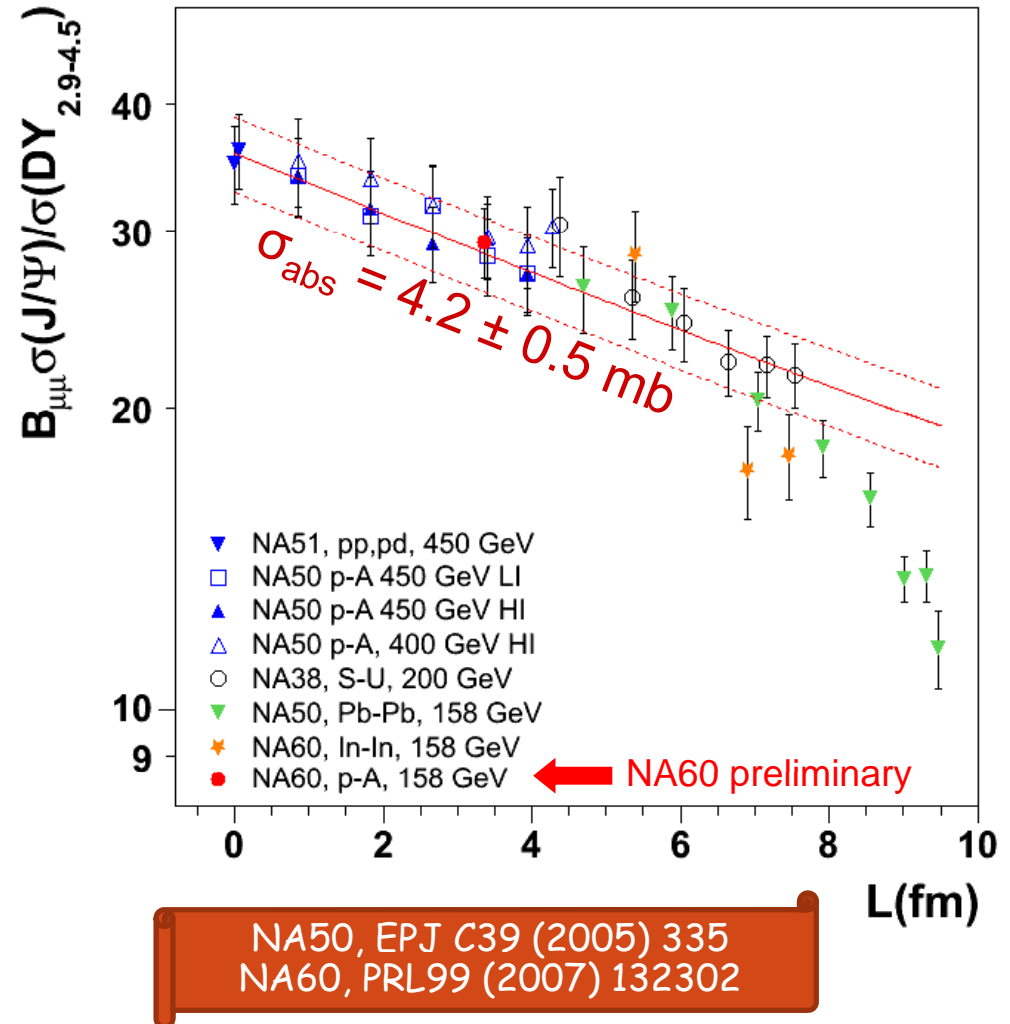
$$\sigma_{\text{abs}} = 4.2 \pm 0.5 \text{ mb}$$



# Cold and hot matters @ SPS



- Normal nuclear absorption alone does a splendid job describing pA, SU and peripheral InIn and PbPb:
  - (including one preliminary pA @ 158 GeV from NA60, final yet to come...)
- $\exp(-\sigma_{abs} \rho_0 L)$ 
  - L nuclear thickness
  - (or in Glauber model)
  - $\sigma_{abs} = 4.2 \pm 0.5 \text{ mb}$



$R_{dAu}$  rapidity dependence @ RHIC

---

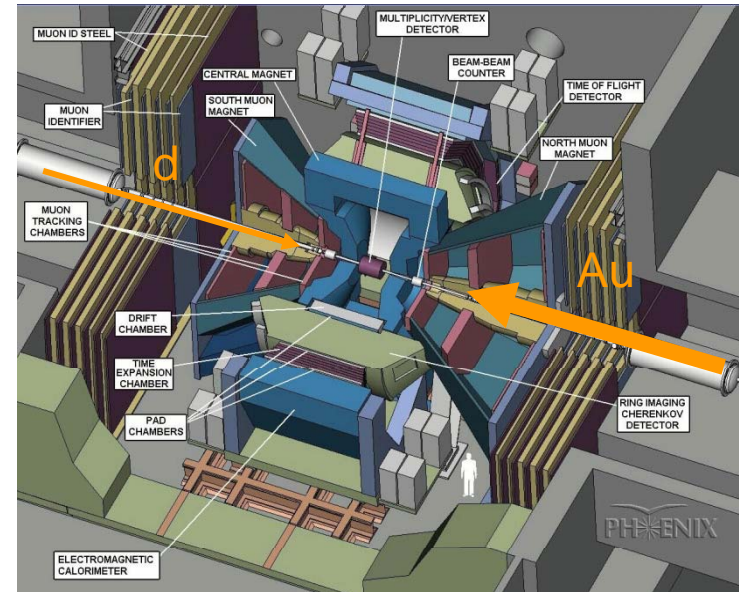
Shading light on shadows

Deuteron →

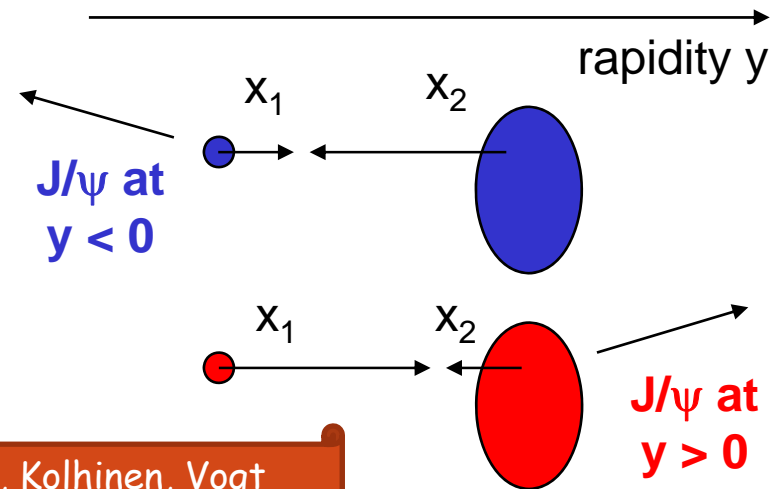
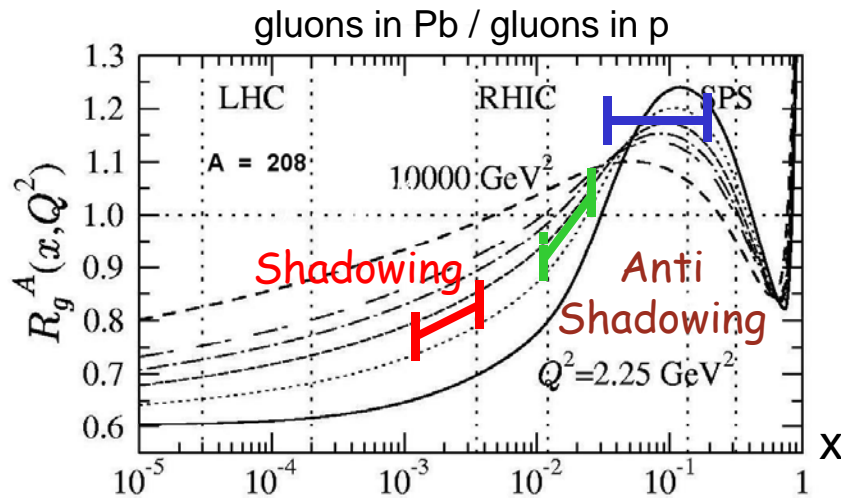
← Gold

- At RHIC,  $J/\psi$  mostly produced by gluon fusion, and thus sensitive to gluon's pdf
- In PHENIX, three rapidity ranges probe different momentum fraction of Au partons
  - South ( $y < -1.2$ ) : large  $x_2$  (in gold)  $\approx 0.05$  to  $0.14$  \*
  - Central ( $y \approx 0$ ) : intermediate  $x_2 \approx 0.011$  to  $0.022$  \*
  - North ( $y > 1.2$ ) : small  $x_2$  (in gold)  $\approx 0.002$  to  $0.005$  \*

\* Neglecting  $p_T$



An example of gluon shadowing prediction



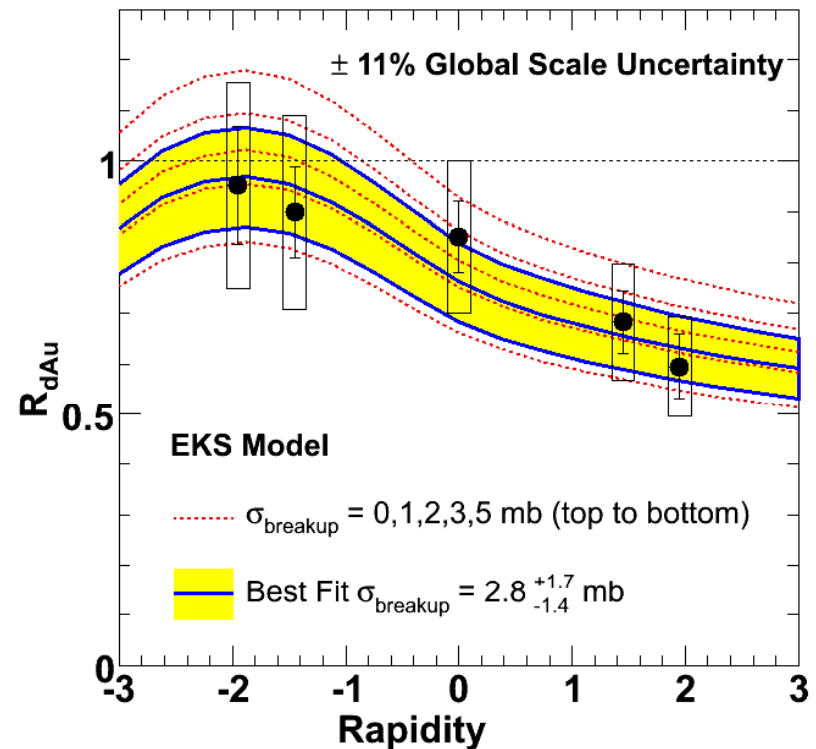
Eskola, Kolhinen, Vogt  
NPA696 (2001) 729

# $R_{dAu}(y)$ @ RHIC

- New analysis of run 3 RHIC data
  - Same p+p reference as Au+Au and Cu+Cu
  - Better (cancellation of) systematics
- Suppression at forward rapidity
  - Shadowing
- Assuming a shadowing scheme, adjust an absorption cross-section

## EKS shadowing

$$\sigma_{\text{abs}} = 2.8^{+1.7}_{-1.4} \text{ mb}$$

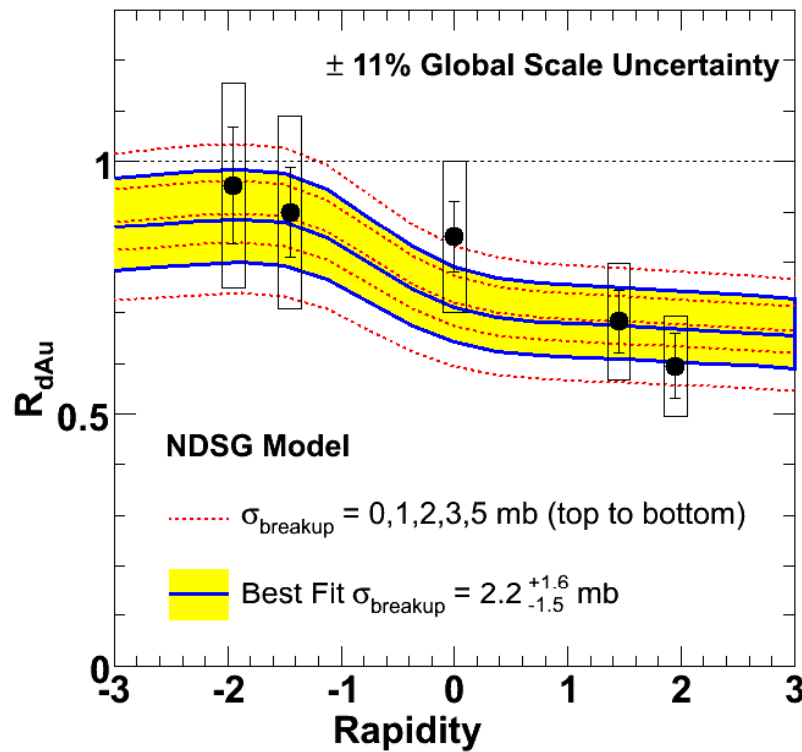


PHENIX, arxiv:0711.3917

$$R_{dAu}(y)$$

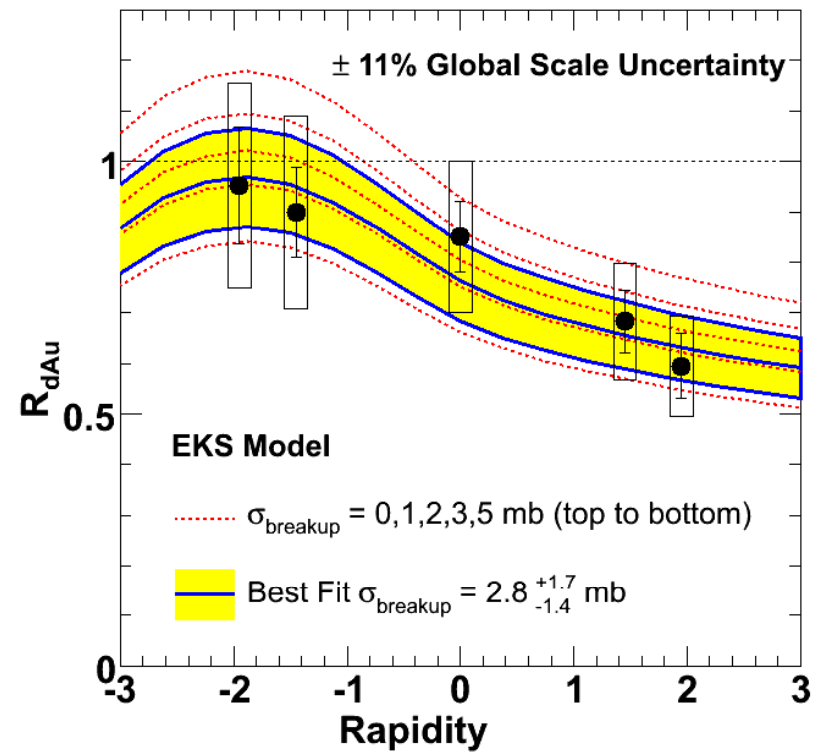
## NDSG shadowing

$$\sigma_{\text{abs}} = 2.2^{+1.8}_{-1.5} \text{ mb}$$



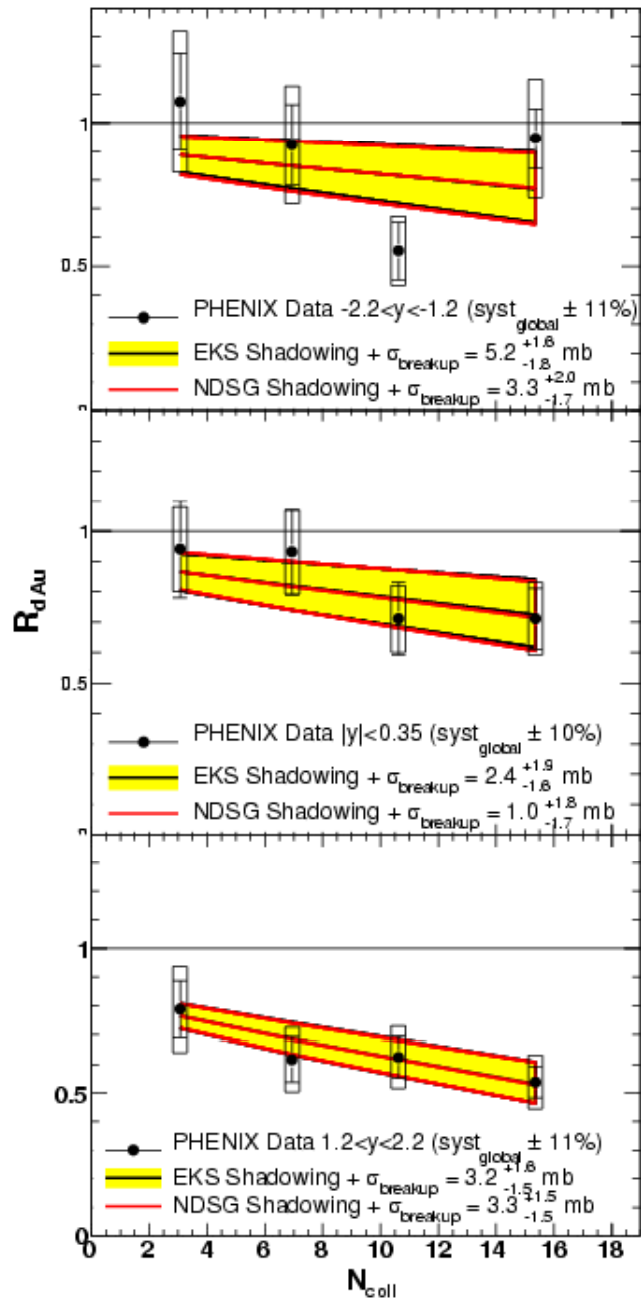
## EKS shadowing

$$\sigma_{\text{abs}} = 2.8^{+1.7}_{-1.4} \text{ mb}$$



PHENIX, arxiv:0711.3917

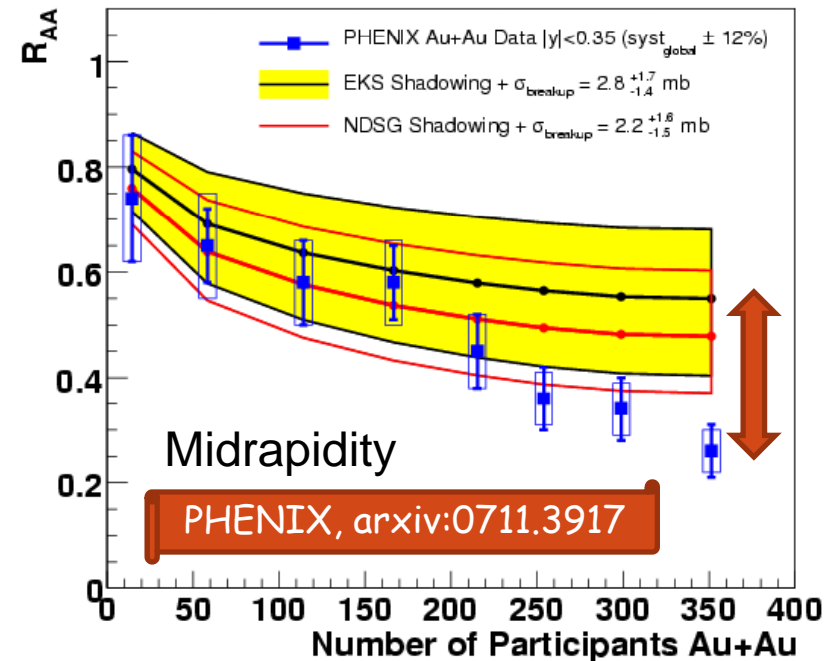
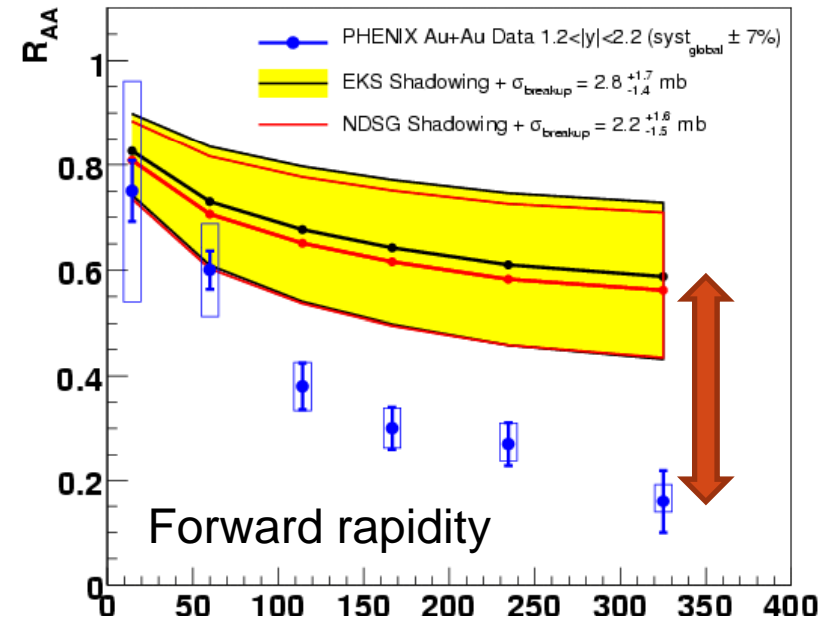
$$R_{dAu}(N_{coll}, y)$$



- Also tried to fit the centrality dependence
  - (assuming a inhomogeneous shadowing scheme)
- Consistent results within (large) uncertainties

$$R_{AA}(N_{part})$$

- Now extrapolate to AuAu collisions →
  - (Also available for CuCu)
  - Mid and forward are correlated through shadowing scheme
  - If you believe this shadowing, large anomalous suppression, larger at forward rapidity
- Net shadowing effect
  - EKS:  $(y=0) \approx (y=1.7)$
  - NDSG:  $(y=0) < (y=1.7)$

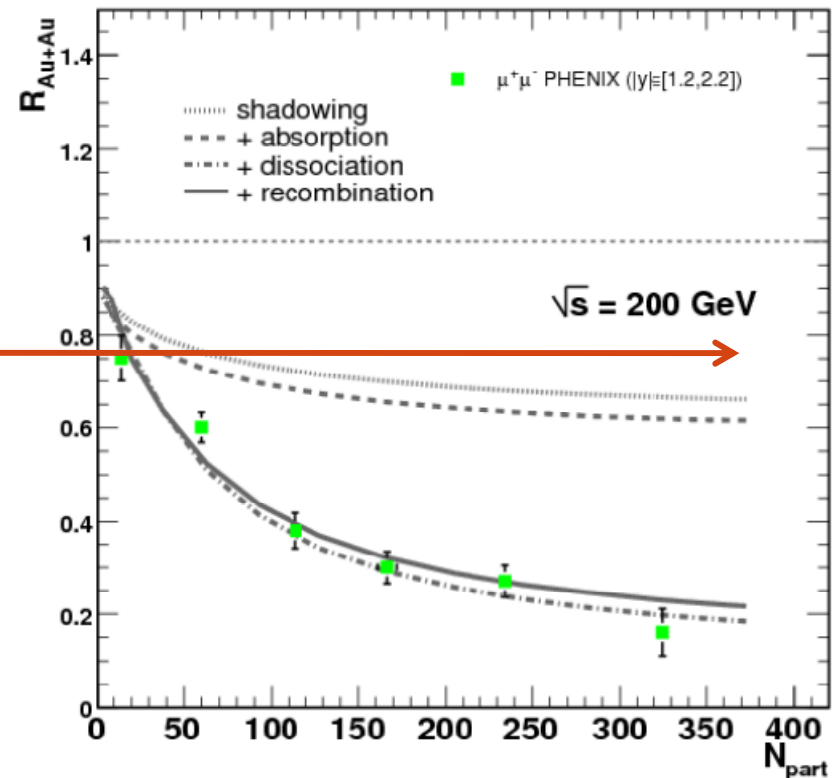
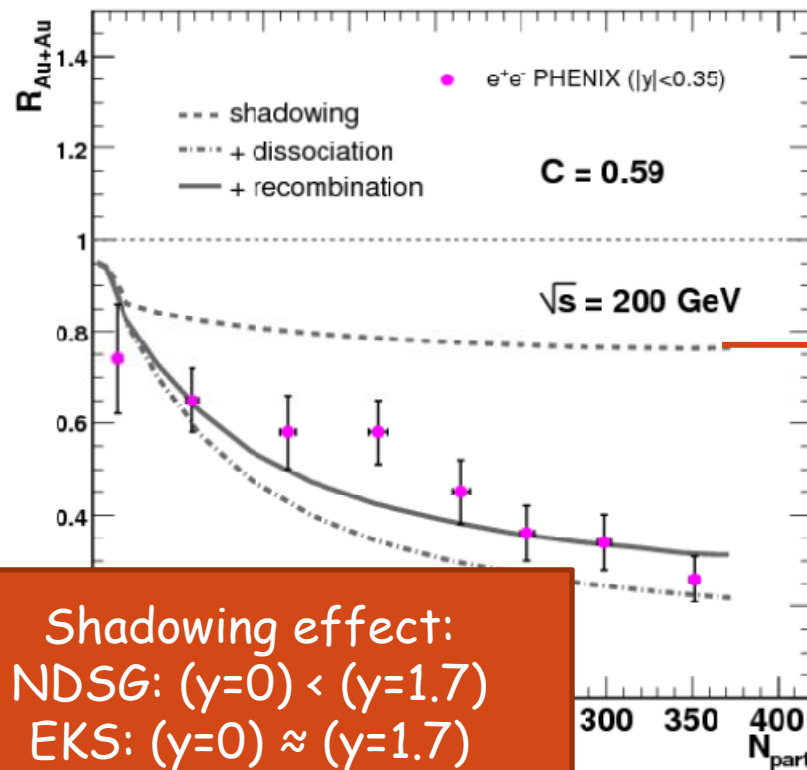


# Yet another shadowing scheme?

Capella et al, arXiv:0712.4331

Shadowing from Schwimmer  
multiple scattering :

+ E-p conservation  
+ regeneration

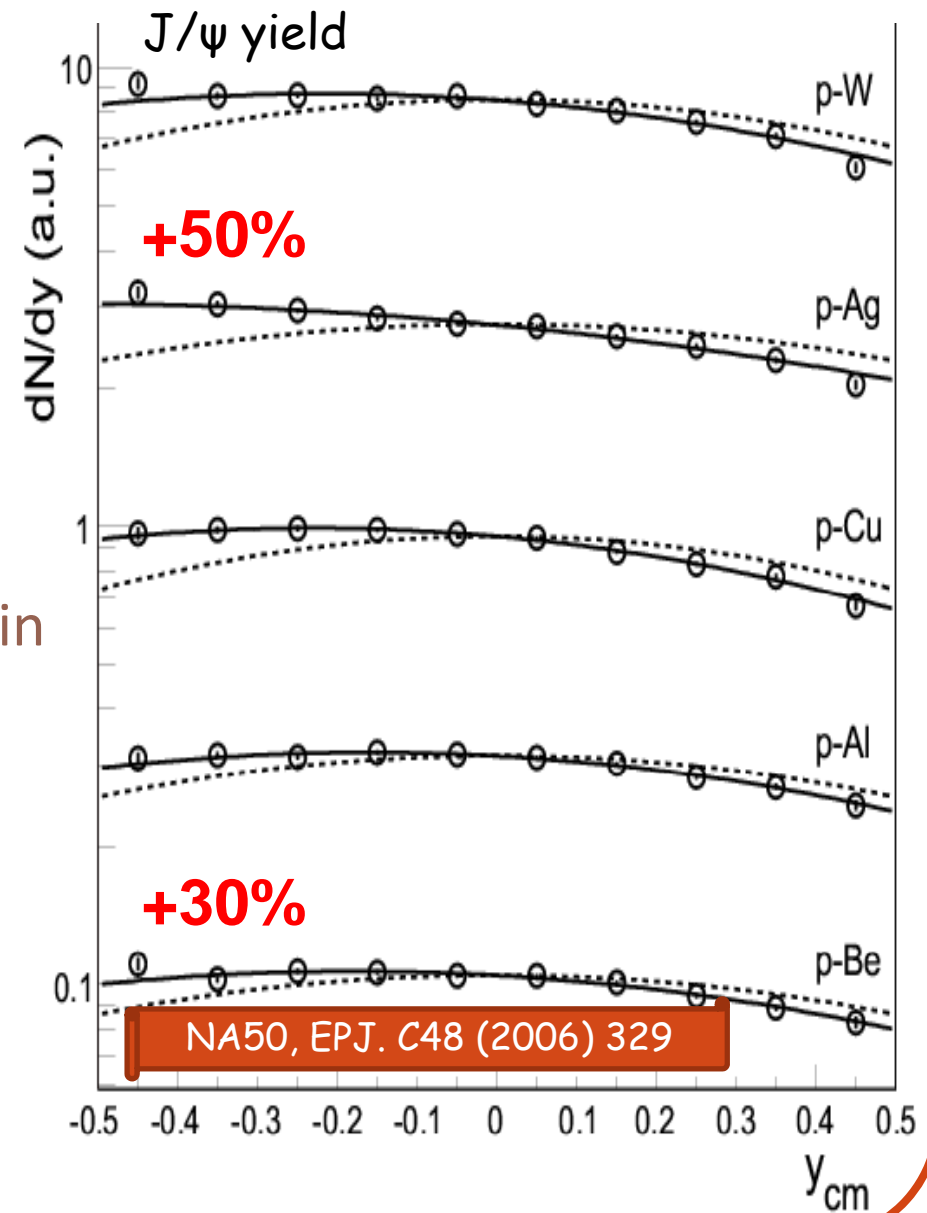


Shadowing effect:  
NDSG:  $(y=0) < (y=1.7)$   
EKS:  $(y=0) \approx (y=1.7)$   
Schwimmer:  $(y=0) > (y=1.7)$



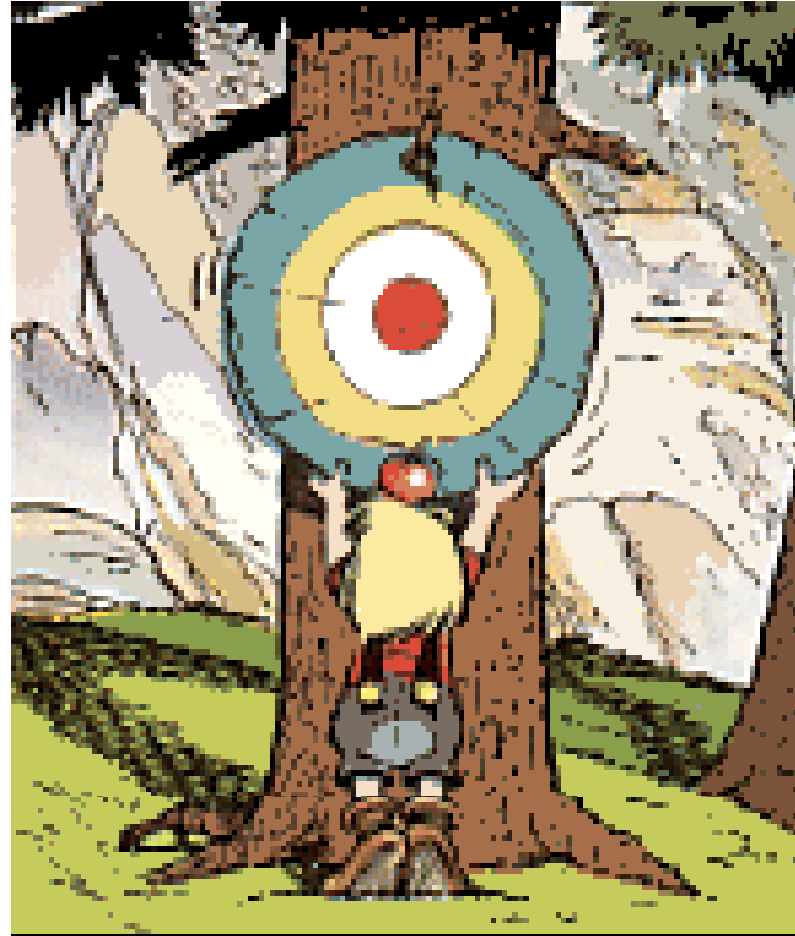
# Something odd @ SPS?

- Do we fully understand CNM @ SPS ?
- Not these surprising rapidity distribution asymmetries →
  - Variation of  $\approx 30$  to  $\approx 50\%$  in one unit of rapidity !
  - Seems large to be (anti)shadowing...
  - Not taken into account in CNM extrapolation...



# $R_{dAu}$ centrality dependence

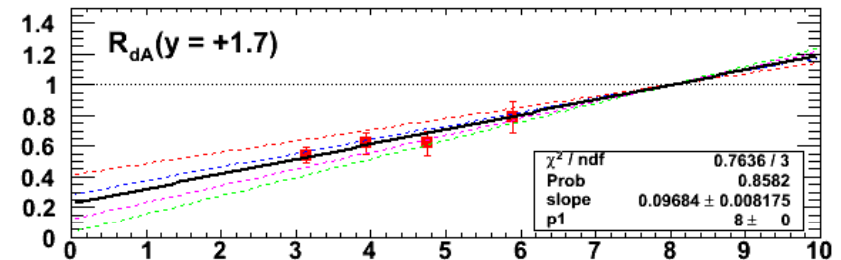
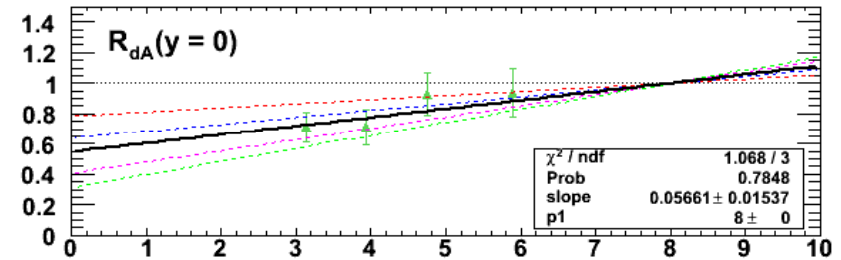
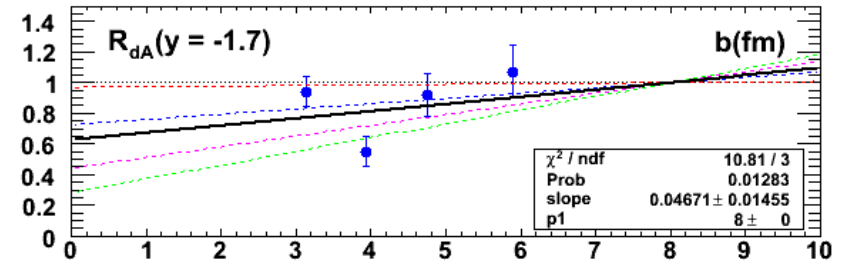
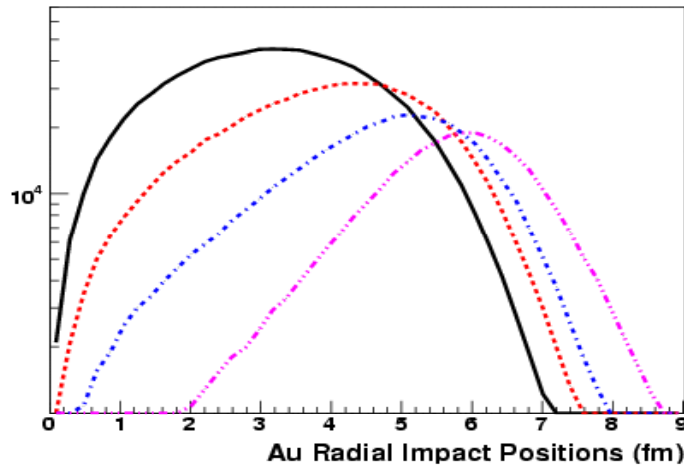
---



© Uderzo & Goscinny,  
Asterix chez les helvètes

# $R_{dAu}(b)$

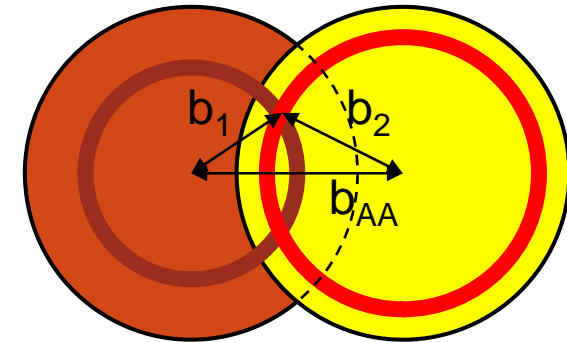
- Re-plot PHENIX  $R_{dA}$  vs local impact parameter  $b$  from Glauber model
- Phenomenological fit to  $R_{dA}(b) \rightarrow$ 
  - $R_{dA}(b) = A*(b-8)+1$  &  $R_{dA}(b>8\text{fm})=1$
  - (other shapes were tried)



RGdC, Quark Matter 06  
PHENIX, arxiv:0711.3917

# From dA to AA

- For a given A+A collision at  $b_{AA}$ , Glauber provides a set of N+N collisions occurring at  $b_i^1$  and  $b_i^2$



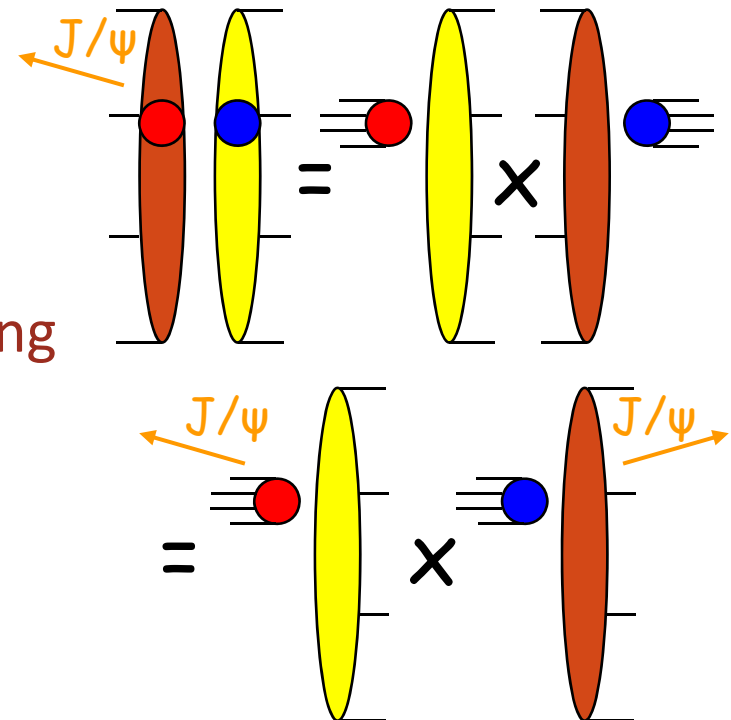
- One minimal assumption is rapidity factorization:  $R_{AA}(|y|, b_{AA}) =$

$$\sum_{\text{collisions}} [R_{dA}(-y, b_i^1) \times R_{dA}(+y, b_i^2)] / N_{\text{coll}}$$

- Works (at least) for absorption & shadowing since production

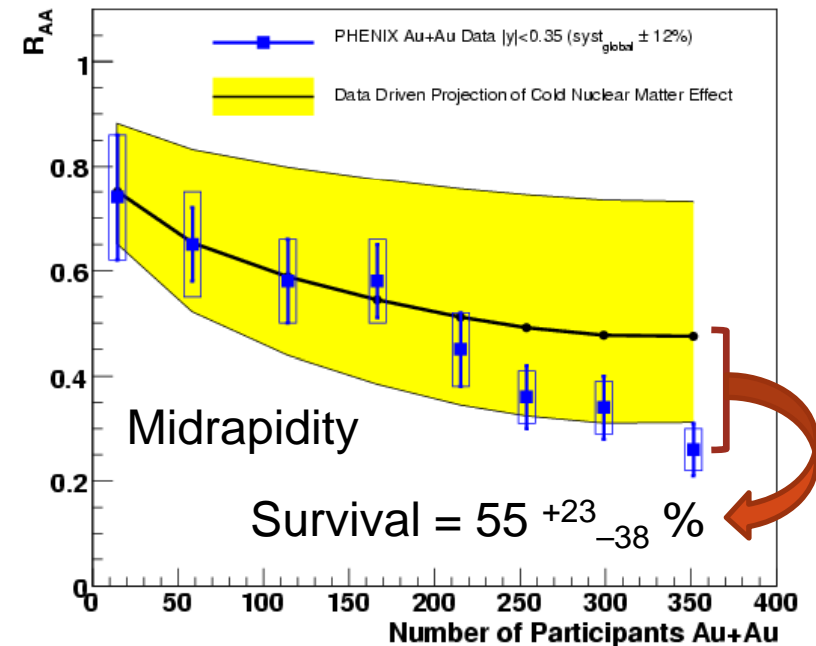
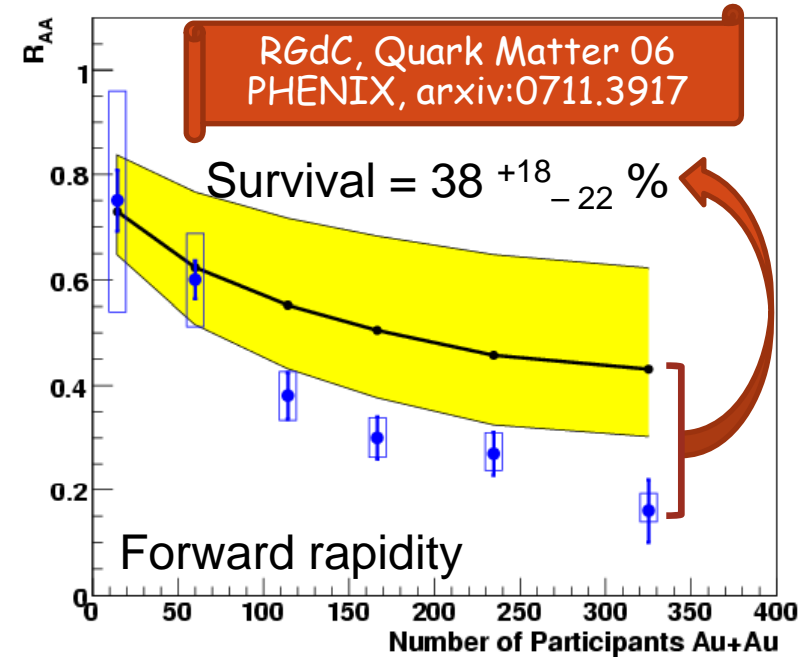
$$\sim \text{pdf1} \times \text{pdf2} \times \exp -\rho\sigma(L_1+L_2)$$

RGdC, QM06, hep-ph/0701222



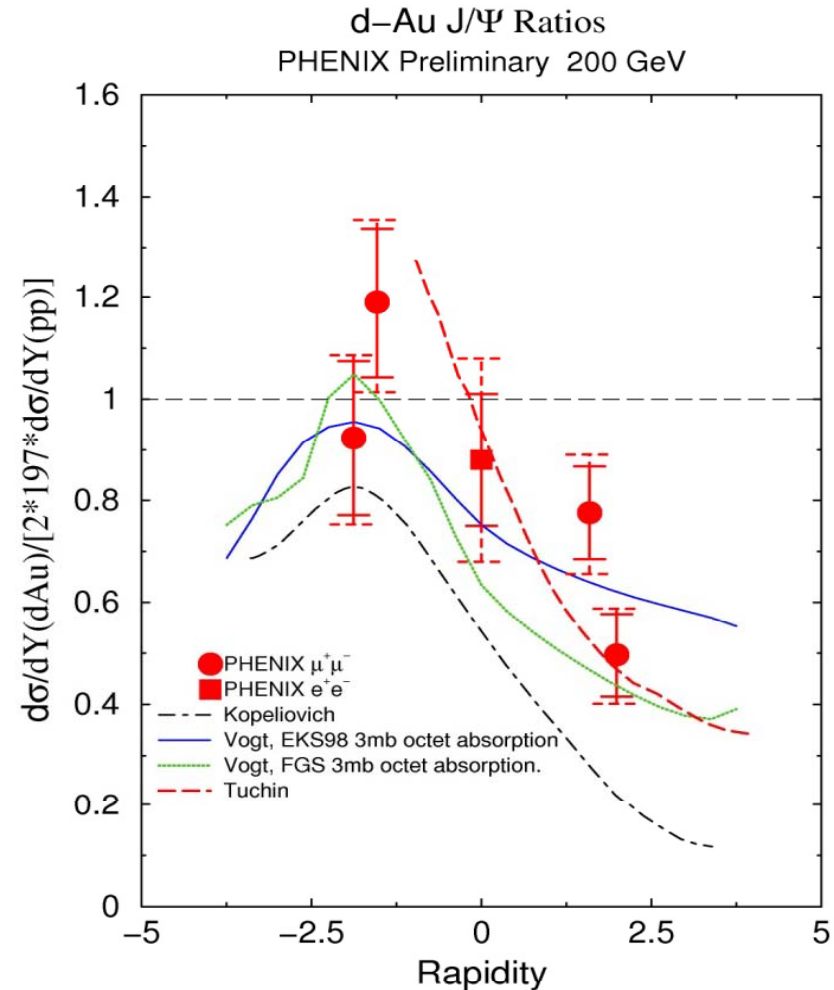
$$R_{AA}(N_{part})$$

- Pros and cons:
  - No shadowing scheme nor absorption scheme
  - Mid and forward are not correlated, less model dependent
  - Larger uncertainties (esp.  $y \approx 0$ )
  - No dCu, so no CuCu
- Anomalous suppression at least at forward rapidity!
- Anomalous suppression could be identical at midrapidity



# Unaccounted cold effects ?

- Could  $R_{dA}(-y) \times R_{dA}(+y)$  factorization be wrong?
- Yes, in case of strong saturation...
- dAu computation  $\rightarrow$
- AuAu computation underway... But:
  - Is saturation at play beyond traditional shadowing at  $x_1 \approx 0.003$ ?
  - How to describe  $x_2 \approx 0.1$ ?

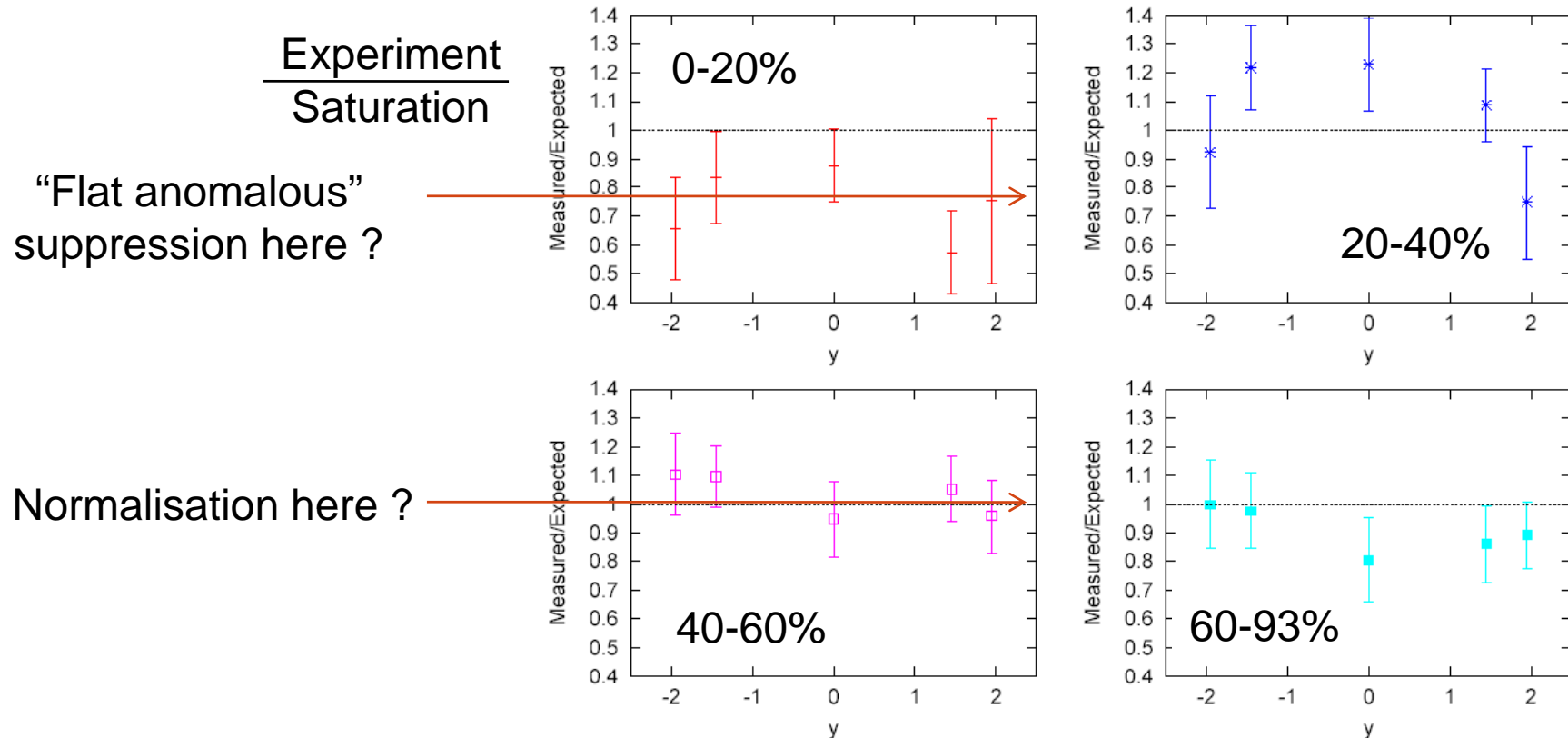


Tuchin, hep-ph/0504133

# Unaccounted cold effects ?

M. Nardi, QM08

- Saturation could suppress forward  $J/\psi$  in AuAu
- First numerical estimate, work in progress...



$R_{dAu}$  transverse momentum

---

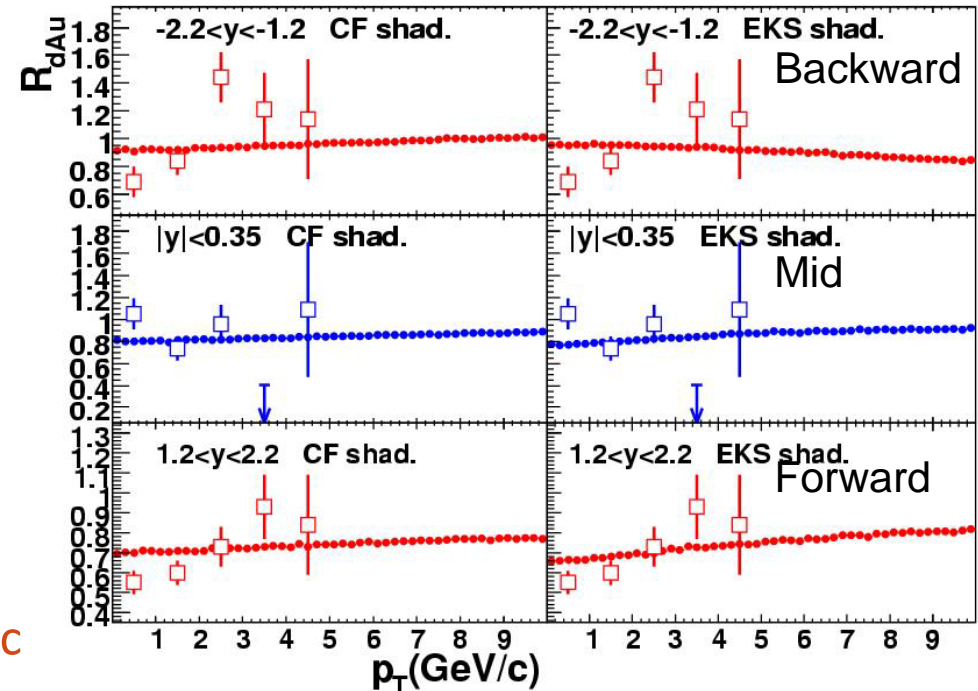


# $R_{dAu}(p_T)$

- $R_{dAu}(p_T)$  not precise but tendency to raise...
- Potential reasons:

$$x_{1,2} = \frac{m_T}{\sqrt{s}} e^{\pm y}$$

- Raising  $x_{Bj}$  = less shadowing
  - 0.02 to 0.05 from 0 to 9 GeV/c
  - See discussion in →
- Also Cronin effect? Further  $p_T$  broadening?

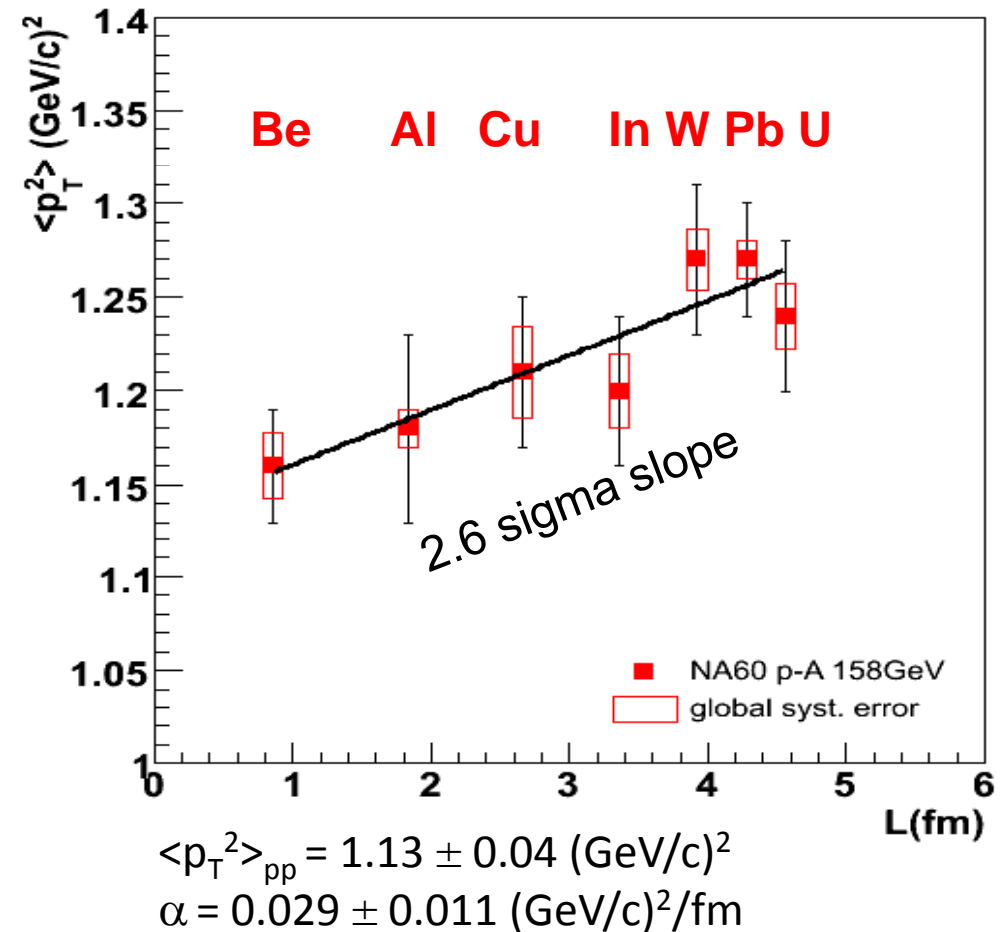


PHENIX, arxiv:0711.3917 compared to  
Ferreiro, Fleuret, Rakotozafindrabe,  
arxiv: 0801.4949

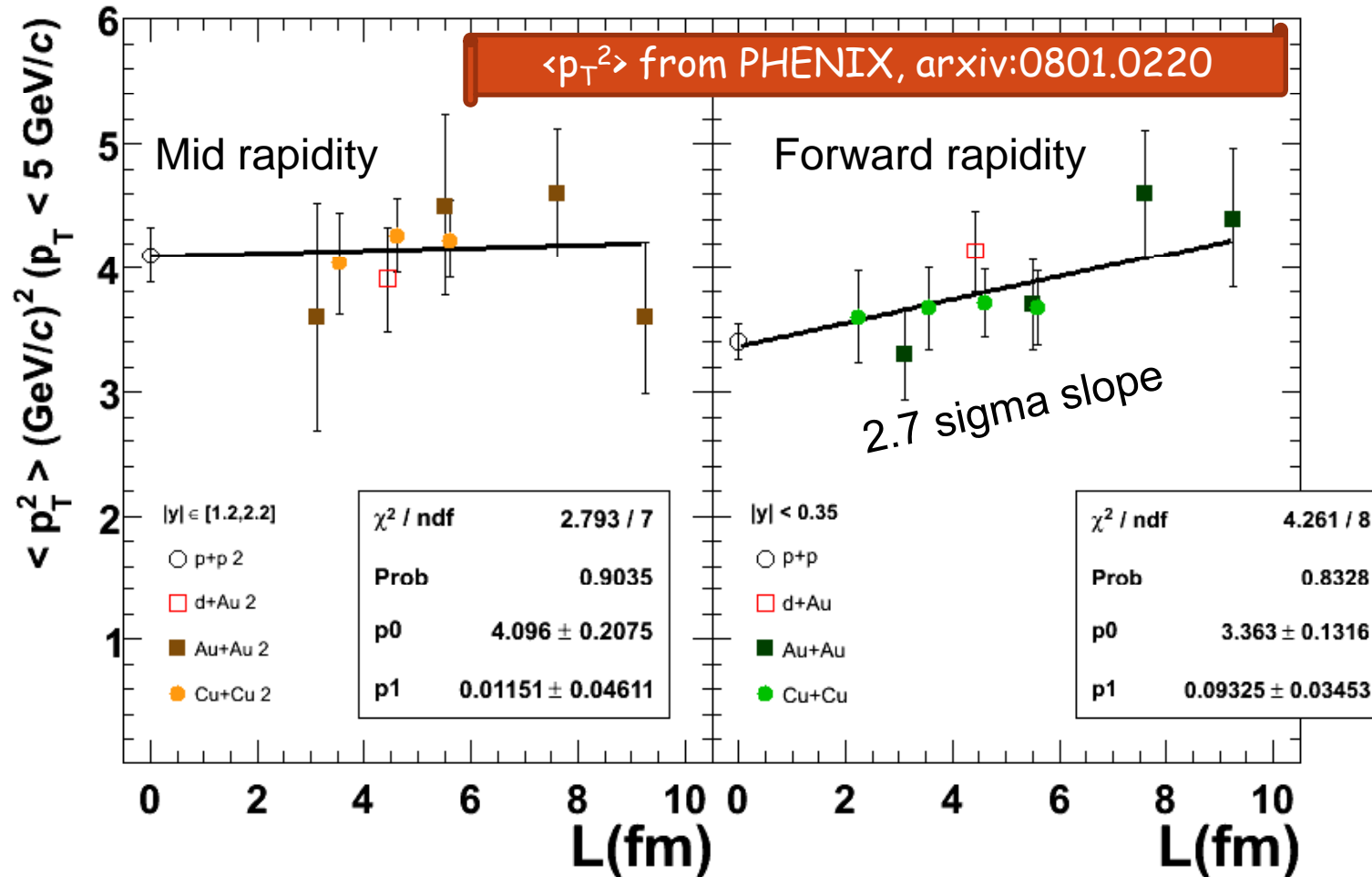
# $\langle p_T^2 \rangle$ vs $L =$ broadening @ SPS

R. Arnaldi, QM08

- Cronin should go like:
  - $\langle p_T^2 \rangle_{AB} = \langle p_T^2 \rangle_{pp} + \alpha \times L$
- Different scaling in pA and AA collisions reported at QM08
  - Something else going on in AA?
  - High  $p_T$  J/ $\psi$  escape?
- Anyway...



# $p_T$ broadening @ RHIC ?



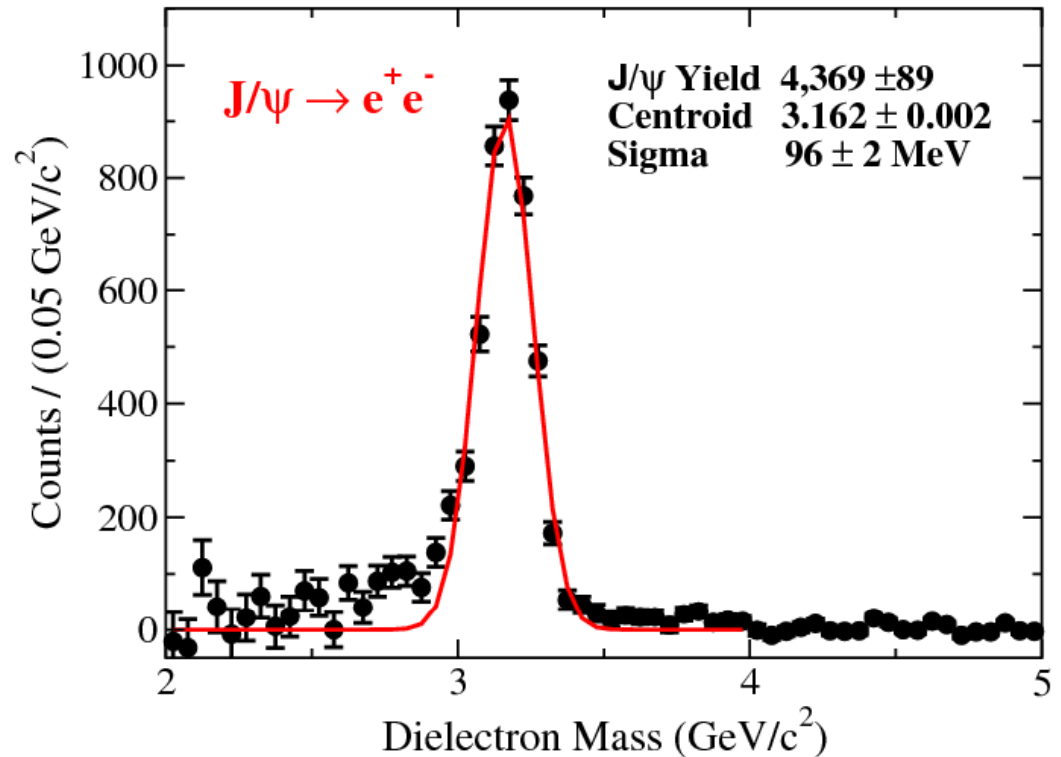
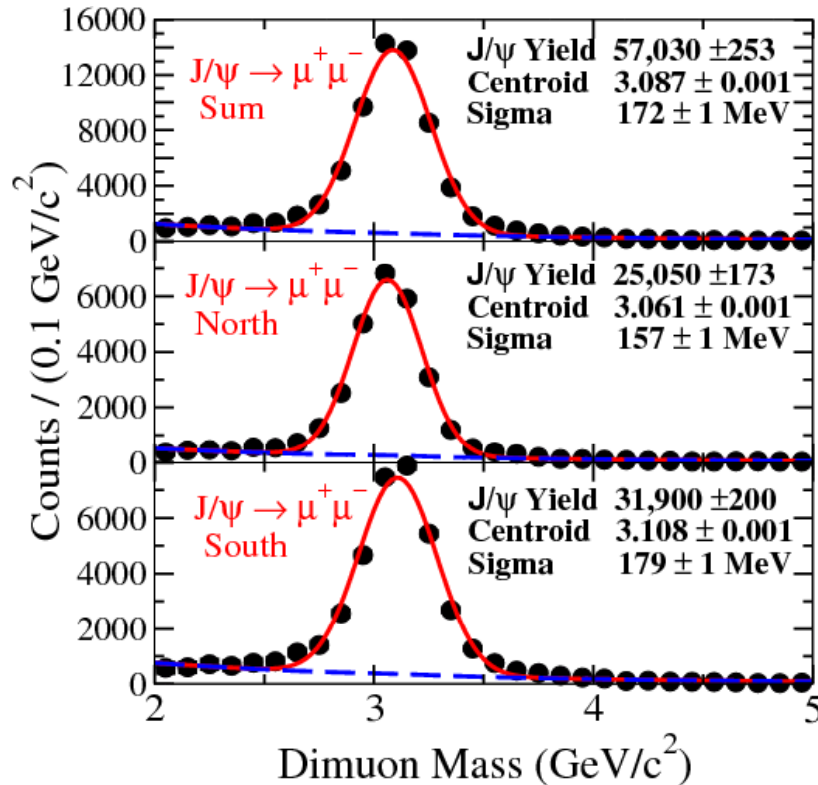
$\alpha = 0.011 \pm 0.046 \text{ (GeV/c)}^2/\text{fm}$ 
← compatible →
 $\alpha = 0.093 \pm 0.034 \text{ (GeV/c)}^2/\text{fm}$

## Three existing cold scenarios could justify the rapidity anomalous dependence

---

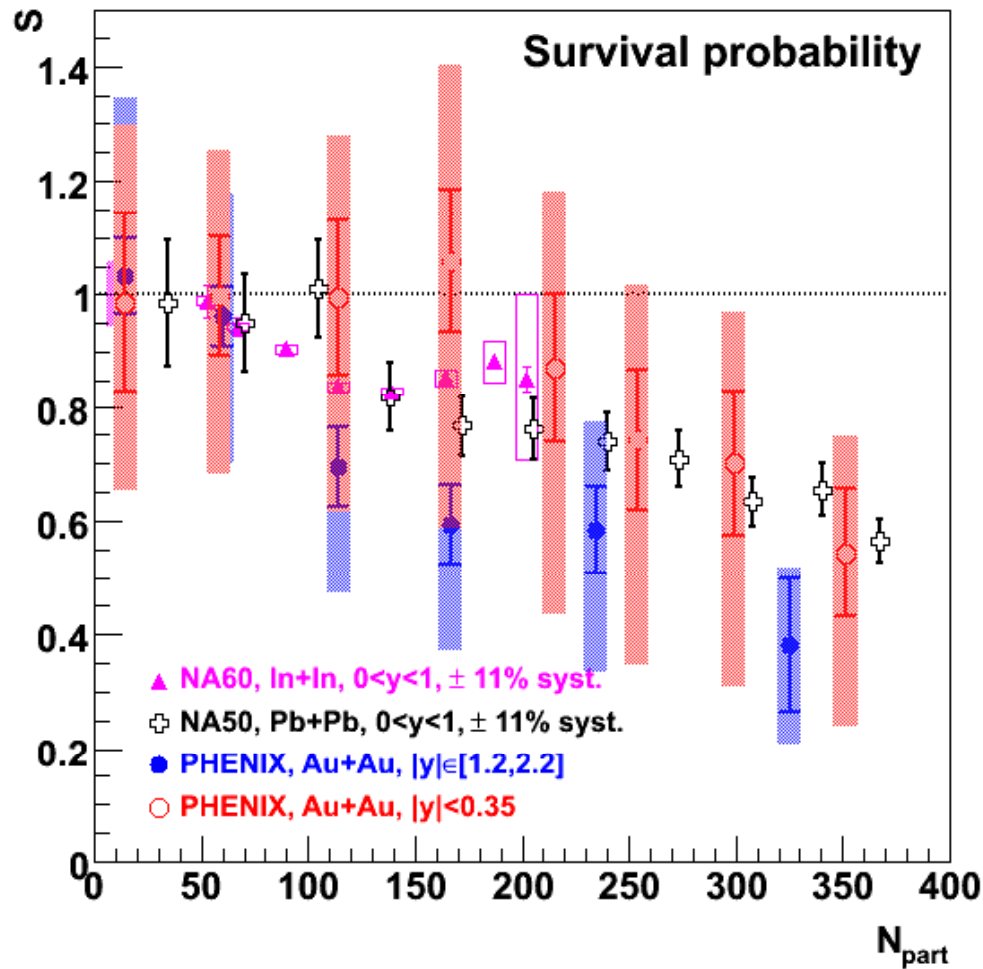
1. The data itself (dAu data driven method)
2. A given shadowing scheme (Schwimmer)
3. Color glass condensate (work in progress)  
[however very related]

Hopefully, RHIC run 8 dAu  $\approx$  30 x run 3 !



- Let's wait for this run analysis before to say more about cold matter (and derive decent survival probabilities)

# Indecent conclusion...



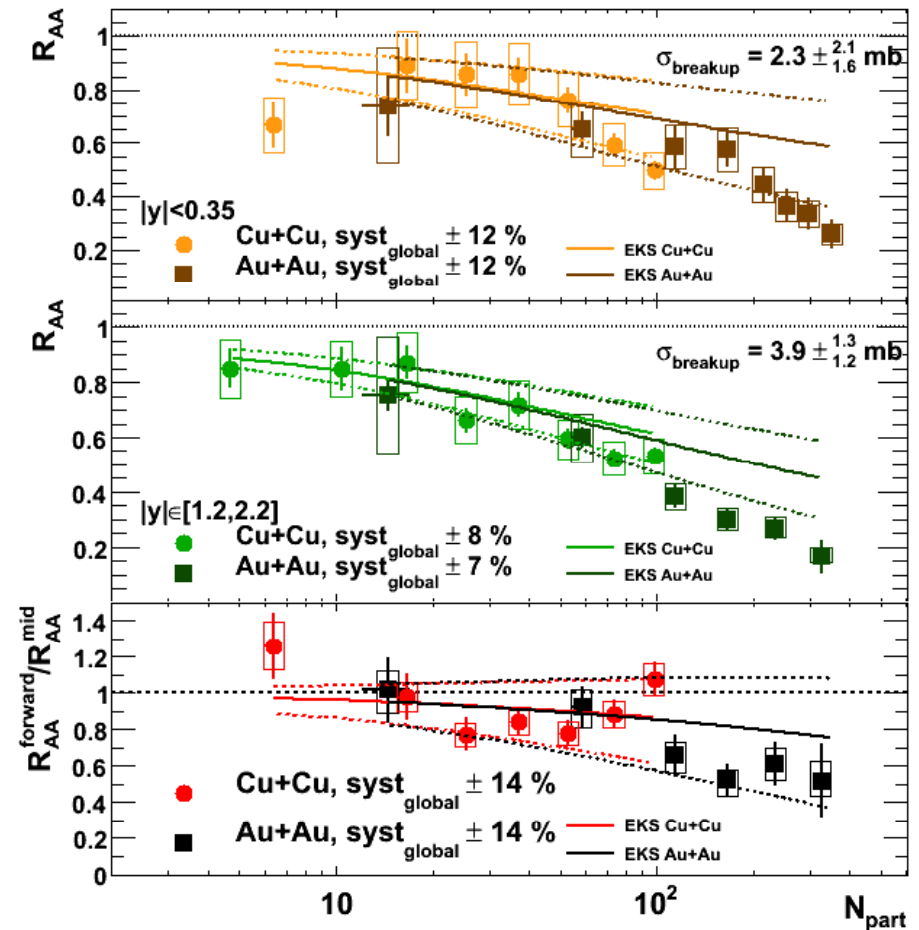
PHENIX, PRL98 (2007) 232301  
divided by  
PHENIX, arxiv:0711.3917  
(data driven method)

- RAA / CNM from data driven method
- Anomalous suppression could :
  - be the same at forward and midrapidity 😊
  - And different between SPS and RHIC 😊

That's all folks

# CuCu data

- Fitting an effective break-up cross section (depending on  $y$ ) and extrapolate to CuCu and AuAu...



PHENIX, arxiv:0801.0220

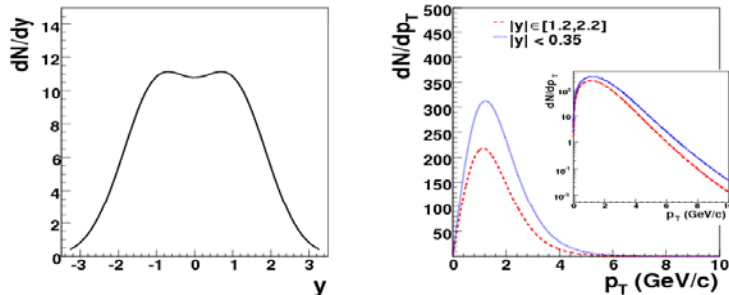


# Transverse momentum dependence of J/Ψ shadowing effects

arXiv:0801.4949

## Method:

J/Ψ  $p_T$  and rapidity extracted from PHENIX 200 GeV proton-proton data  
 Phys. Rev. Lett.98, 232002 (2007)



$$x_{1,2} = \frac{m_T}{\sqrt{s_{NN}}} e^{\pm y} < 1$$

with  $m_T = \sqrt{m^2 + p_T^2}$

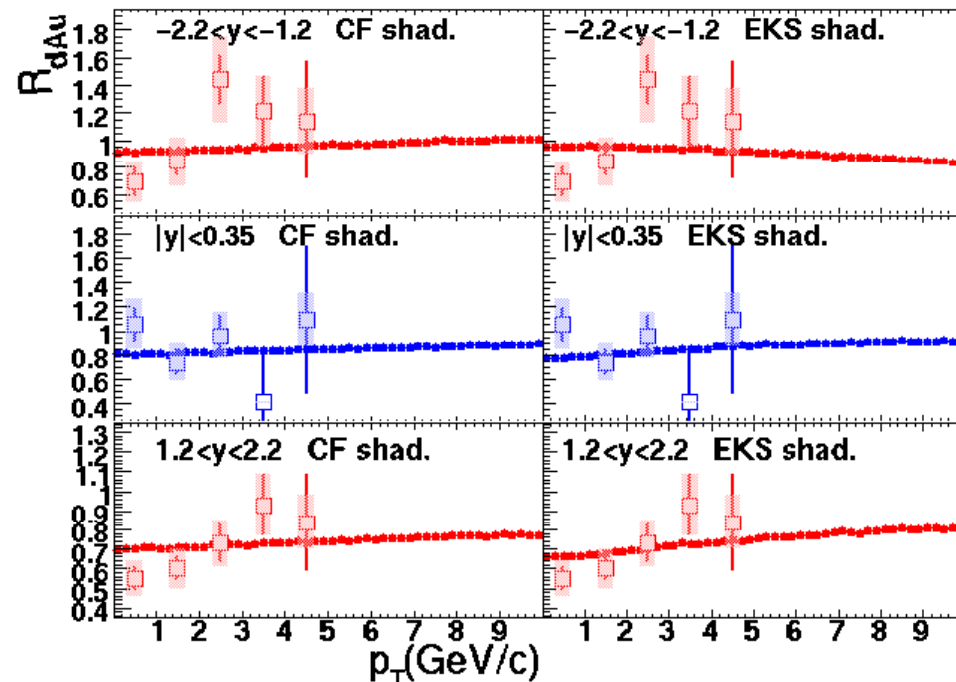
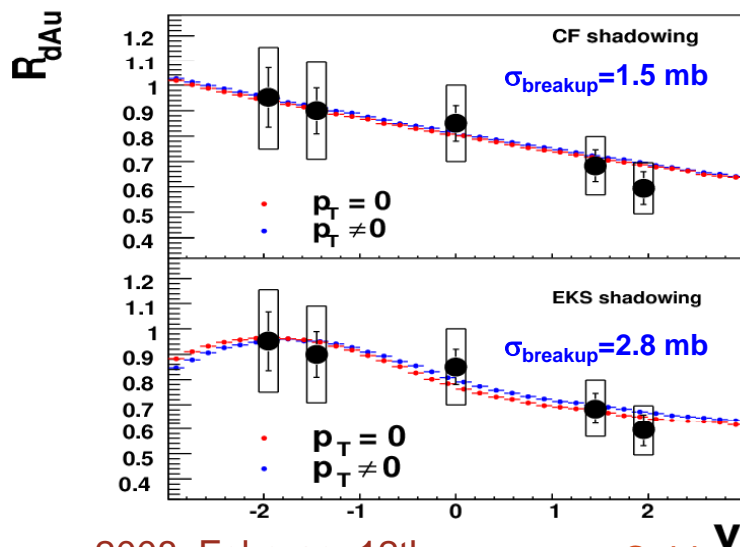
Glauber  
 Monte Carlo

Test 2 shadowing models :

- CF : Eur. Phys. J. C42, 419 (2005)
- EKS : Eur. Phys. J. C9, 61 (1999)

## Results:

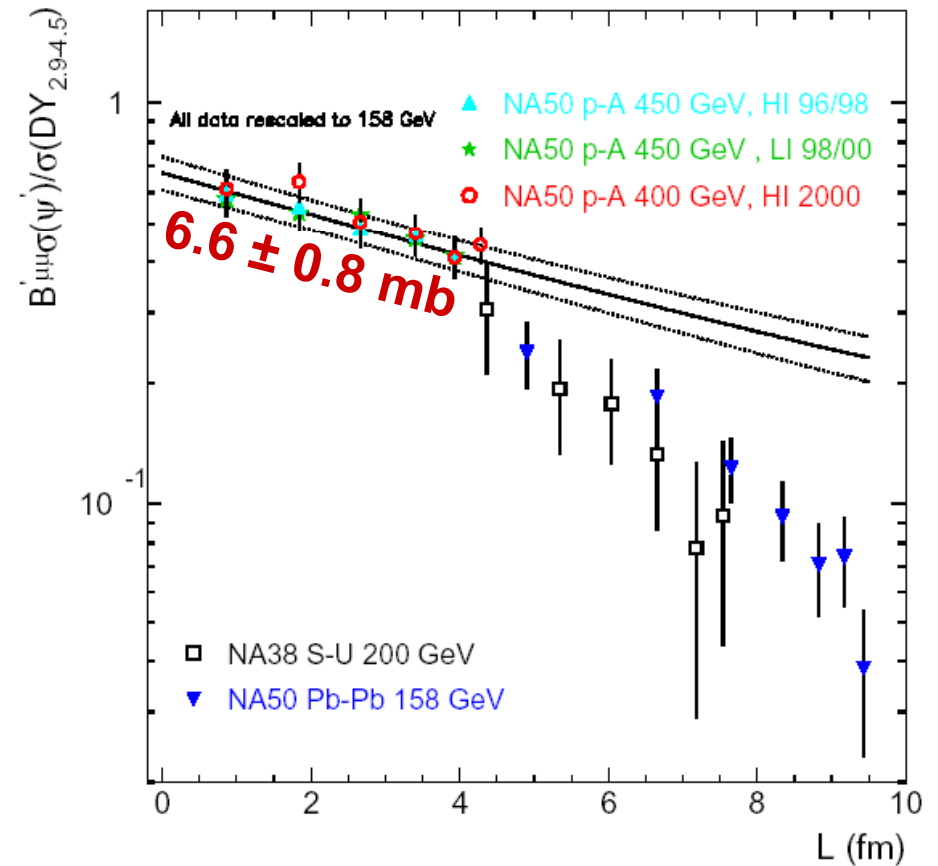
Expectations for 200 GeV d+Au data  
 Comparing with PHENIX 200 GeV d+Au data  
 arXiv:0711.3917



(E.G. Ferreira, F. Fleuret, A. Rakotozafindrabe)

**$R_{dAu}$  vs.  $p_T$  gives new constraints on shadowing models**

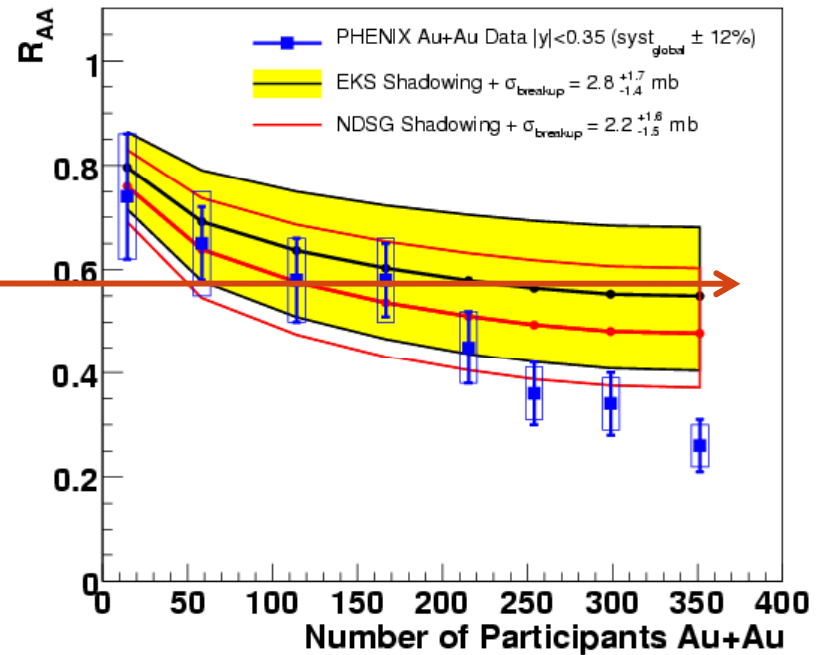
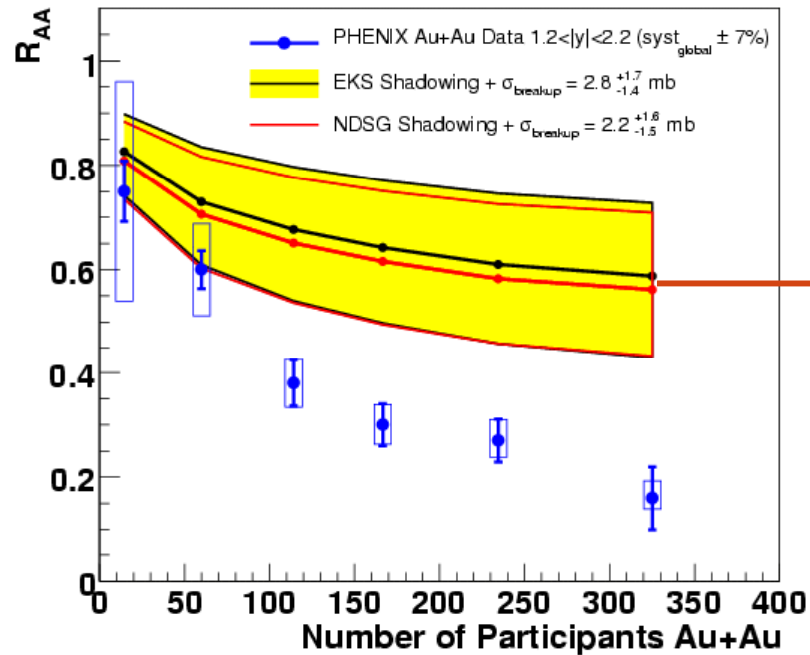
# $\Psi'$ absorption at SPS



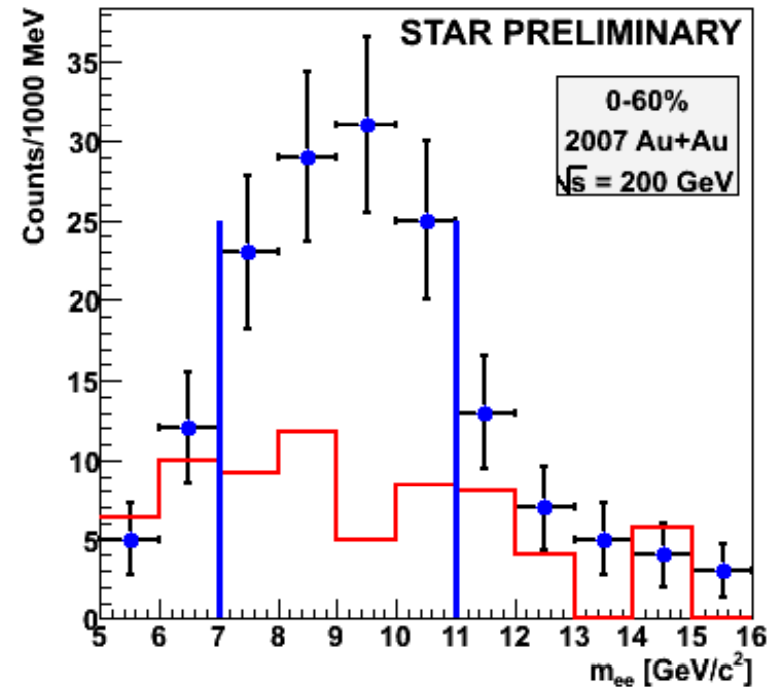
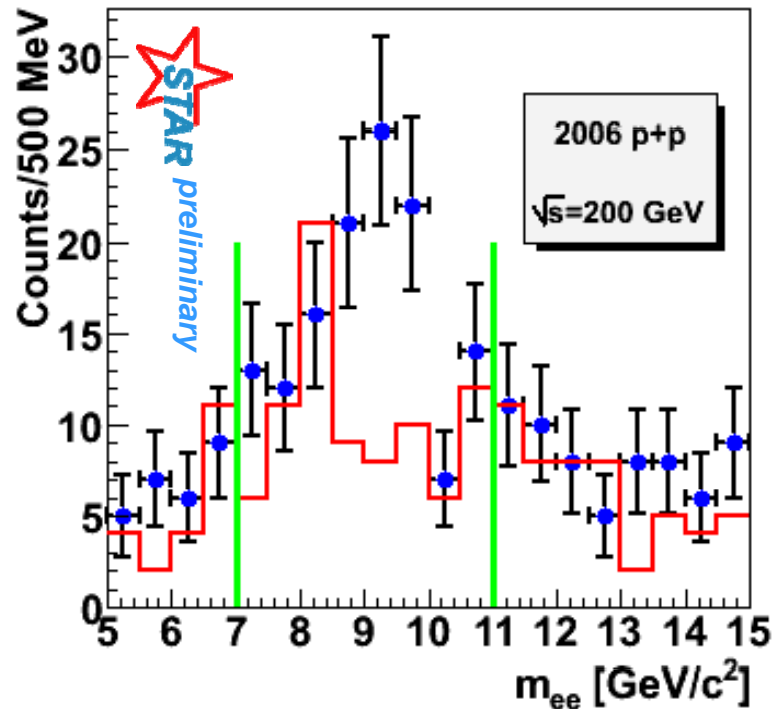
# Face to face

EKS ( $y=0$ )  $\approx$  ( $y=1.7$ )

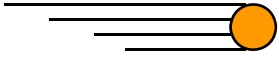
NDSG ( $y=0$ )  $<$  NDSG ( $y=1.7$ )



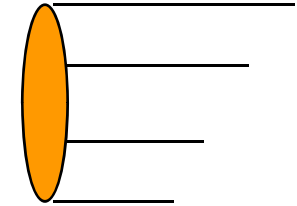
# STAR upsilon's



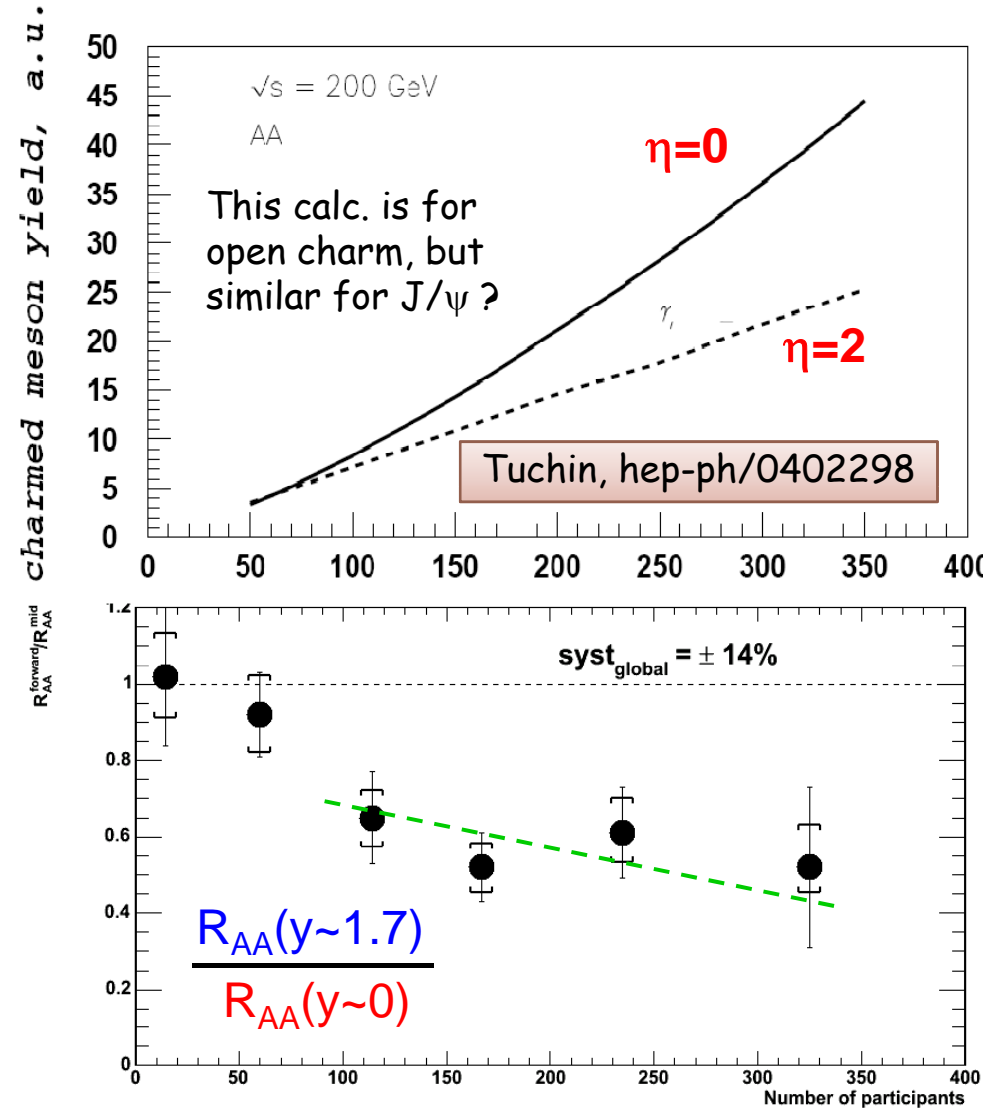
- Proof of principle: dozens of  $Y$  in p+p & A+A!
- Nuclear modification factor to come soon
- Suffers less from cold matter ( $x=0.02$  to  $0.1$ =EKS antishadow)
  - (should be checked with run8 d+Au)
- Should measure (unseparated) excited states melting



# Unaccounted CNM ?

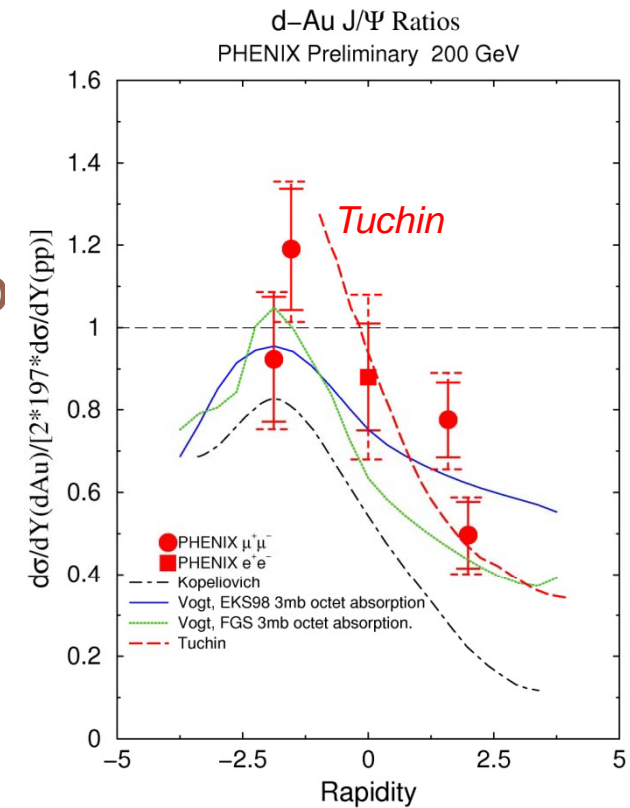
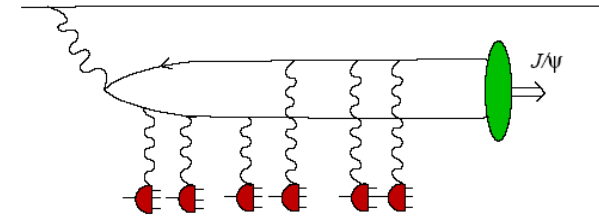


- Strong initial states effect ala color glass condensate ?
  - But they have to violate rapidity symmetrisation  $R_{AA}(|y|) = R_{dA}(-y) \times R_{dA}(+y)$
  - (otherwise taken into account in CNM extrapolation)
- Could this + sequential melting produce  $R_{AA}(y \sim 0)$  and  $R_{AA}(y \sim 1.7)$  ?
- Double ratio should drop...
- A possibility...



# Tuchin & Kharzeev

- Hard probes 2004
  - [hep-ph/0504133](https://arxiv.org/abs/hep-ph/0504133)
- Coherent production of charm (open or closed)
  - ( $y < 0$  production time too low to make computation)
  - Shadowing from CGC computation...



# Tuchin & Kharzeev...

+ absorption for  
SPS & fermilab

