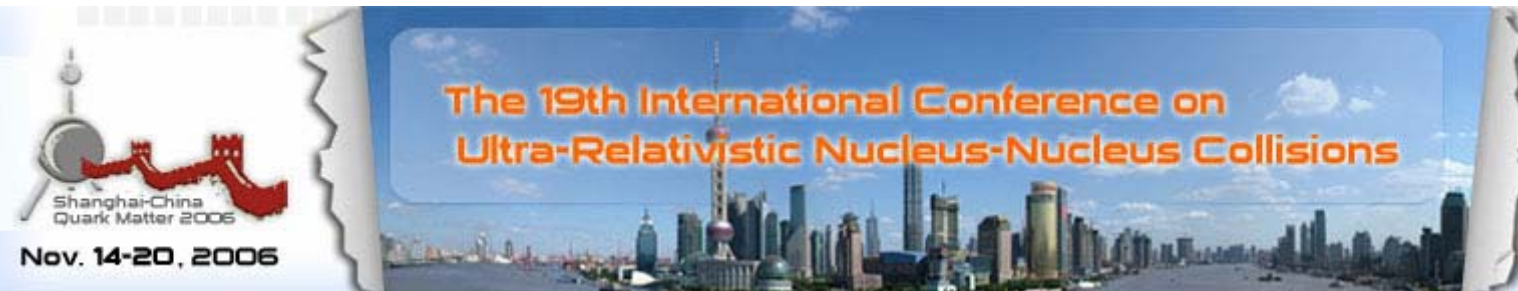


A d+Au data-driven prediction of cold nuclear matter (CNM) effects on J/ψ production in Au+Au collisions at RHIC

Raphaël Granier de Cassagnac
LLR - École polytechnique / IN2P3

(thanks to Jean Gosset
& Klaus Reygers)



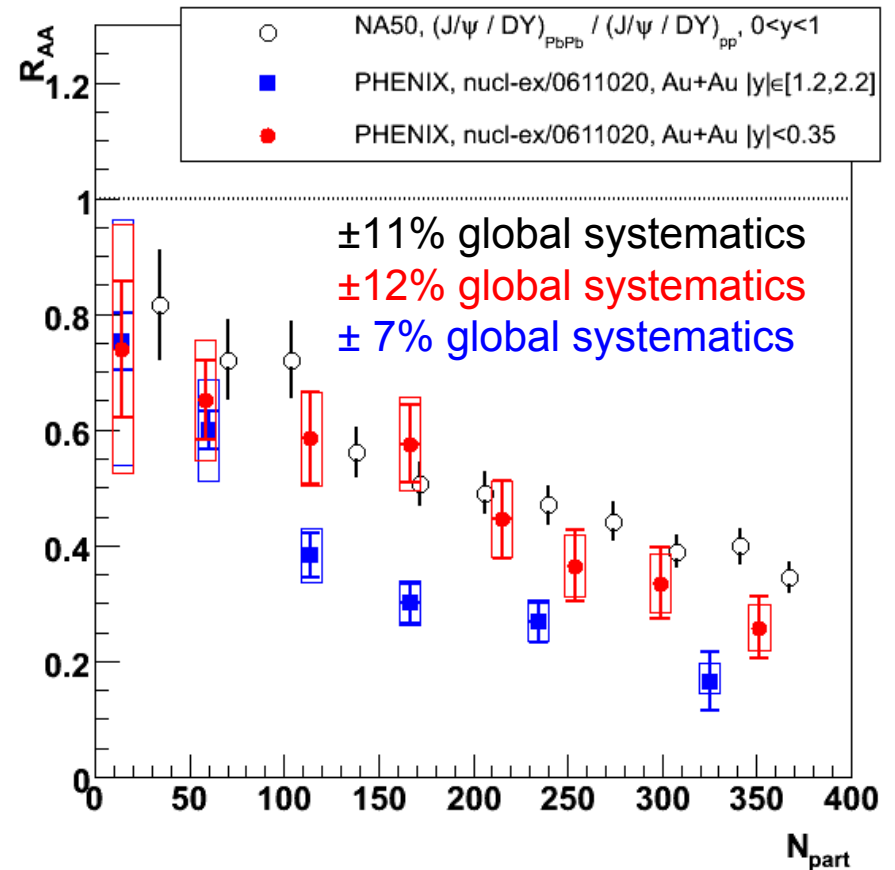
A red stylized signature or logo consisting of several connected, curved lines.

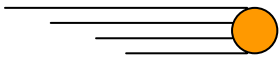
Motivation & outline

- Subtracting cold nuclear matter (CNM) effects is crucial to interpret J/ψ suppression

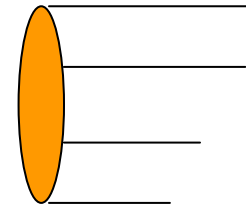
- How is it done @ SPS ?
- How is it done @ RHIC ?
- A new method for RHIC...

J/ψ nuclear modification factor





Cold nuclear matter @ SPS

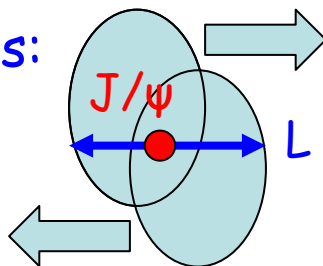


Normal nuclear absorption
alone does a splendid job
in describing pA, SU, and
peripheral PbPb & InIn...

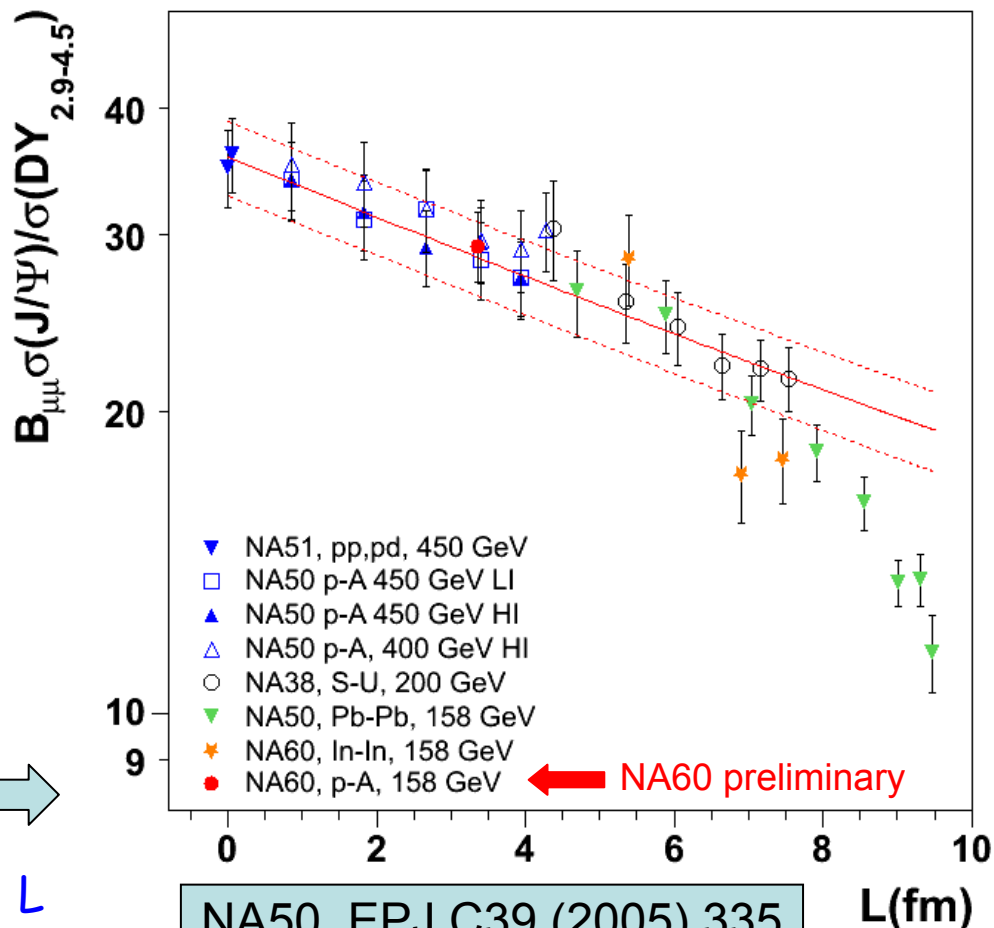
- (including preliminary pA @ 158 GeV from NA60)

- $\exp(-\sigma_{abs} \rho_0 L)$

- (or in Glauber model)
- $\sigma_{abs} = 4,18 \pm 0,35 \text{ mb}$
- L nuclear thickness:



J/ψ / DY rescaled to 158 GeV

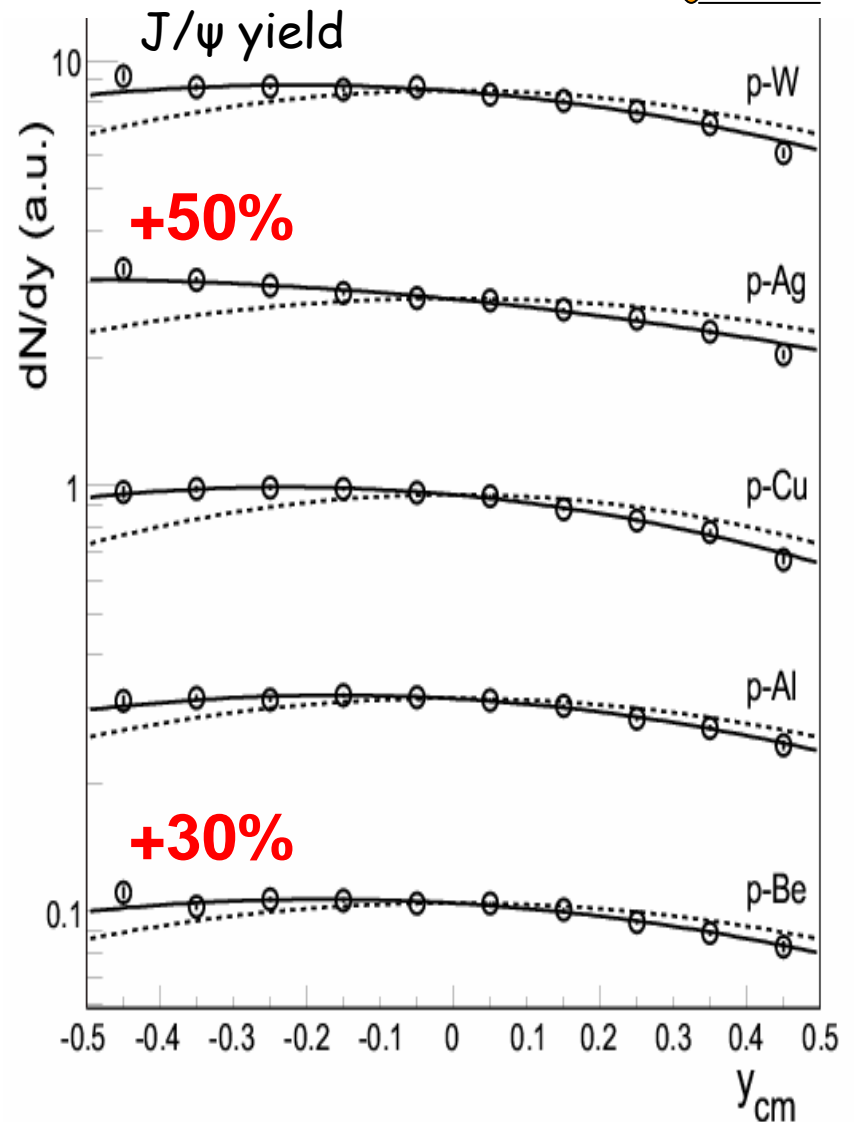


NA50, EPJ C39 (2005) 335
& E. Scomarini's talk

(Anti) shadowing @ SPS ?

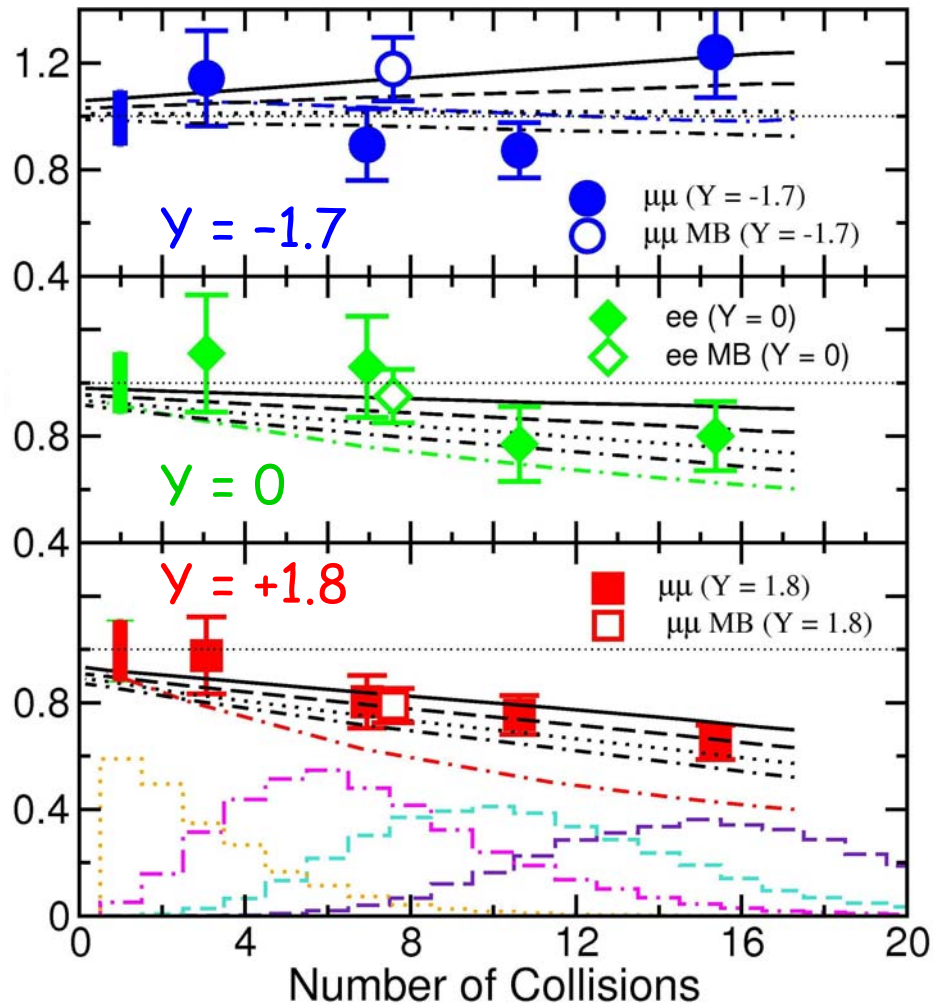
- Do we fully understand CNM @ SPS ?
- Not these surprising rapidity distribution asymmetries →
 - Variation of ~30 to ~50% in one unit of rapidity !
 - Is it (anti)shadowing ?
 - Not taken into account in CNM extrapolation...

NA50, CERN-PH-EP/2006-018,
to appear in Eur. Phys. J. C.



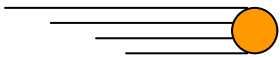
Cold nuclear matter @ RHIC

R_{dAu}

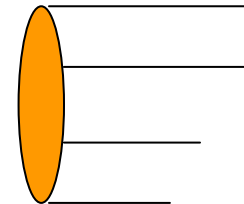


- First centrality dependence in dA (or pA) of J/ψ production!
- Reproduced by Ramona Vogt
 - Black lines: EKS98 shadowing + $\sigma_{abs} = 0$ to 3 mb
 - Colored lines: FGS shadowing + $\sigma_{abs} = 3$ mb
- Favoring moderate shadowing + moderate absorption...

PHENIX, PRL96 (2006) 012304
 Klein, Vogt, PRL91 (2003) 142301



From dA to AA @ RHIC



What is on the market ?

1. Model of nuclear absorption + inhomogeneous (anti)shadowing

(Ramona Vogt, nucl-th/0507027)

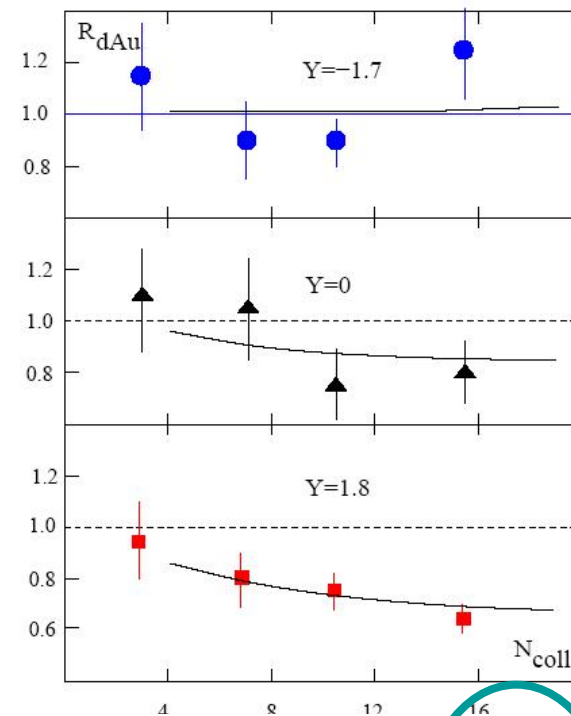
2. $\exp - [(\sigma_{\text{diss}}(y) + \sigma_{\text{diss}}(-y)) \rho_0 L]$

- (Karsch, Kharzeev & Satz
PLB637(2006)75)

- σ_{diss} from fits on dA data \rightarrow

- (unrealistic error bars)

- But shadowing doesn't go like L...



$$\sigma_{\text{diss}}(y = 1.8) = 3.1 \pm 0.2 \text{ mb}$$

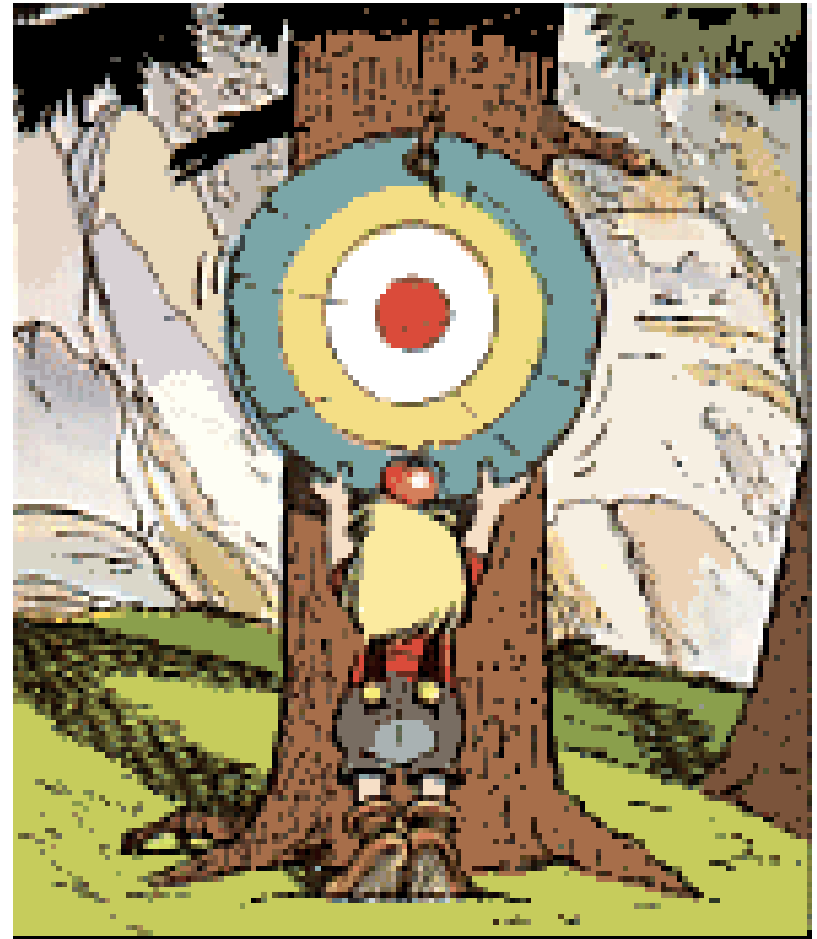
$$\sigma_{\text{diss}}(y = 0) = 1.2 \pm 0.4 \text{ mb}$$

$$\sigma_{\text{diss}}(y = -1.7) = -0.1 \pm 0.2 \text{ mb}$$

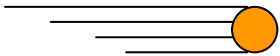
KKS, PLB637(2006)75

3. Another approach...

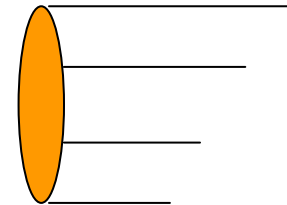
- Goal
 - Predict R_{AA} from R_{dA}
- Concerns
 - Stay as much as possible data-driven
 - Take full advantage of dAu centrality-dependence...



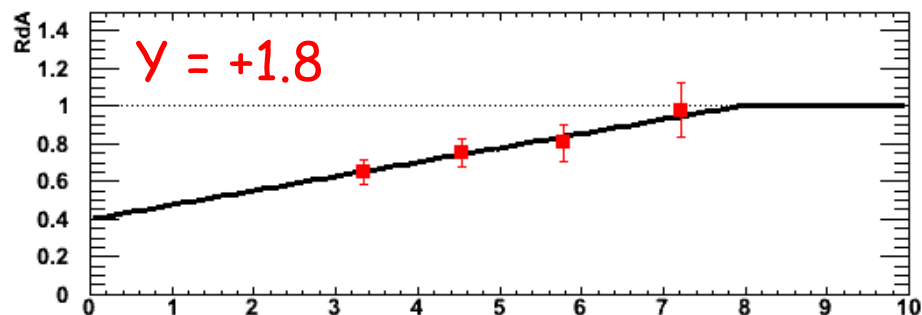
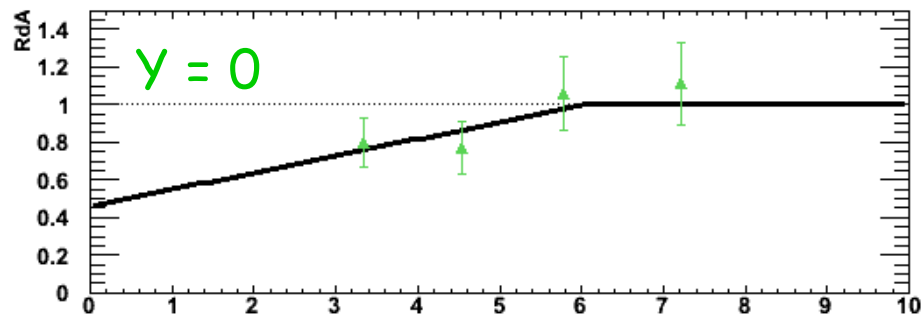
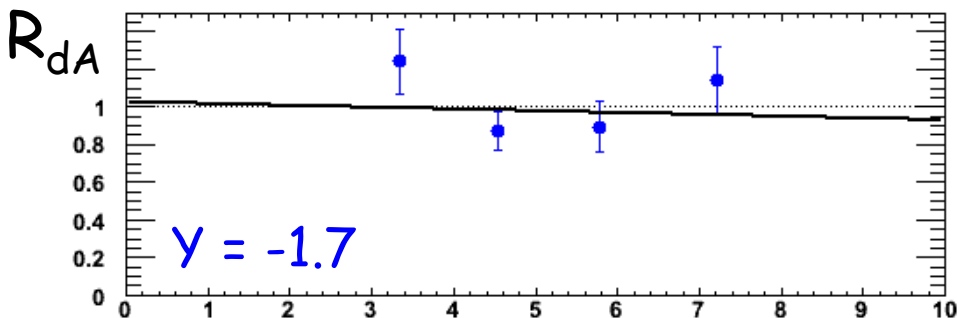
© Uderzo & Goscinny,
Asterix chez les helvètes



R_{dA} vs impact parameter b

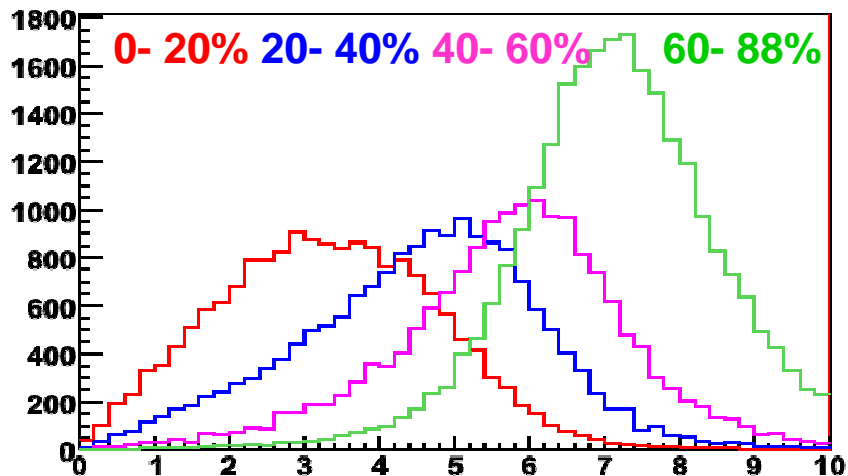


- Re-plot PHENIX R_{dA} vs impact parameter b from Glauber model
- Phenomenological fit to $R_{dA}(b) \rightarrow$
- Cut off $R_{dA}=1$ at high b
 - Physically expected
 - OK for an upper bound of CNM

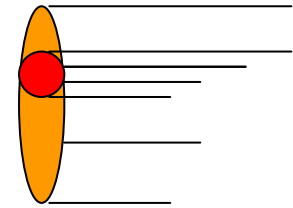
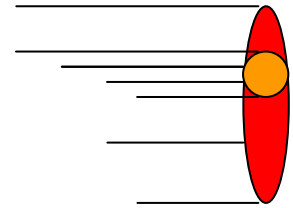


$b(\text{fm})$

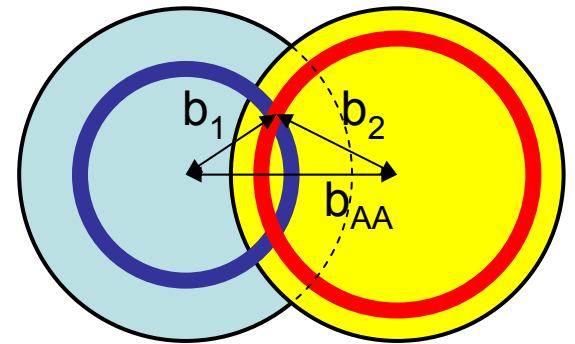
$b(\text{fm})$



Plugged in Glauber model



- Glauber provides, for a given A+A collision at b_{AA} , a set of N+N collisions occurring at $b_1^{i_1}$ and $b_2^{i_2}$.

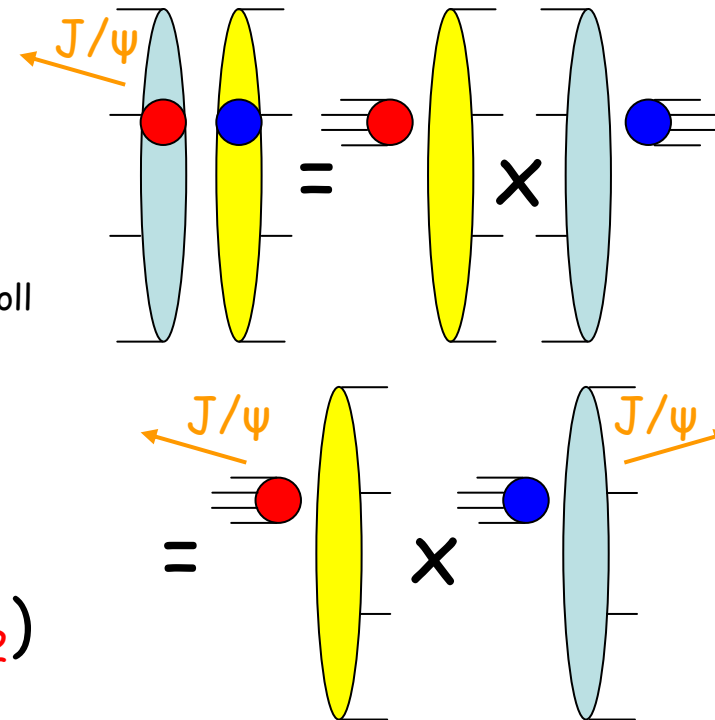


- One minimal assumption is rapidity factorization: $R_{AA}(|y|, b_{AA}) =$

$$\Sigma_{\text{collisions}} [R_{dA}(-y, b_1^{i_1}) \times R_{dA}(+y, b_2^{i_2})] / N_{\text{coll}}$$

- Works (at least) for absorption & shadowing since production

$$\sim \text{pdf1} \times \text{pdf2} \times \exp -\rho\sigma(L_1+L_2)$$

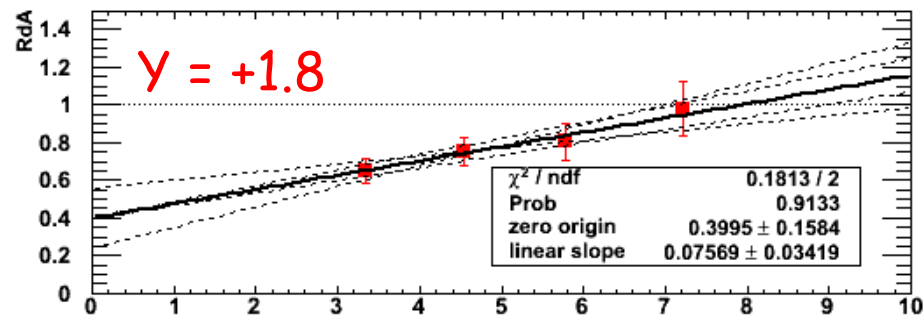
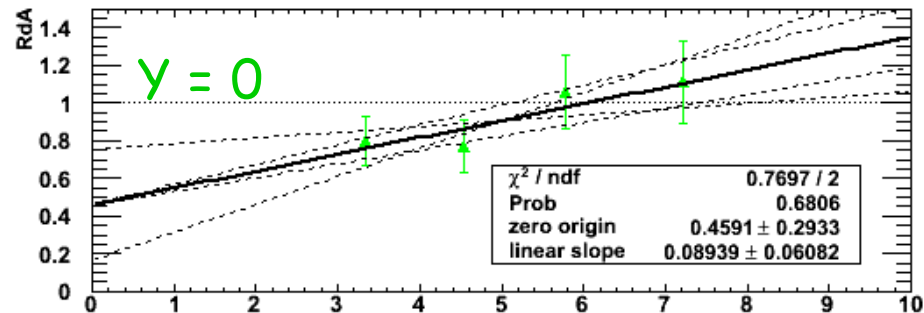
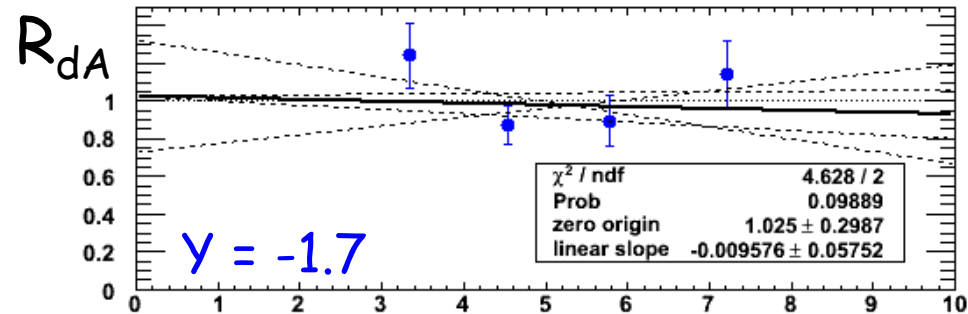


Propagating dA error

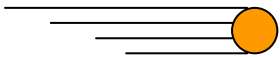
1. Varying the fit parameters

- Uncorrelated
- Within $\pm 1\sigma$

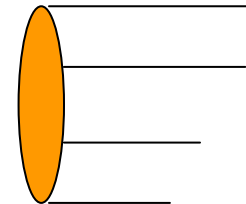
to propagate the statistic + systematic dAu uncertainties throughout the Glauber computation



$b(\text{fm})$



Systematic uncertainties

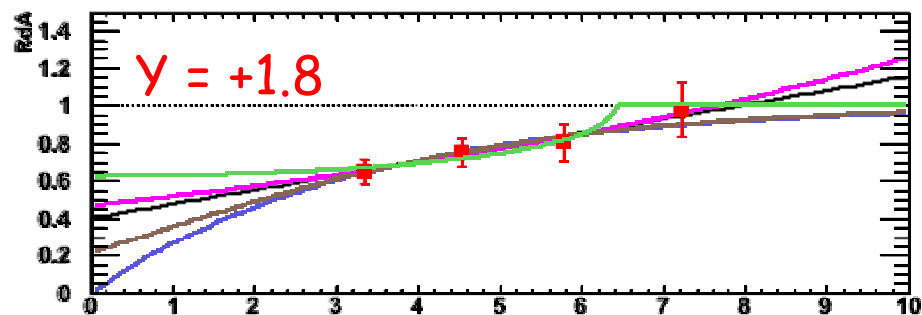
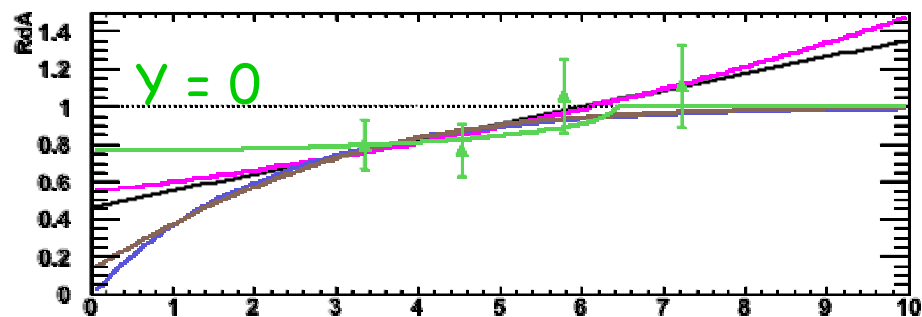
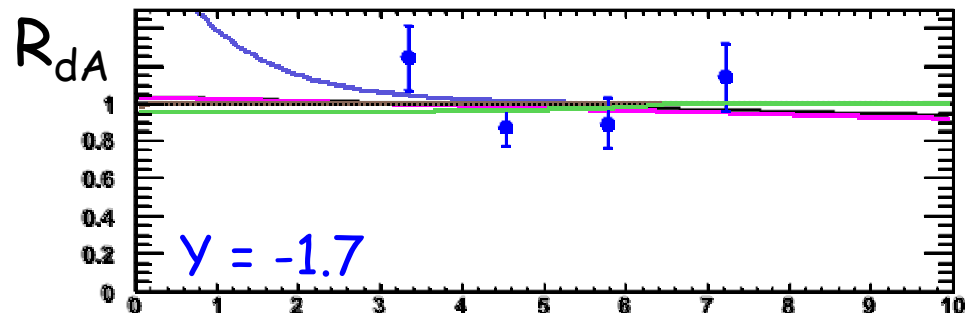


2. Varying line shapes \rightarrow

- 3 to 5% @ $y \sim 1.7$
 - 3 to 12% @ $y \sim 0$
- (asymmetric and depending on centrality)

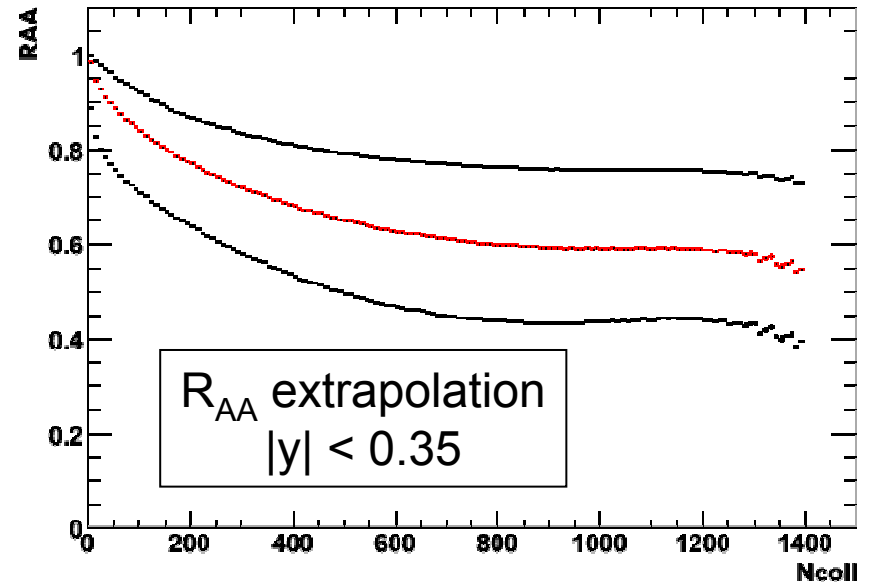
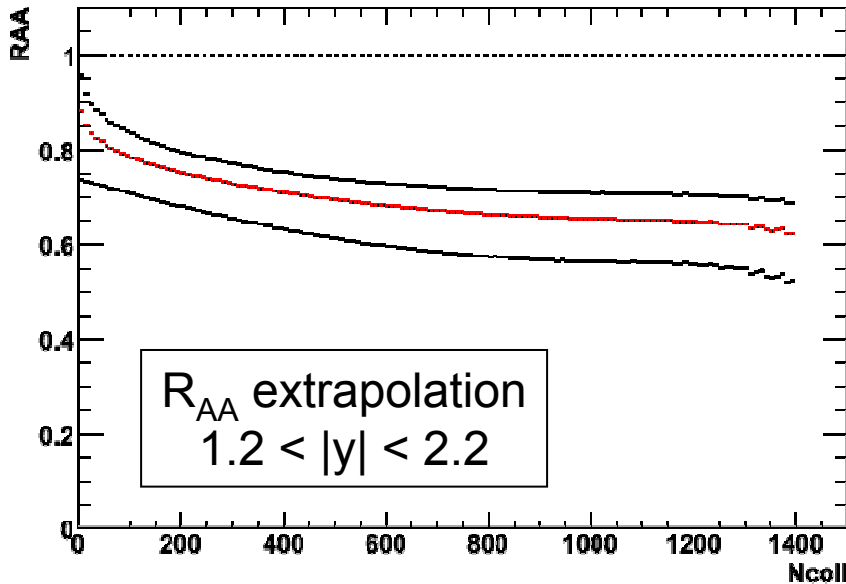
3. Varying Glauber parameters (pp total cross section, Woods Saxon parameters,...)

- 2% @ $y \sim 1.7$
- 4% @ $y \sim 0$



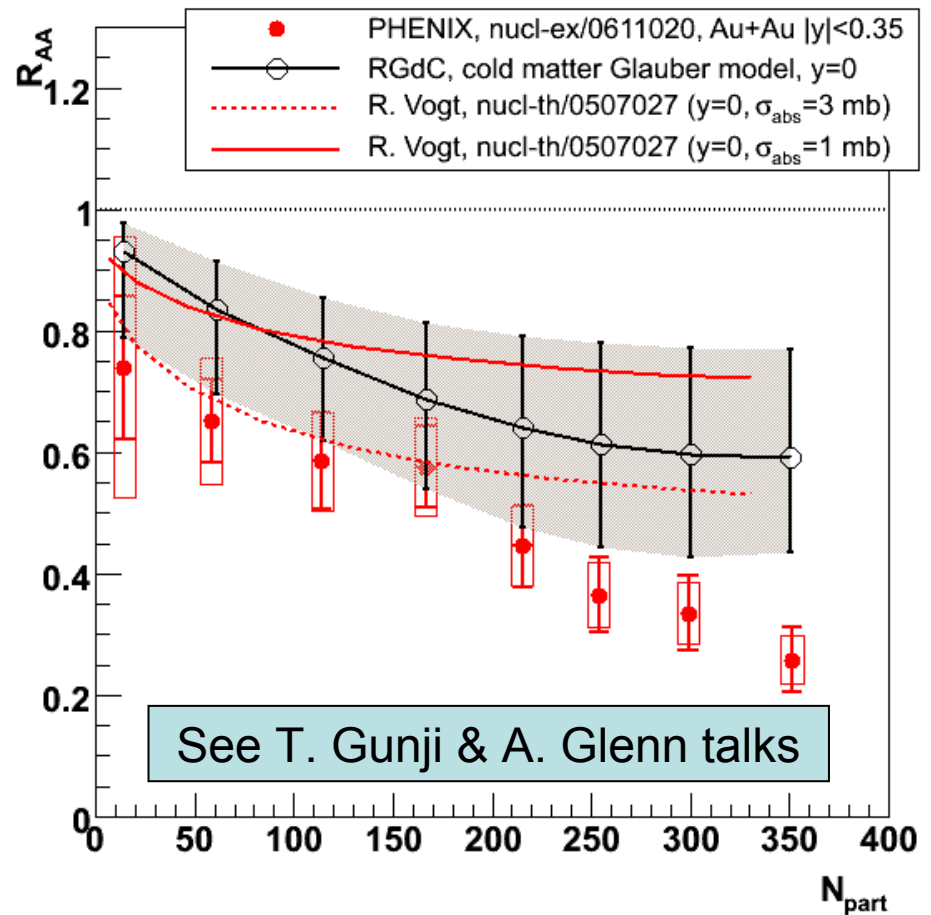
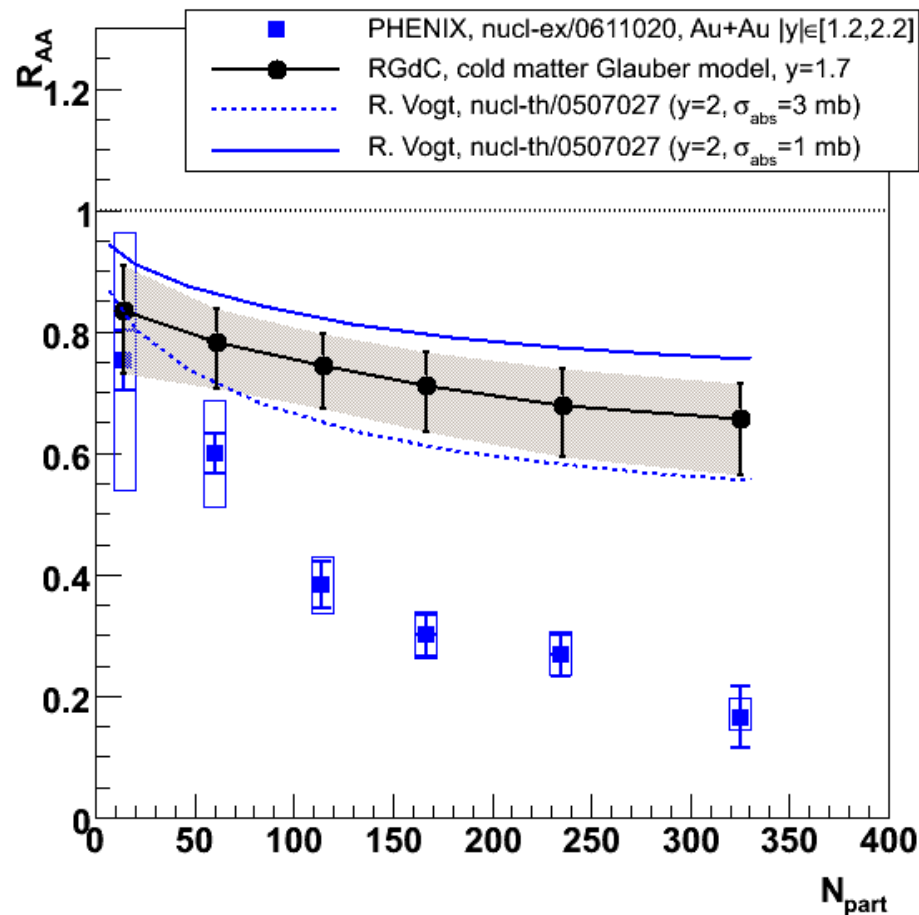
$b(\text{fm})$

The output: $R_{AA}(N_{coll})$



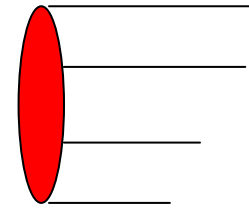
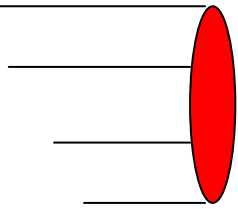
- Black curves reflect stat. and syst. errors from dAu
- Much less constrained @ $y \sim 0$ because:
 - $R_{dA}(0)$ measurements are less precise than $R_{dA}(-1.7)$ and $R_{dA}(+1.7)$
 - and squared while computing $R_{dA}(-y) \times R_{dA}(+y)$
- Then, take the average in experimental centrality classes

Comparison to models & data



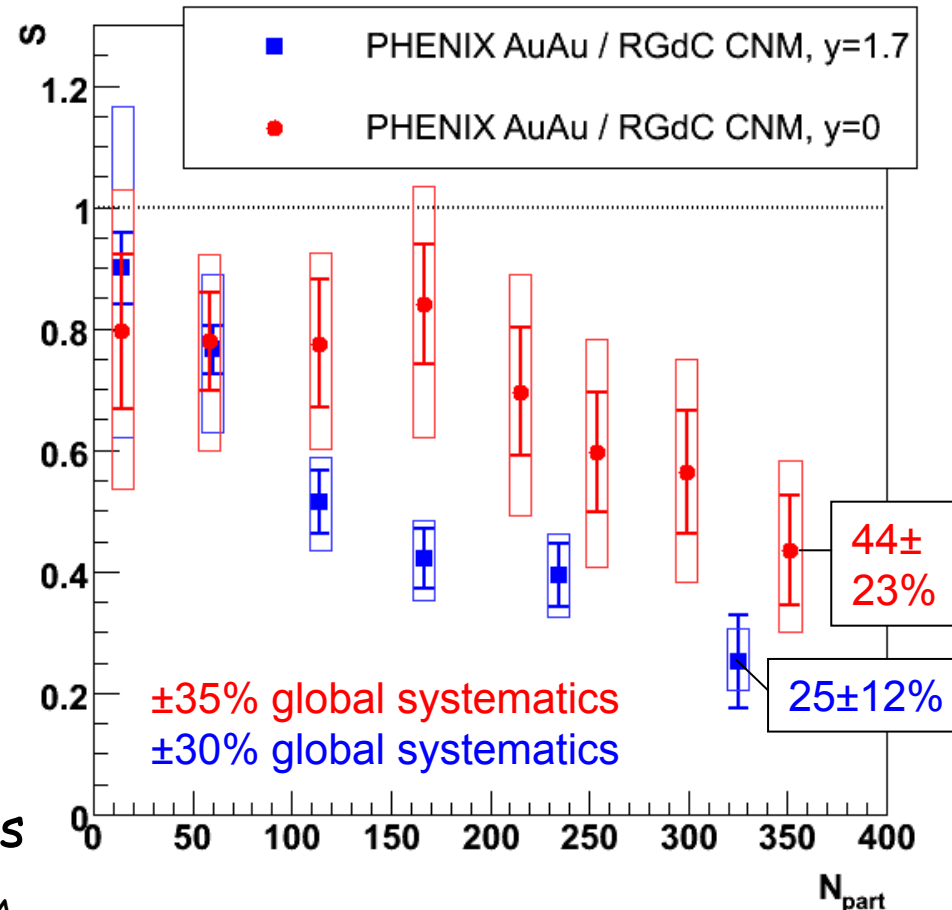
- Consistent with Vogt's prediction (EKS shadowing + 1 or 3 mb σ_{abs})
- Prediction @ $y \sim 1.7$ is much more powerful than @ $y \sim 0$

$R_{AA} / \text{CNM} @ \text{RHIC}$



- First R_{AA}/CNM extraction including (proper) error propagation
- Boxes are correlated errors from AuAu & dominant CNM
- Important: missing overall global relative uncertainty
 - 30% @ $y \sim 1.7$ / 35% @ $y \sim 0$
 - Due to different pp references that don't cancel in R_{dA} and R_{AA}
$$R_{AA}(|y|) / R_{dA}(-y) \times R_{dA}(+y)$$

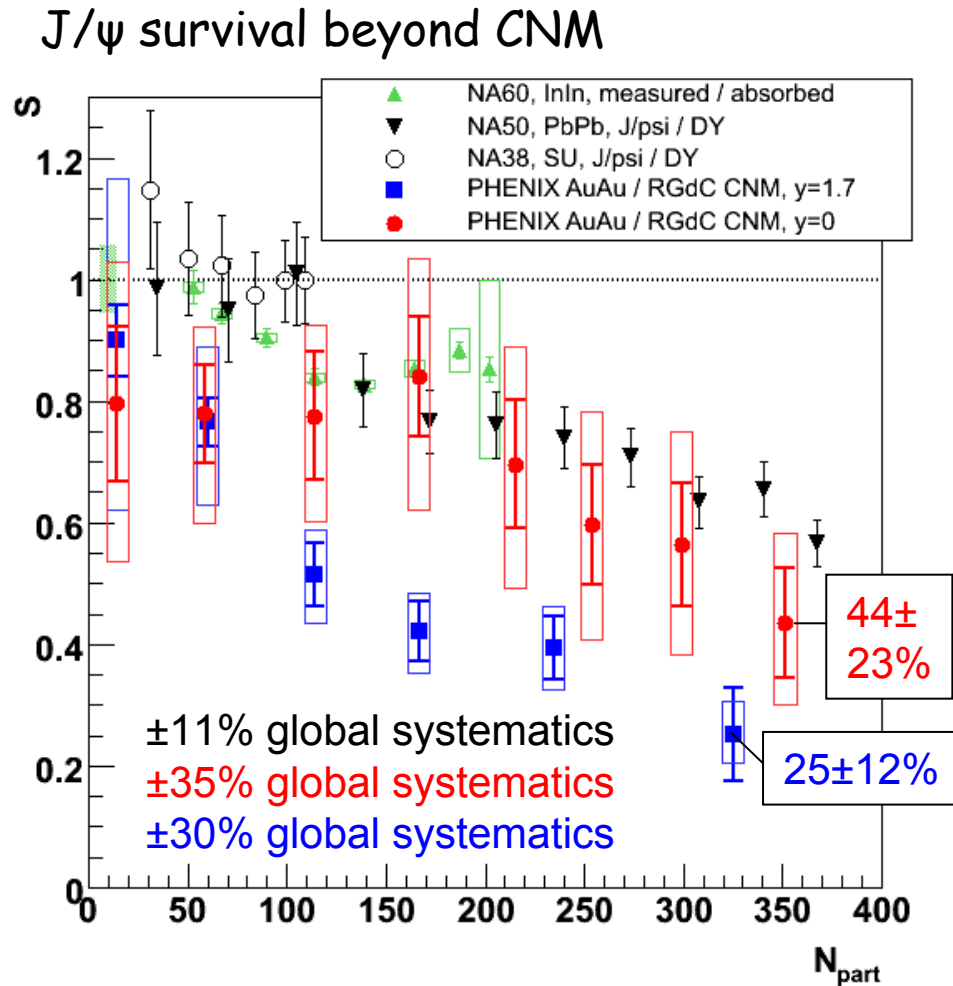
J/ ψ survival beyond CNM



← This will decrease by recomputing R_{dA} with new pp data (A.Bickley's talk)

Quick comparison to SPS

- **At mid-rapidity**, the amount of surviving J/ψ @ RHIC is compatible with SPS (~60%) but depends a lot on CNM (and pp references)...
- **At forward rapidity**, RHIC anomalous suppression is much stronger!





Conclusions



• Pro's

- Very little model dependence (apart from Glauber)
 - (no σ_{abs} , no shadowing scheme, $y=0$ & 1.7 independence,...)
- Proper error propagation from dAu (and pp)
- Proper centrality selection (experimental classes)
→ J/ ψ survival of $25 \pm 12\%$ @ $y=1.7$ & $44 \pm 23\%$ @ $y=0$

• Con's

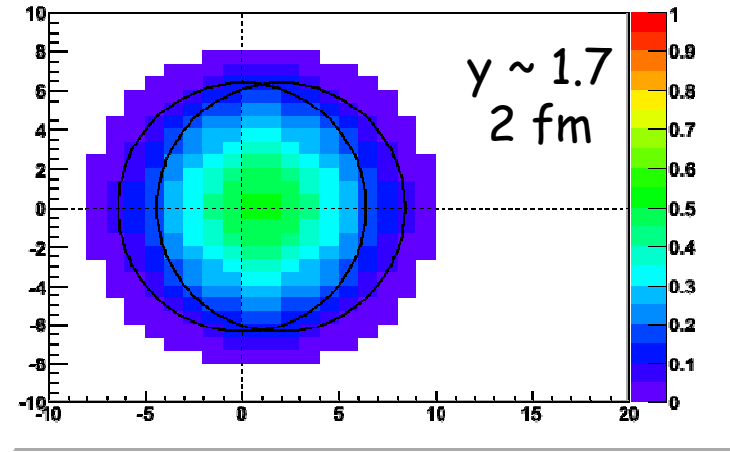
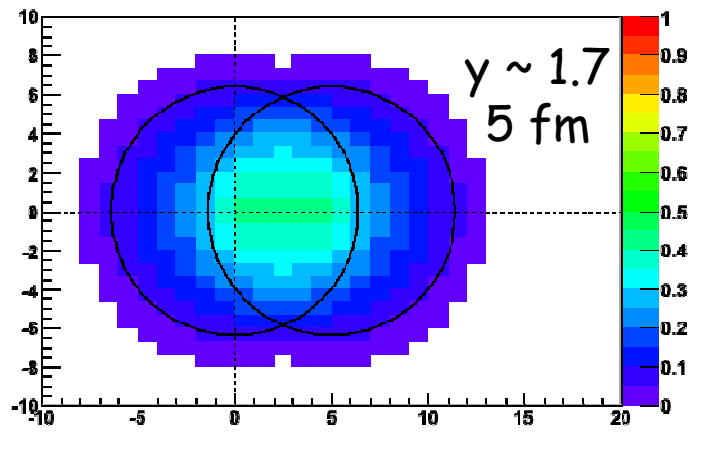
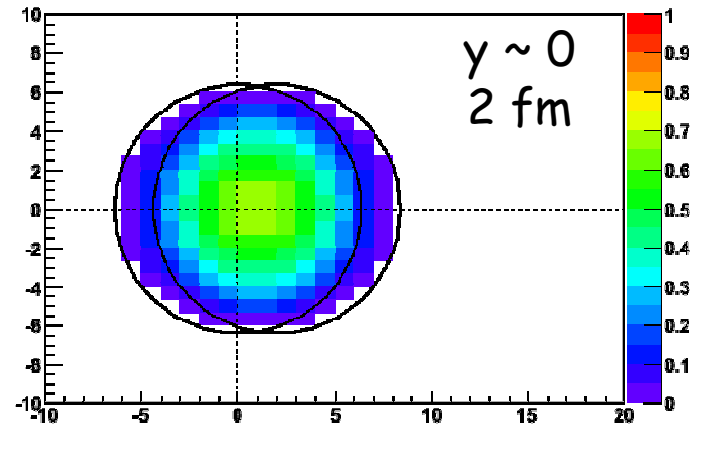
- Not applicable without p+A (or d+A) centrality dependence at same energy and at both +y and -y wrt A+A collisions (thus not at SPS or LHC)
- Limited by dAu statistic ! We need more !
 - Especially @ $y \sim 0$ (and dCu to apply this to CuCu)

Back-up slides

Collision display

- Disappearance probability
 - No assumption on production point
 - Weighted by Woods Saxon

Disappearance probability



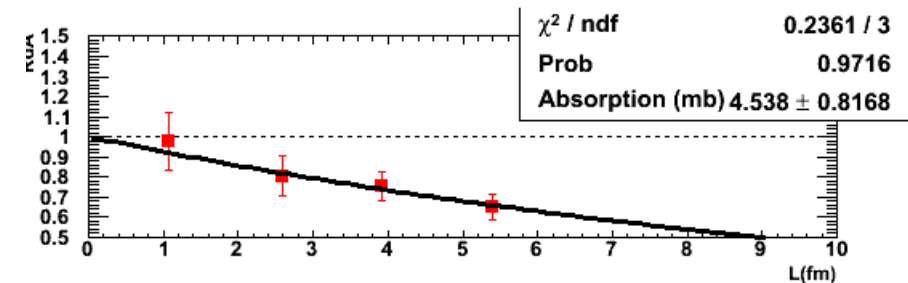
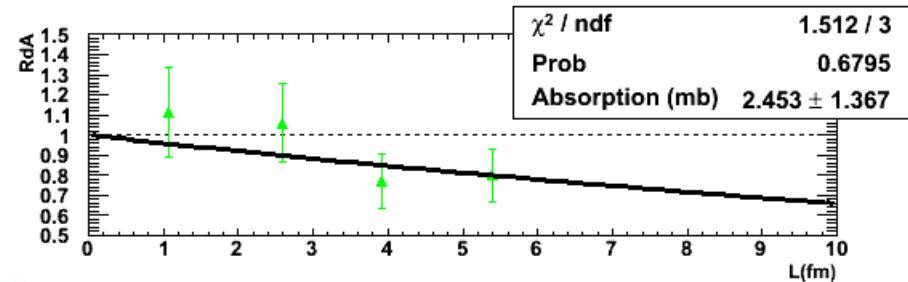
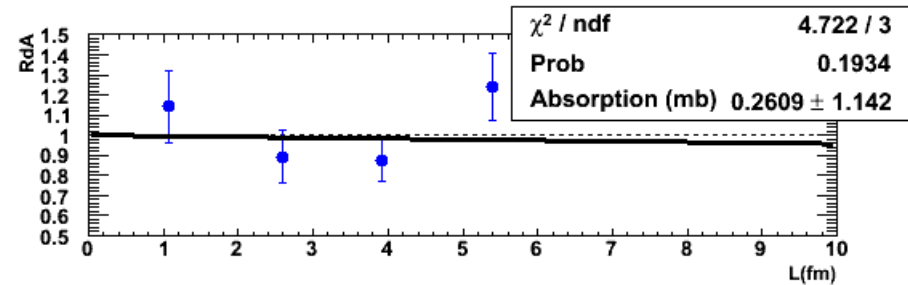
Nuclear absorption only

- Compute L with Glauber model
- Fit $\exp(-\sigma_{\text{abs}} \rho_0 L)$
- Results are different wrt KKS numbers

Rapidity	KKS fit [4]	My fit
$y = -1.7$	-0.1 ± 0.2 mb	0.3 ± 1.1 mb
$y = 0$	1.2 ± 0.4 mb	2.4 ± 1.4 mb
$y = 1.8$	3.1 ± 0.2 mb	4.5 ± 0.8 mb

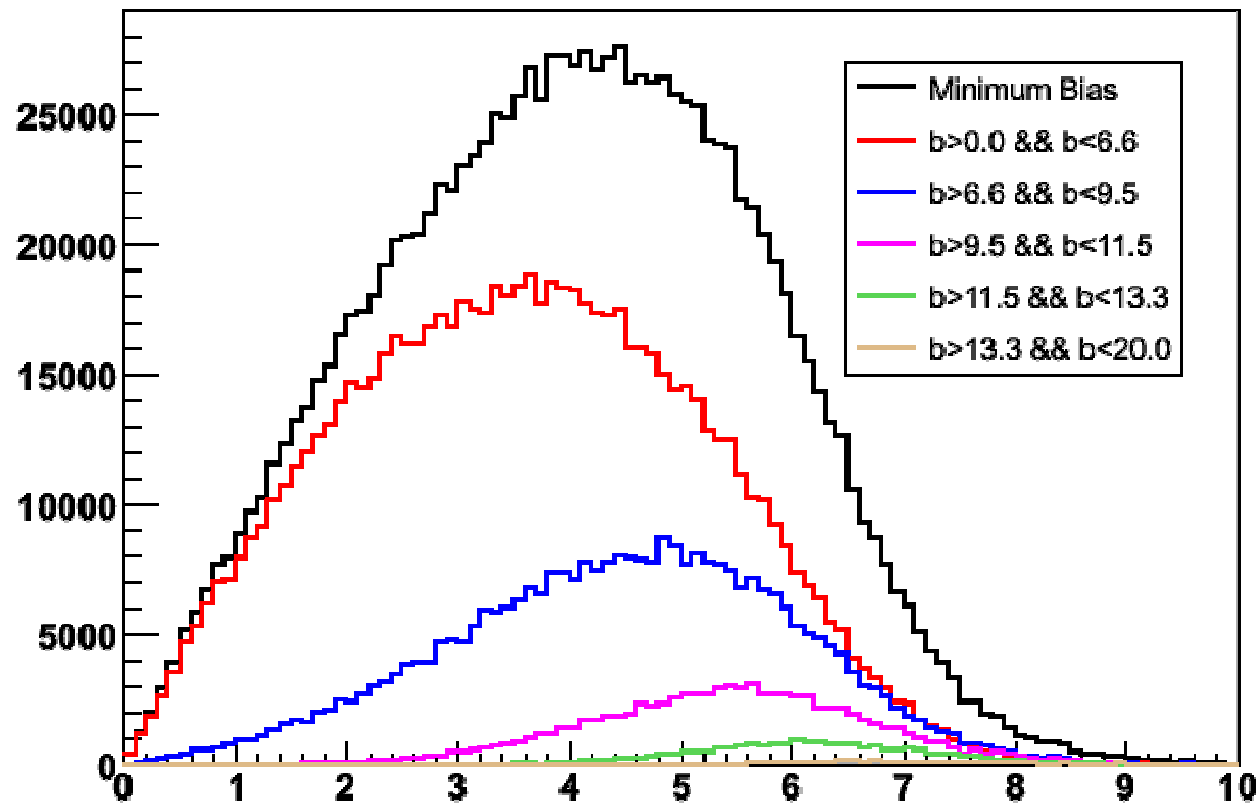
TAB. 1 – σ_{diss} values from KKS and my analysis.

KKS, PLB637(2006)75

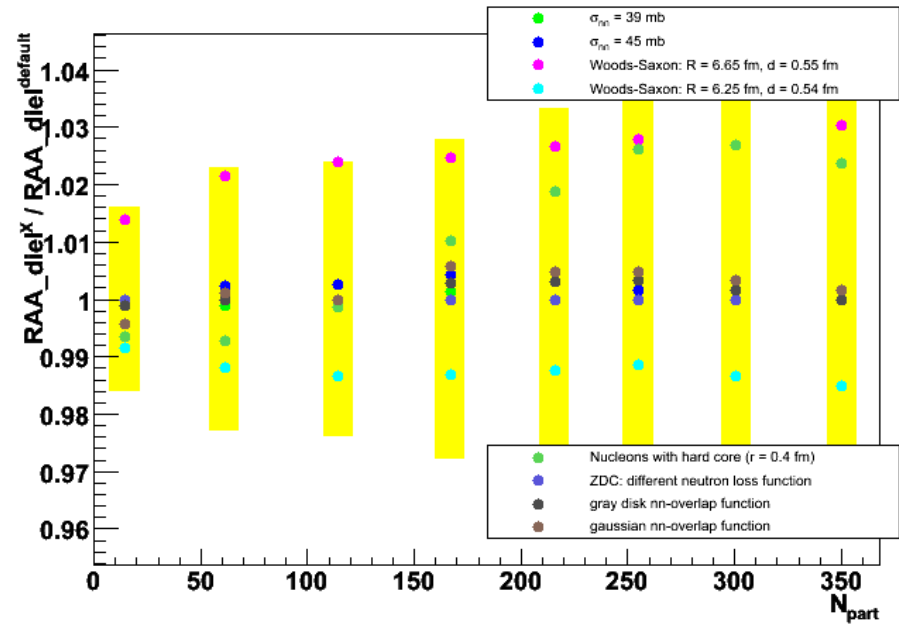
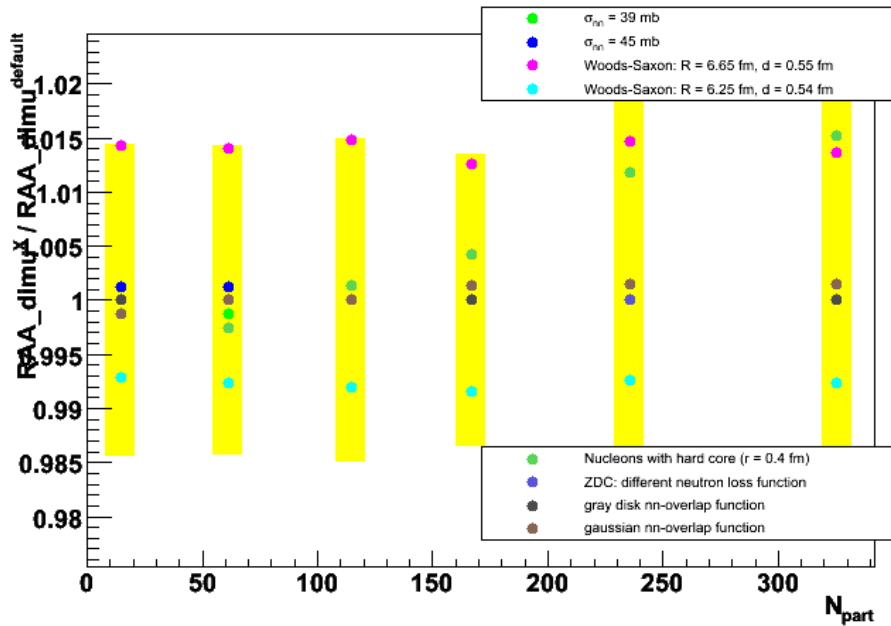


Sampling local impact parameter in AuAu

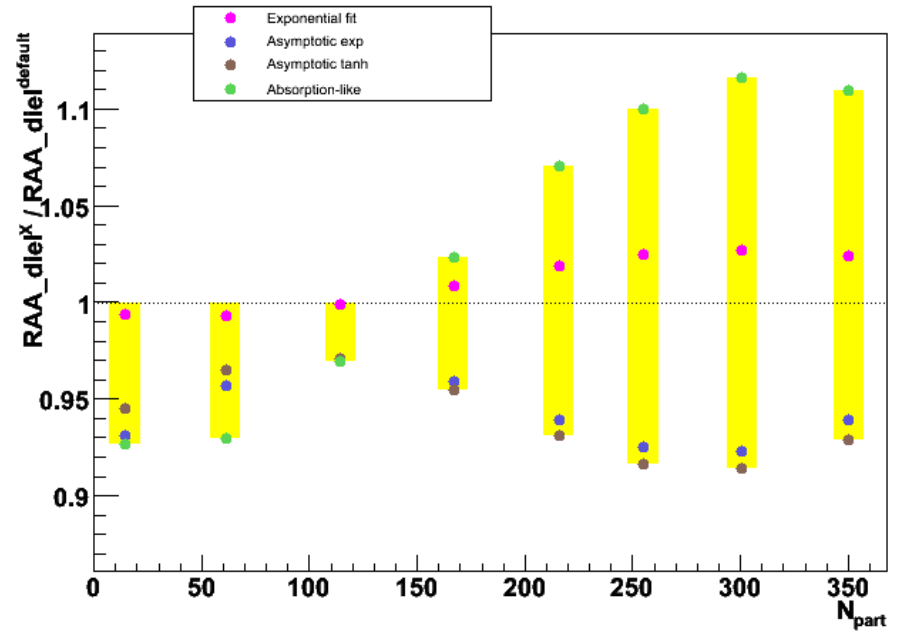
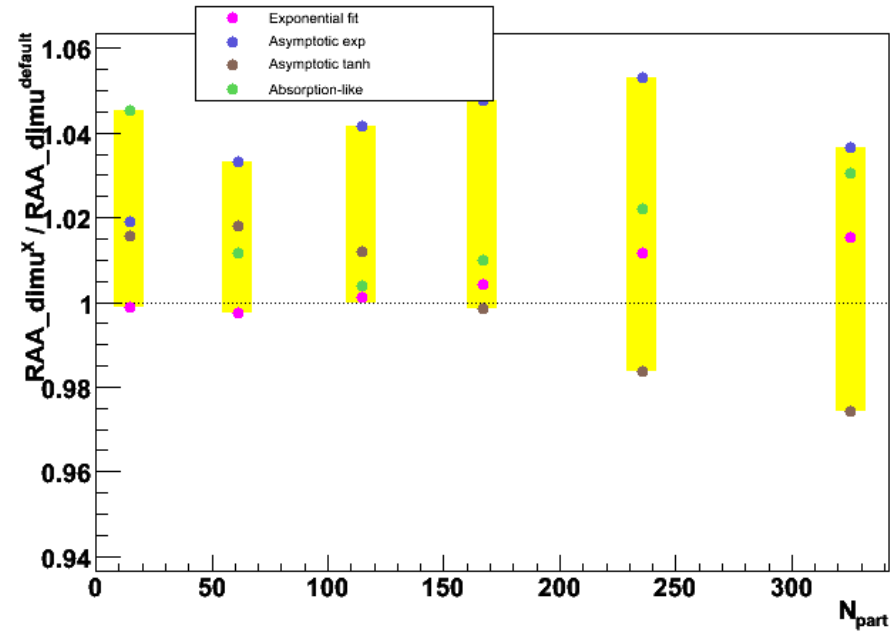
Local impact parameters (fm)

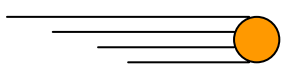


Varying Glauber parameters



Varying the line shape

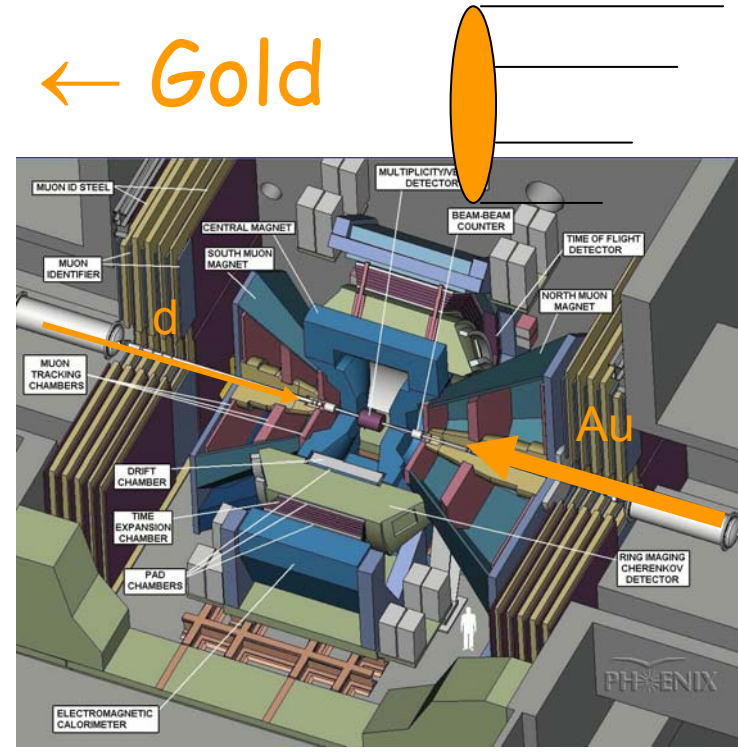




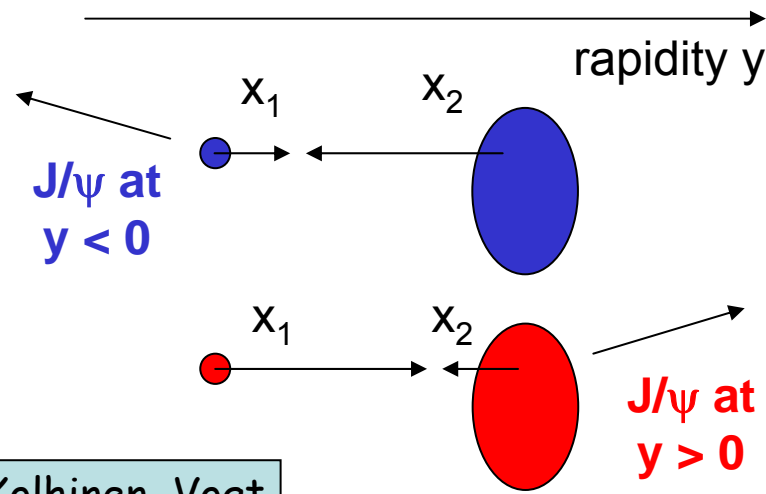
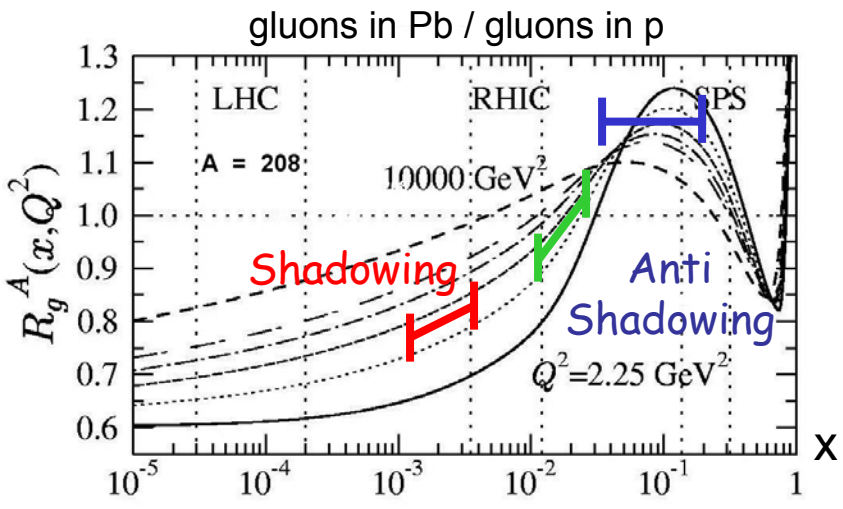
Deuteron →

← Gold

- In PHENIX, J/ψ mostly produced by gluon fusion, and thus sensitive to gluon pdf
- Three rapidity ranges probe different momentum fraction of Au partons
 - South ($y < -1.2$) : large x_2 (in gold) ~ 0.090
 - Central ($y \sim 0$) : intermediate x_2 ~ 0.020
 - North ($y > 1.2$) : small x_2 (in gold) ~ 0.003



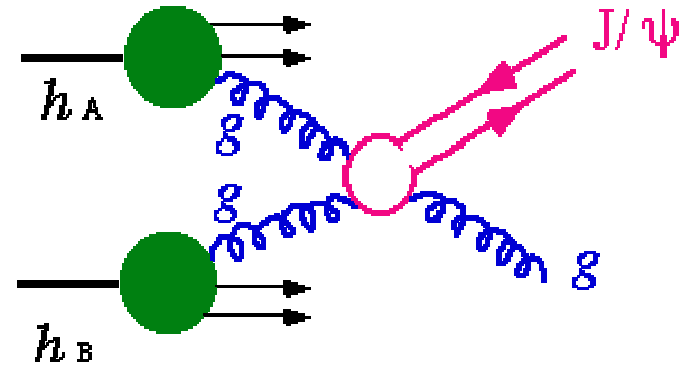
An example of gluon shadowing prediction



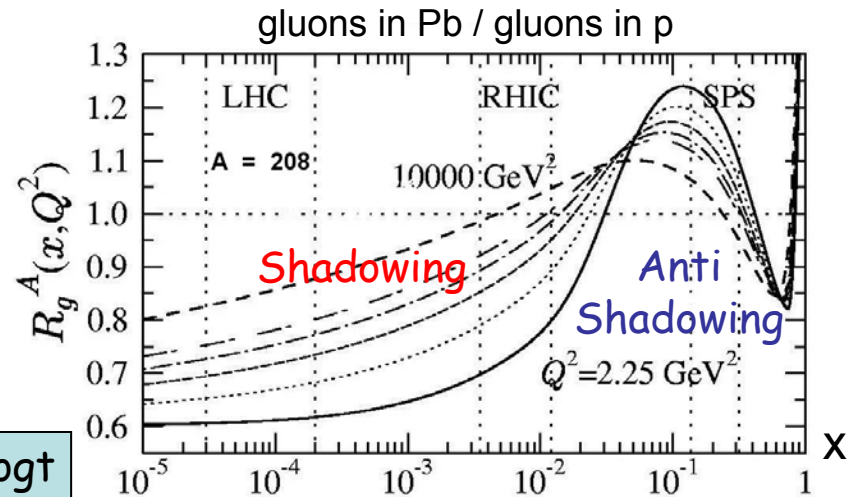
Eskola, Kolhinen, Vogt
NPA696 (2001) 729

Cold nuclear matter effects ?

- J/ψ (or $c\bar{c}$) absorption
- (Anti) shadowing
(gluon saturation, CGC...)
- Energy loss of initial parton
- p_T broadening (Cronin effect)
- Complications from feeddown ψ' & χ_c ?
- Something else ?



An example of gluon shadowing prediction



Eskola, Kolhinen, Vogt
NPA696 (2001) 729

Cold nuclear matter effects ?

A real puzzle ! Especially when one goes to low x_2 , high x_F ...

$$\sigma_\psi(pA) = \sigma_\psi(pp) \times A^\alpha$$

